



# Indications of Neutrino Oscillation in the K2K Neutrino Oscillation Experiment



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Les Rencontre de Physique  
de la Vallee d' Aoste,

Taku I SHI DA (KEK I PNS)  
for K2K Collaboration



# K2K Collaboration

- **JAPAN:** High Energy Accelerator Research Organization (KEK)  
Institute for Cosmic Ray Research (ICRR), University of Tokyo  
Kobe University / Kyoto University  
Niigata University / Okayama University  
Tokyo University of Science / Tohoku University
- **KOREA:** Chonnam National University  
Dongshin University / Korea University  
Seoul National University
- **U.S.A.:** Boston University / University of California, Irvine  
University of Hawaii, Manoa  
Massachusetts Institute of Technology  
State University of New York at Stony Brook  
University of Washington at Seattle
- **POLAND:** Warsaw University / Solton Institute

Since November 2002

- **JAPAN:** Hiroshima University
- **EUROPE:** Rome / Saclay / Barcelona / Valencia / Geneva
- **RUSSIA:** Dubna



# Introduction

- K2K is the first accelerator-based long-baseline neutrino oscillation experiment to investigate the neutrino oscillation observed in atmospheric neutrinos.

$$P(\nu_\mu \rightarrow \nu_x) = \sin^2 2\theta \cdot \sin^2(1.27 \Delta m^2 L / E\nu)$$

	Atm.- $\nu$	K2K
<b>L</b>	10~10 <sup>4</sup> km	250 km (fix.)
<b>E<math>\nu</math></b>	0.1~100 GeV	~ 1.3 GeV
<b><math>\Delta m^2</math></b>	10 <sup>-1</sup> ~ 10 <sup>-4</sup> eV <sup>2</sup>	> 2 · 10 <sup>-3</sup> eV <sup>2</sup>
<b><math>\nu_e / \nu_\mu</math></b>	50 %	~1%

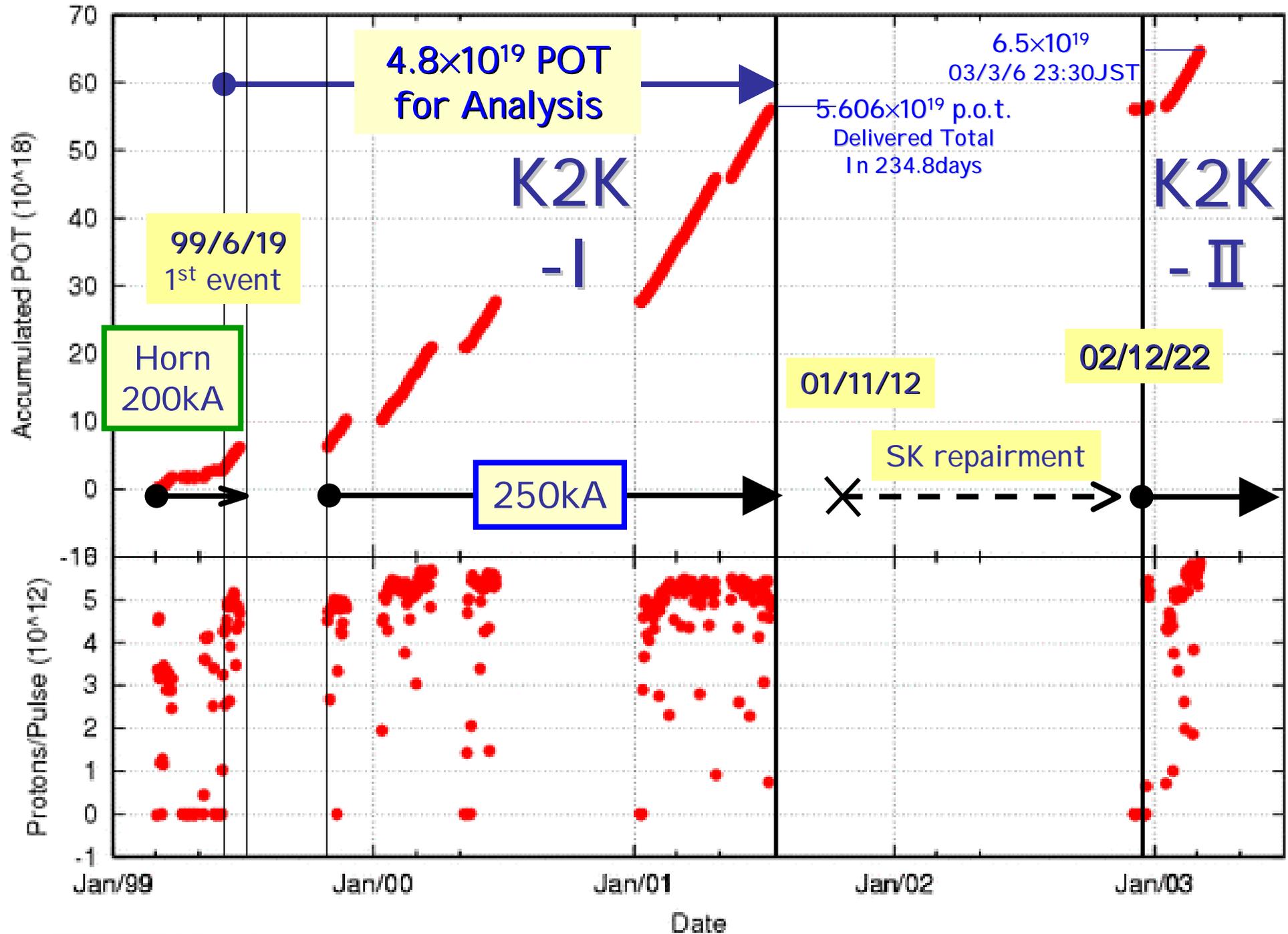
Super-Kamiokande



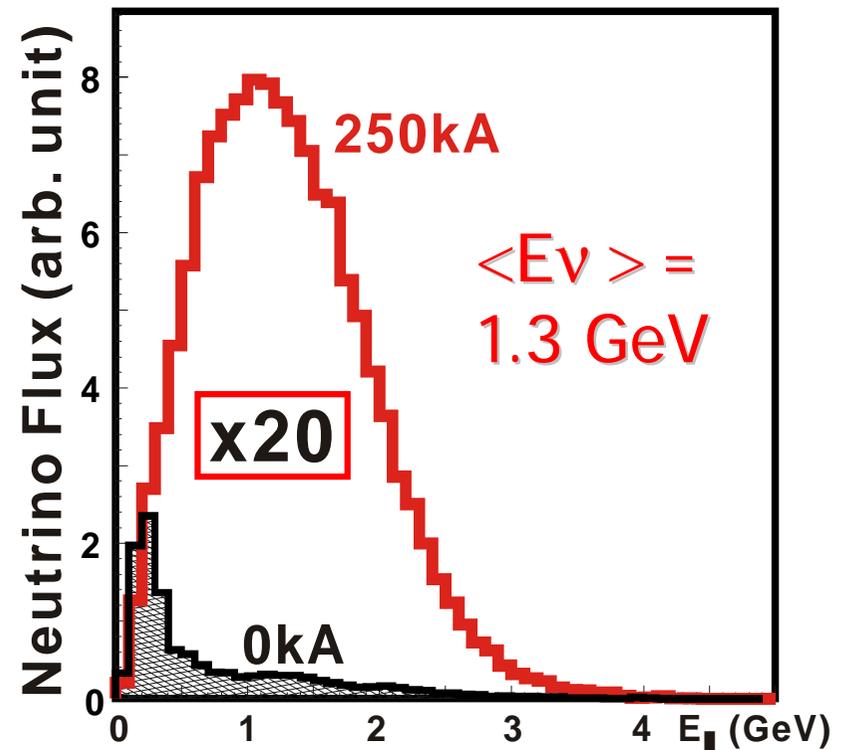
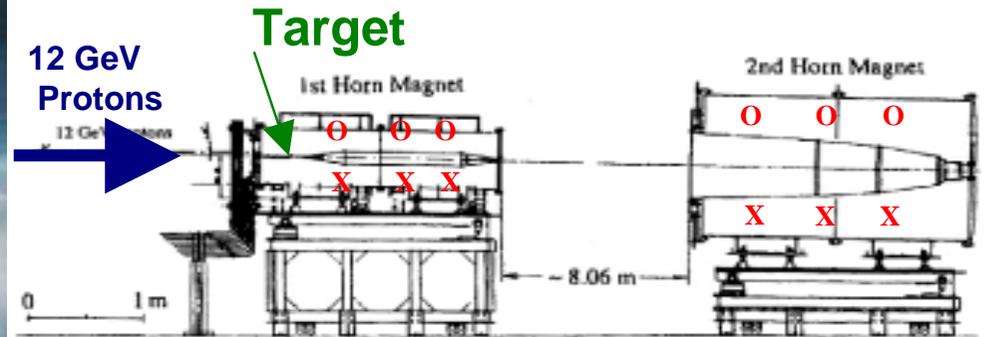
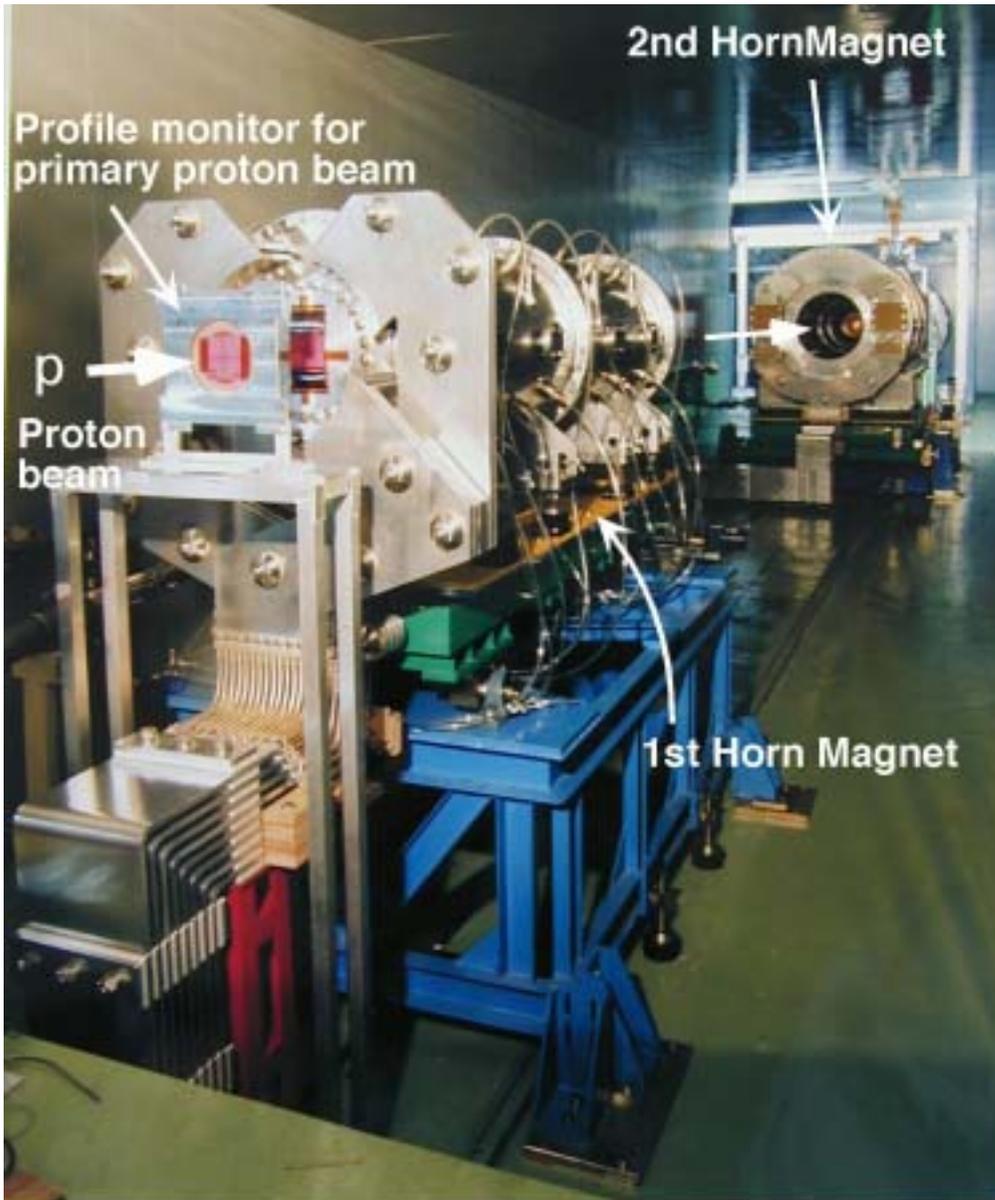
KEK-12GeV PS

- Do  $\nu_\mu$  events decrease ?
- Is  $E\nu$  spectrum distorted ?

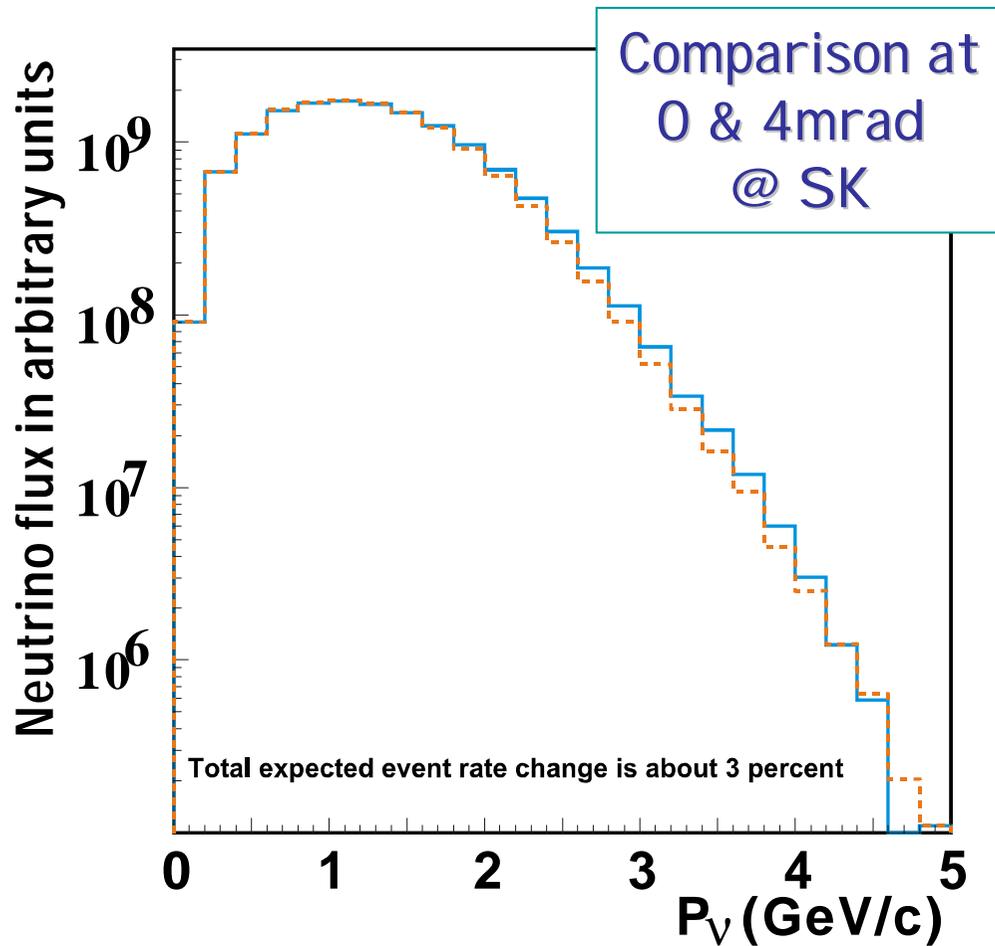




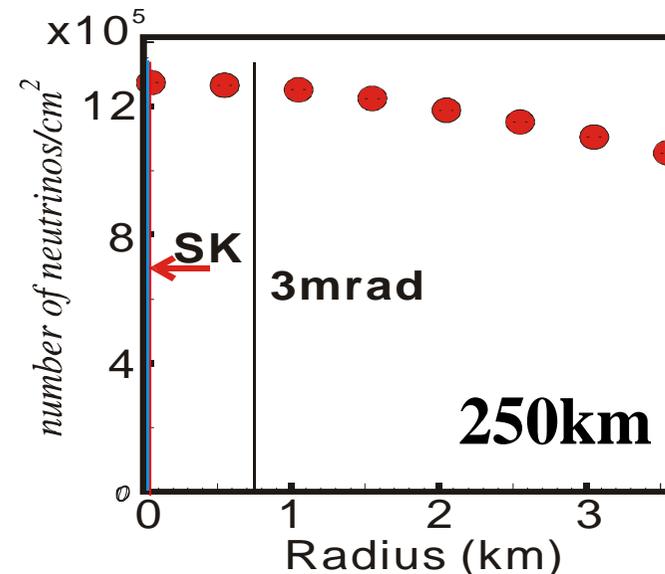
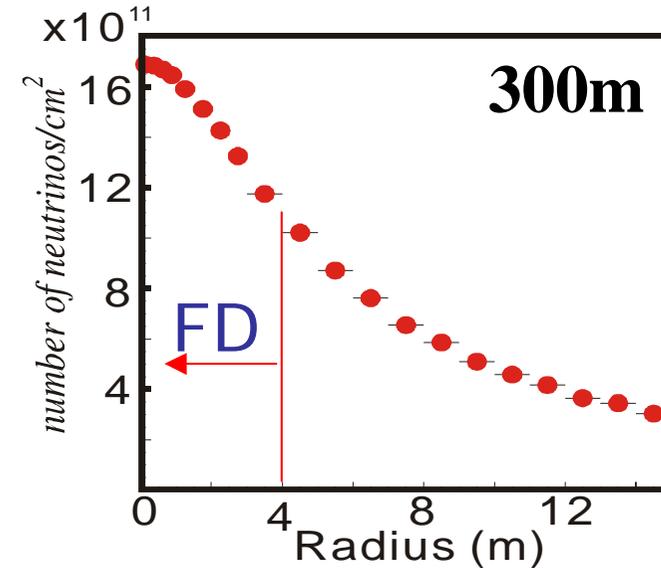
# Target (Al) and Two HORNS



# Expected (MC) Neutrino Spectra and Radial Distributions at 300m/250km

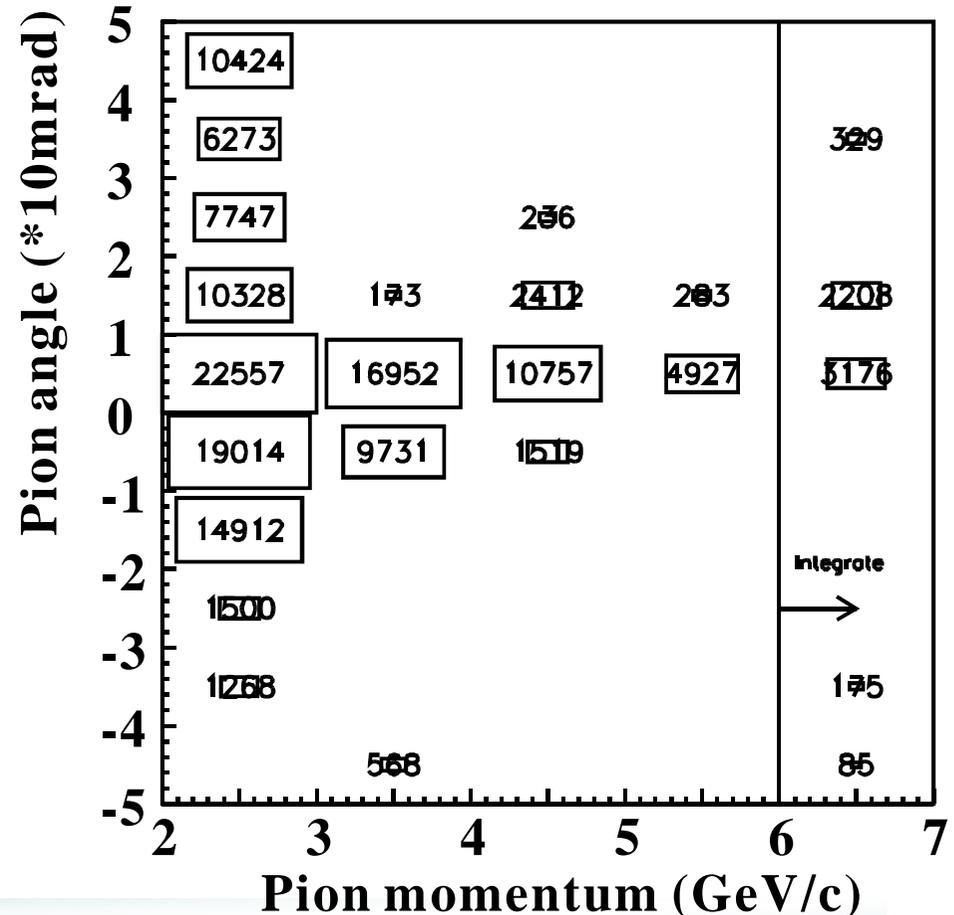
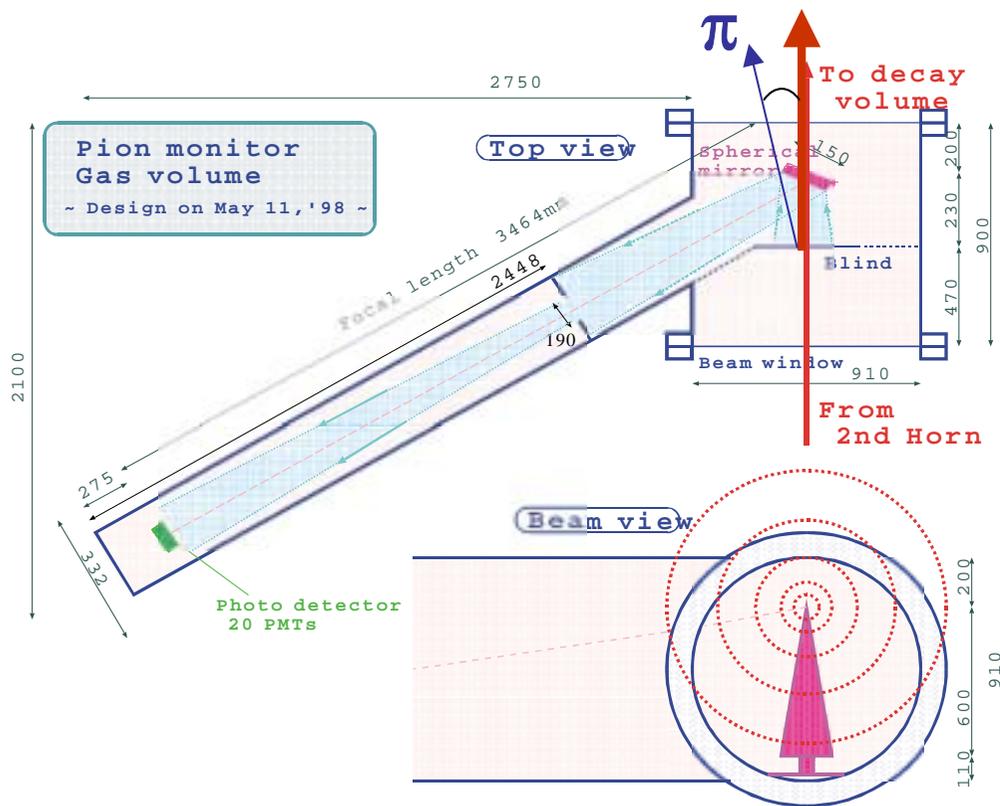


No change in rate and spectrum

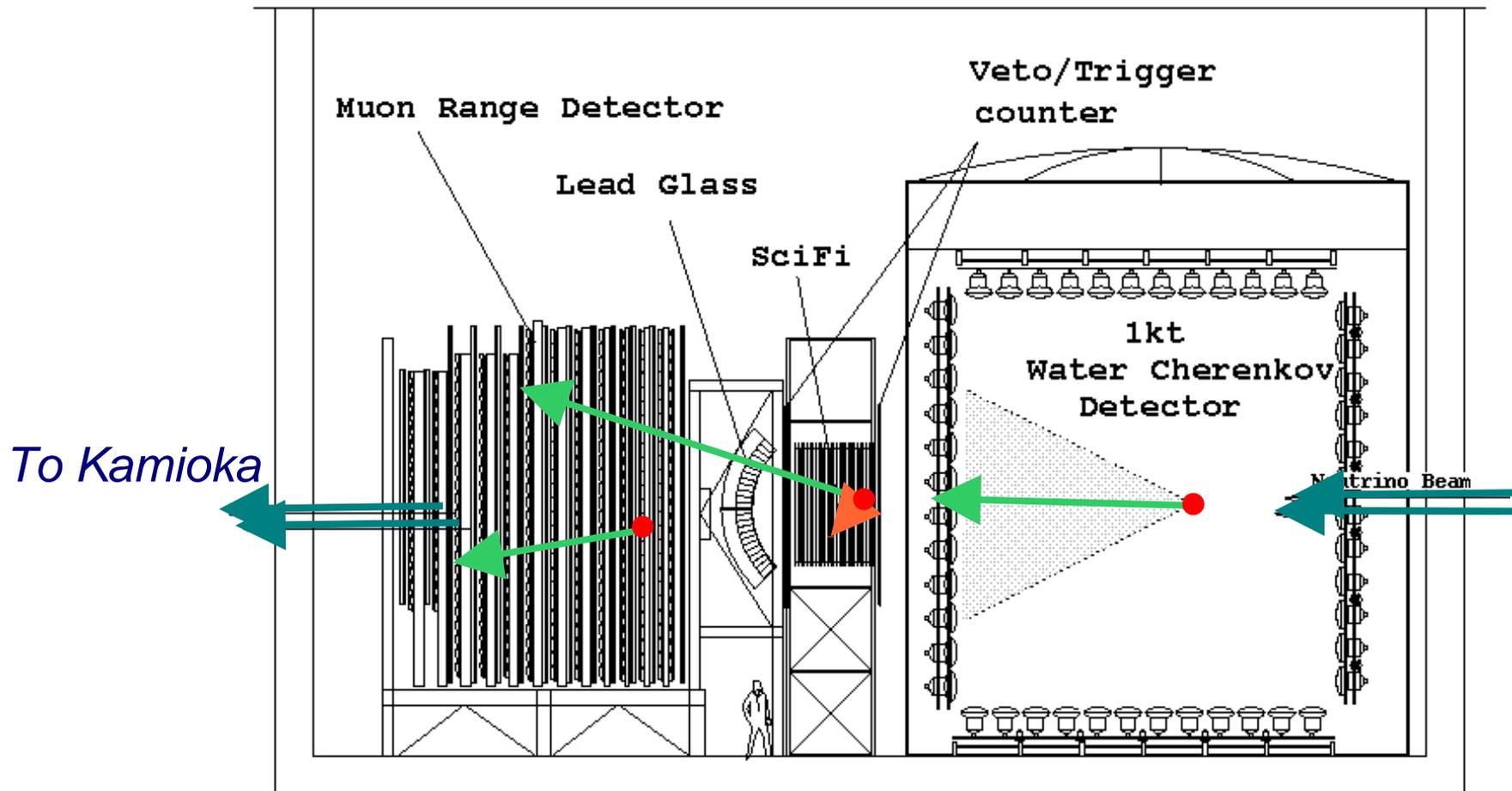


# Pion Monitor

- Gas Cherenkov detector: Insensitive to primary protons.
- Measure momentum and angular distribution of pions  
 $N(p_\pi, \theta_\pi)$  just after the 2<sup>nd</sup> horn ( $p_\pi > 2 \text{ GeV}/c$ ).
- Near to far extrapolation:  $F/N(E_\nu)$

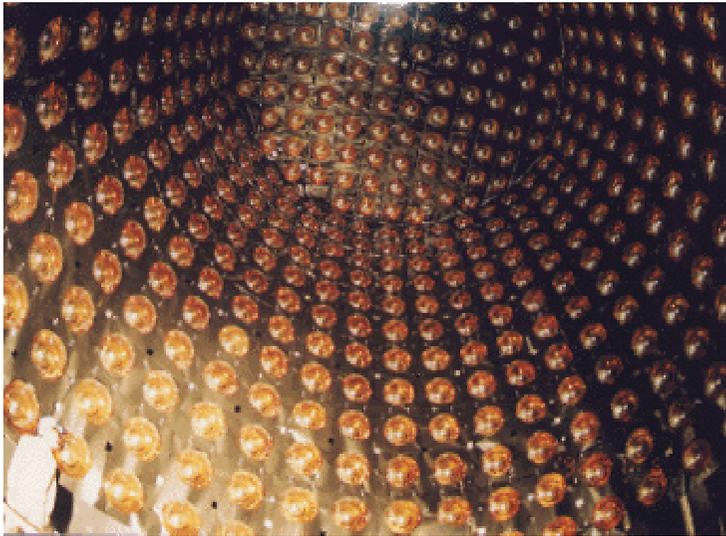


# Front neutrino Detectors at KEK

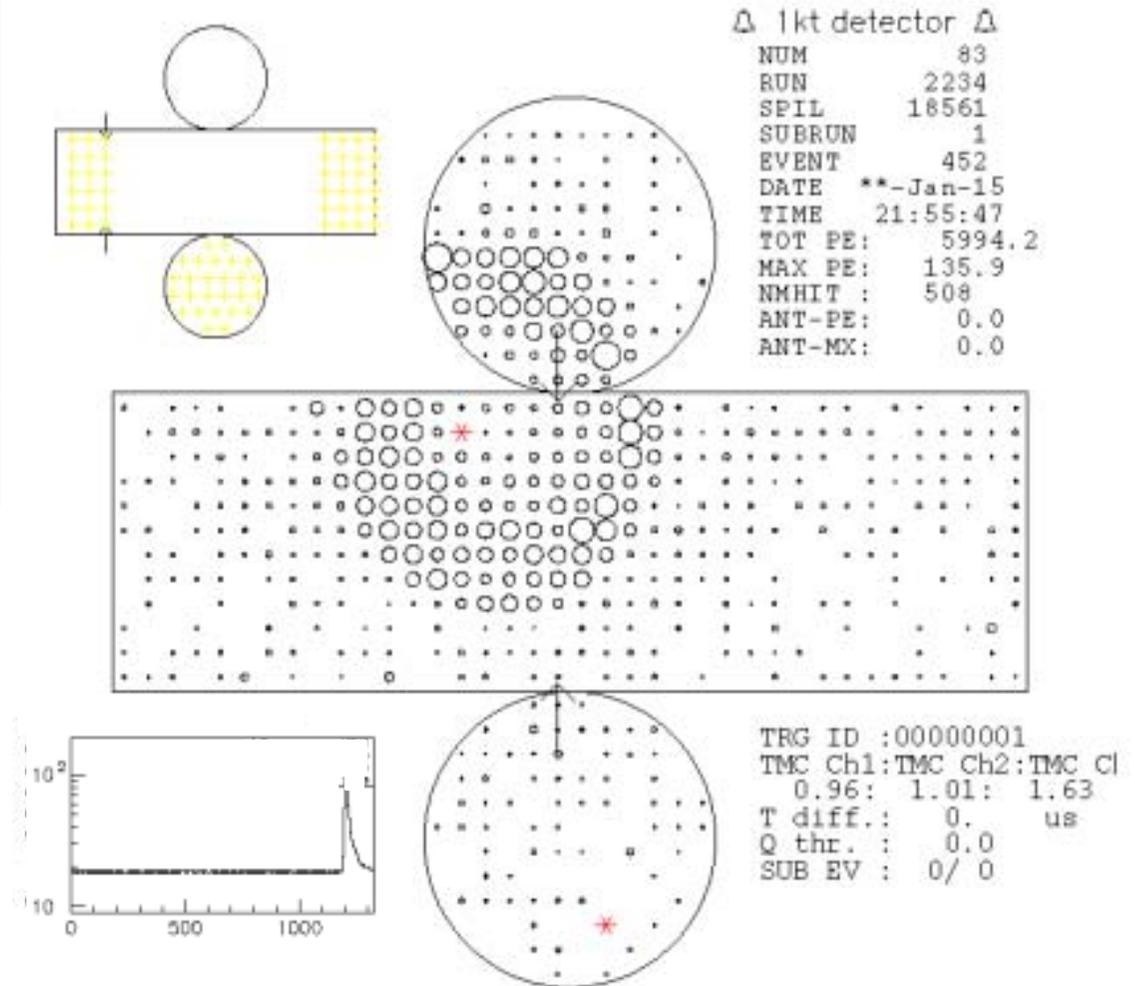


- ▶ 1kt Water Cherenkov detector (KT) fiducial: 25 ton H<sub>2</sub>O
- ▶ Water target/Scintillating fiber tracker (SciFi) 5.9 ton H<sub>2</sub>O
- ▶ Muon range detector (MRD) 700 ton Iron
- ▶ Lead glass detector (LG)

# Water Cherenkov Detector (1kt)

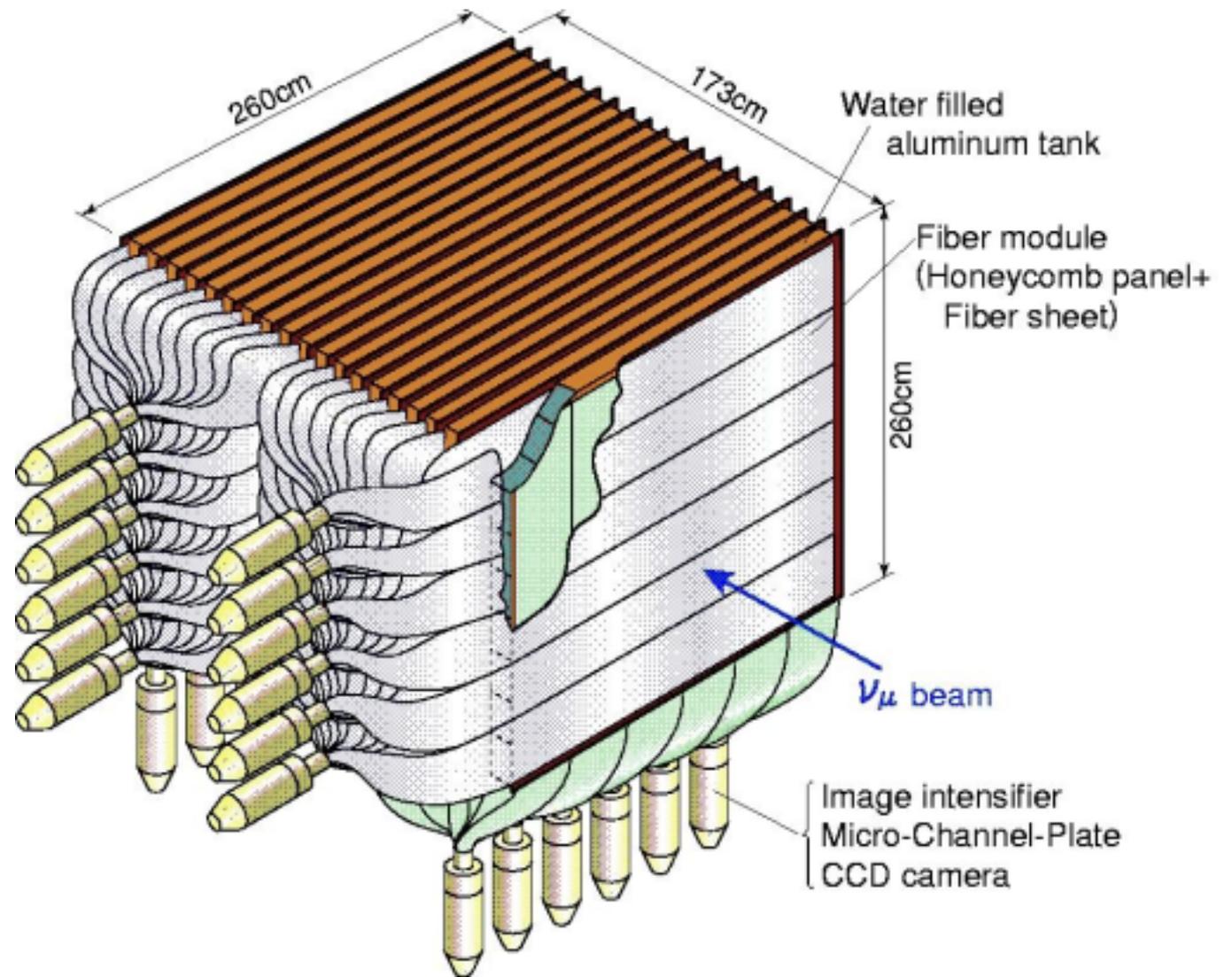
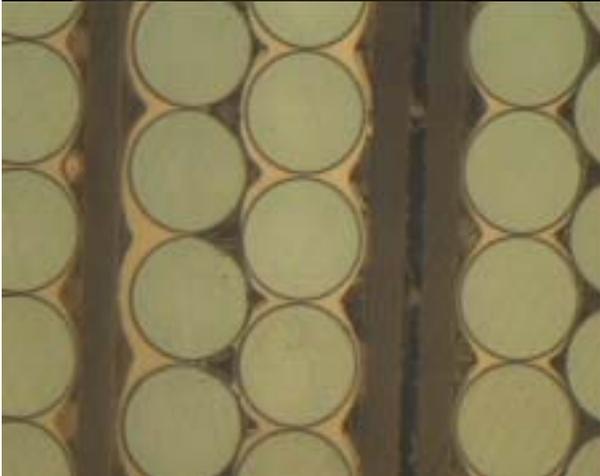
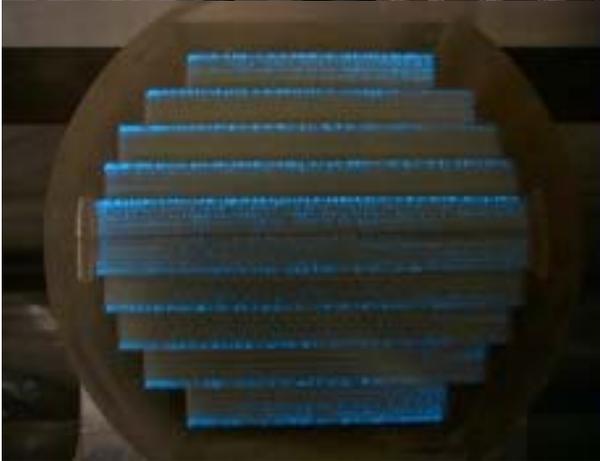


- A miniture of Super-Kamiokande detector with 1/50 volume
- 680 20" PMTs with 70cm spacing (same as SK)
- Inner Volume : 496 tons  
Fiducial Volume : 25.1 tons  
( $r=2\text{m}$  cylindrical volume along beam)

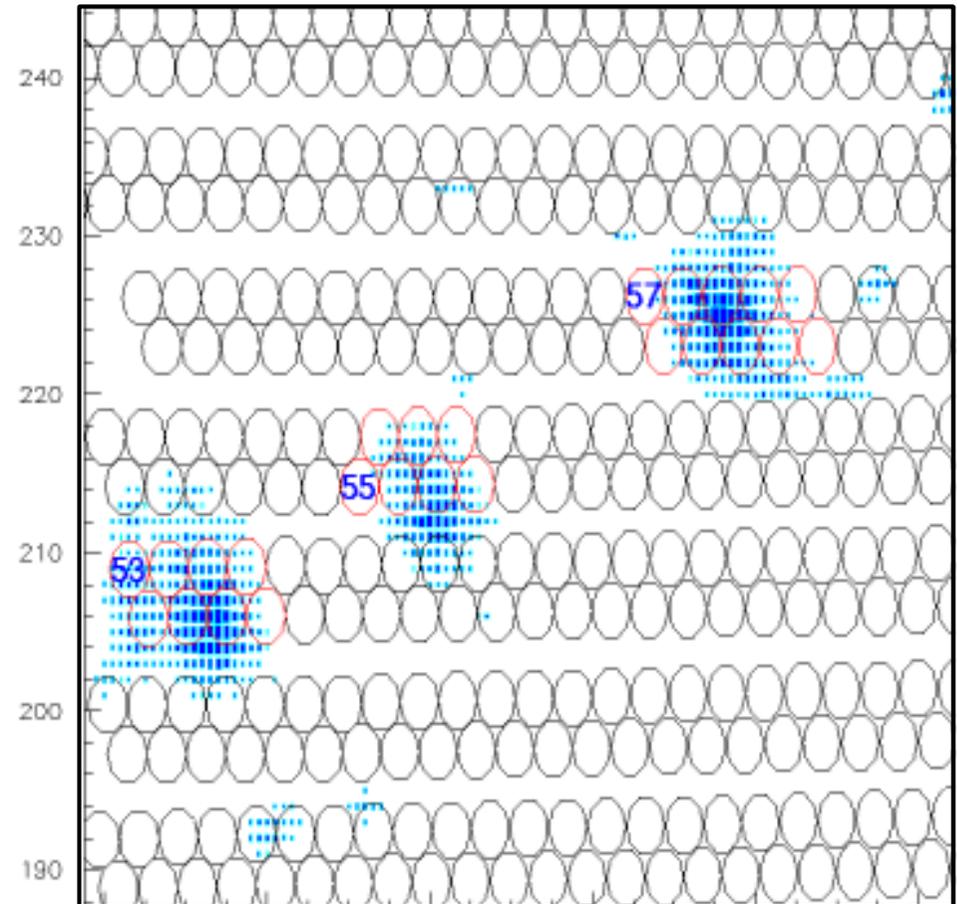
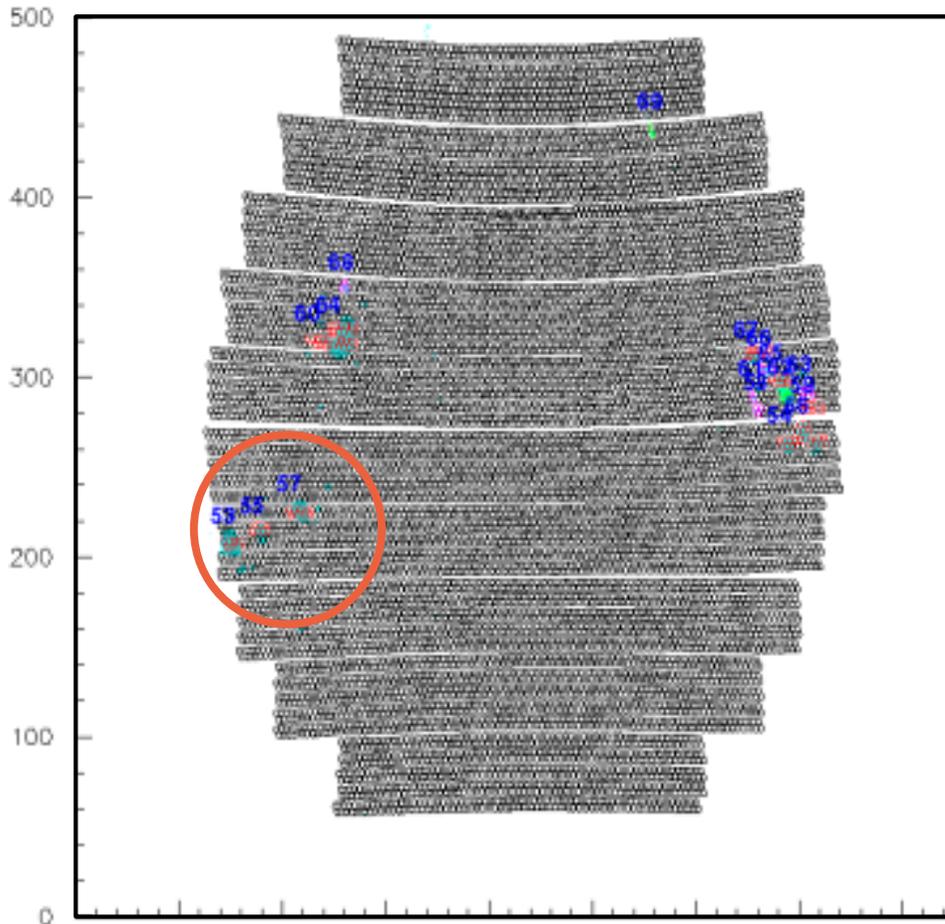


Typical 1R event

# Scintillating Fiber(SciFi) Tracker



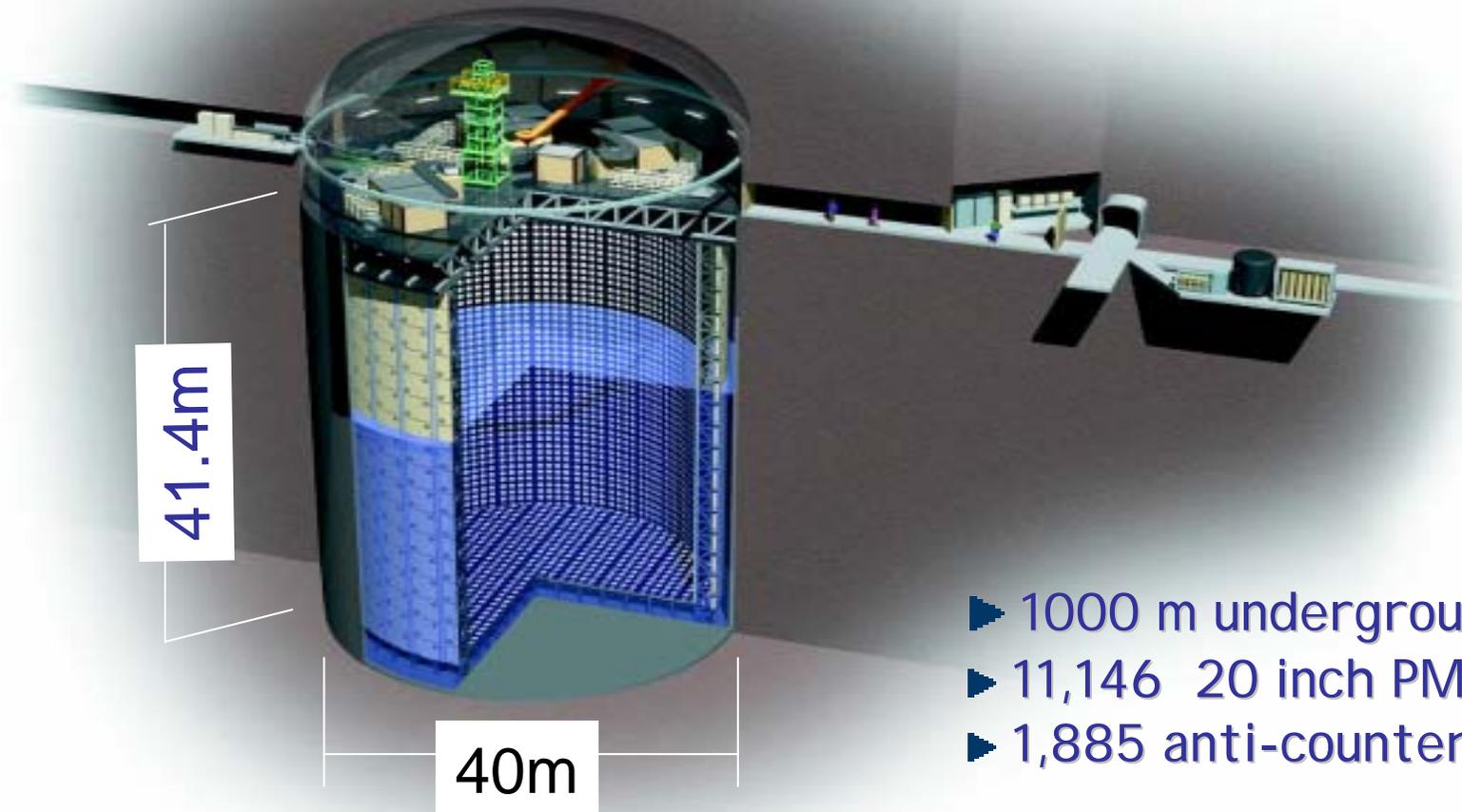
# Typical CCD Pixel Image



# Super-Kamiokande

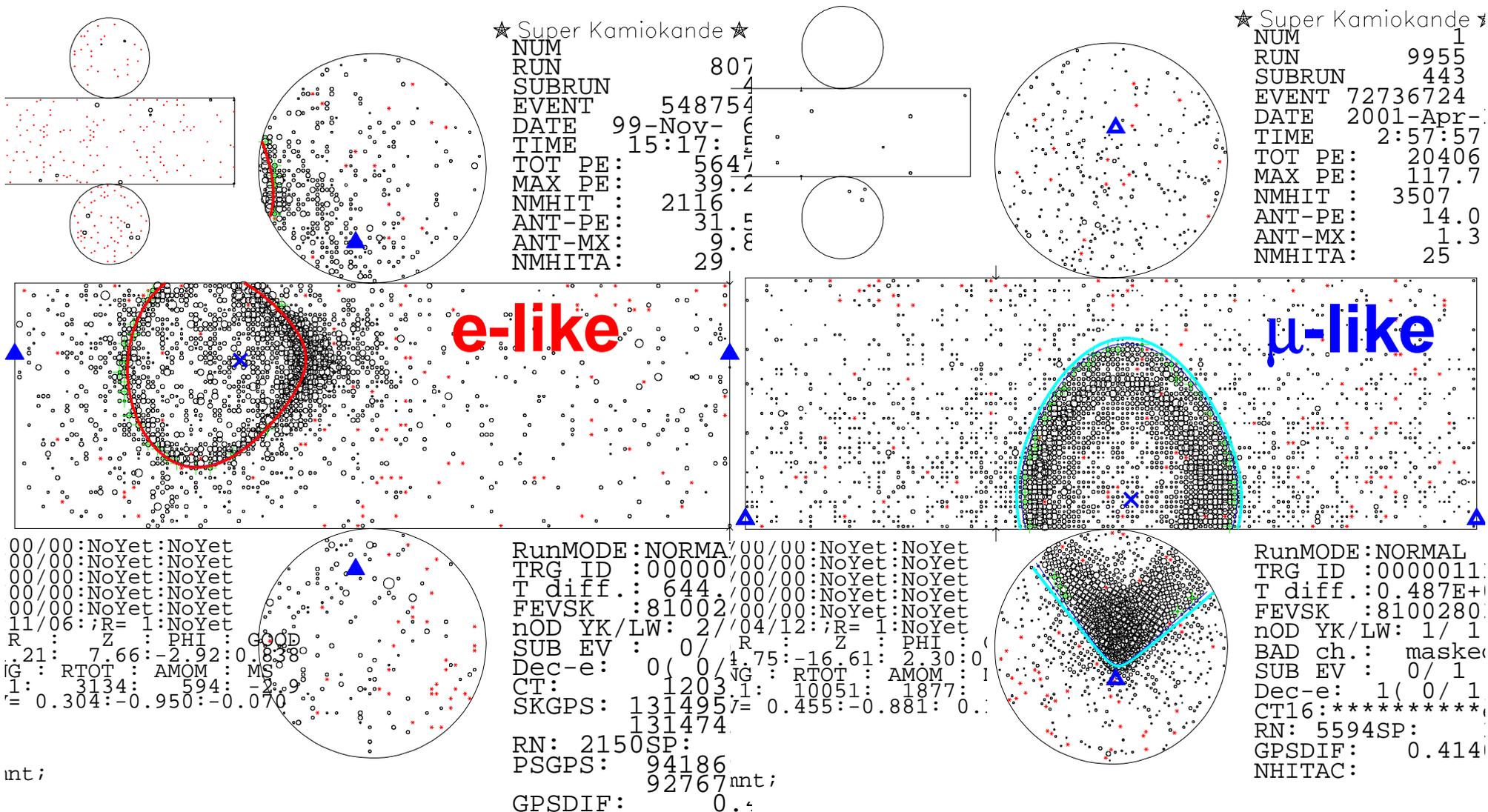
(April 1996 commissioned)

- 50,000 ton water Cherenkov detector (22.5 kton fiducial volume)
- Optically separated **INNER** and **OUTER** detector

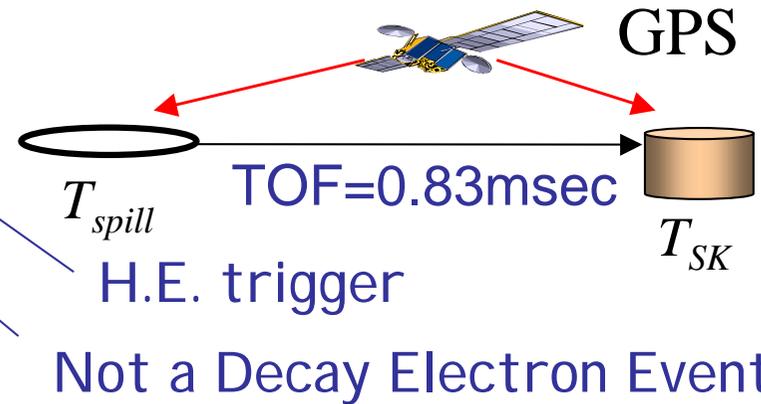
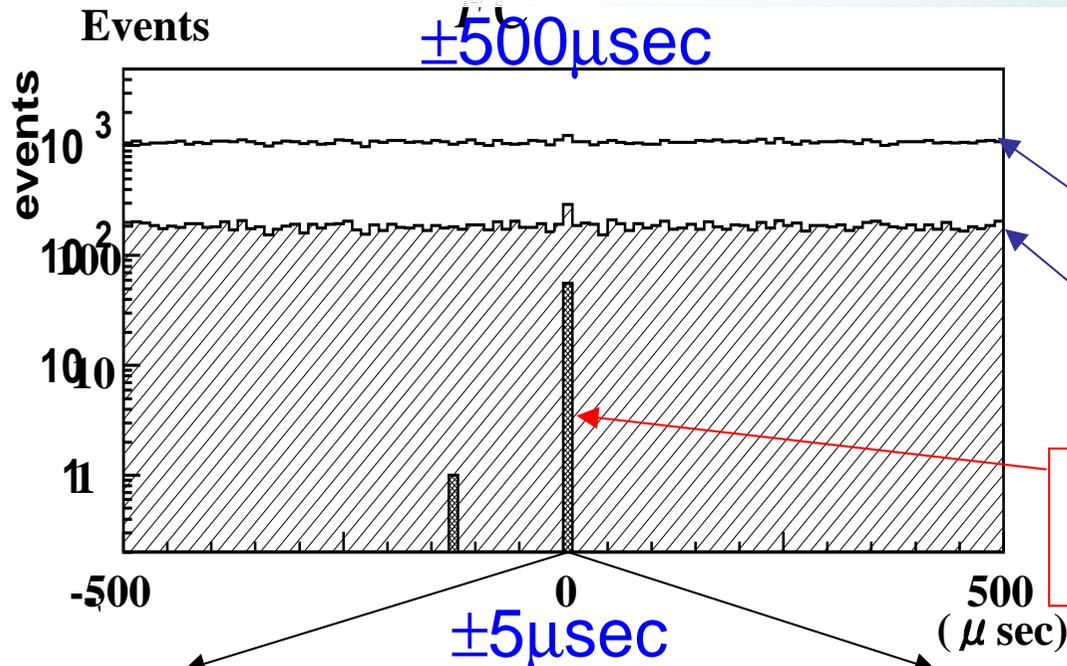


- ▶ 1000 m underground
- ▶ 11,146 20 inch PMTs
- ▶ 1,885 anti-counter PMTs

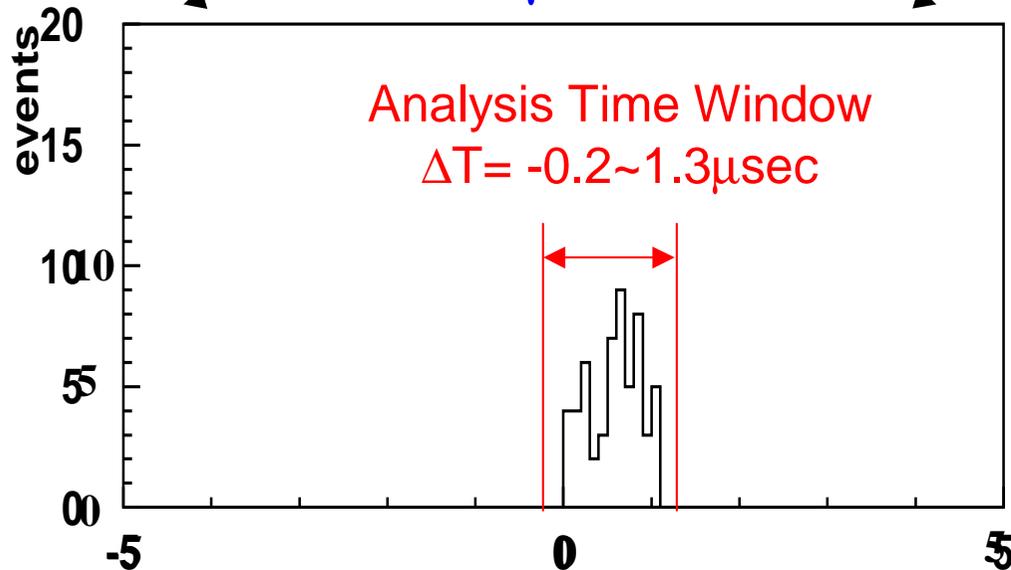
# e-like and $\mu$ -like events



# Super-K Event Selection by GPS



No Activity in Outer Detector (FC)  
Vertex is in Fiducial Volume (FV)



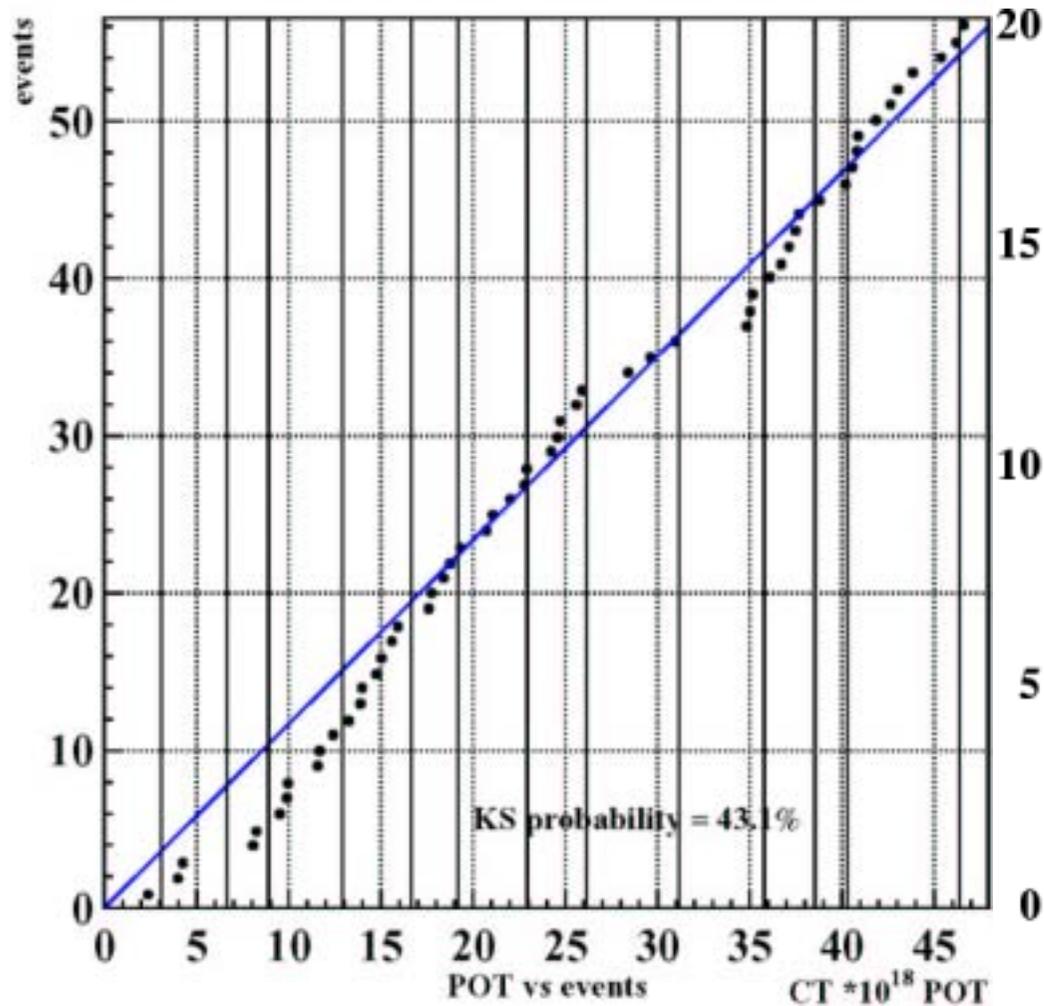
**56 FCFV Events**

where

- ⊕ 1 Ring = 32
- ⊕ μ-like 30 / e-like 2
- ⊕ Multi-Ring = 24

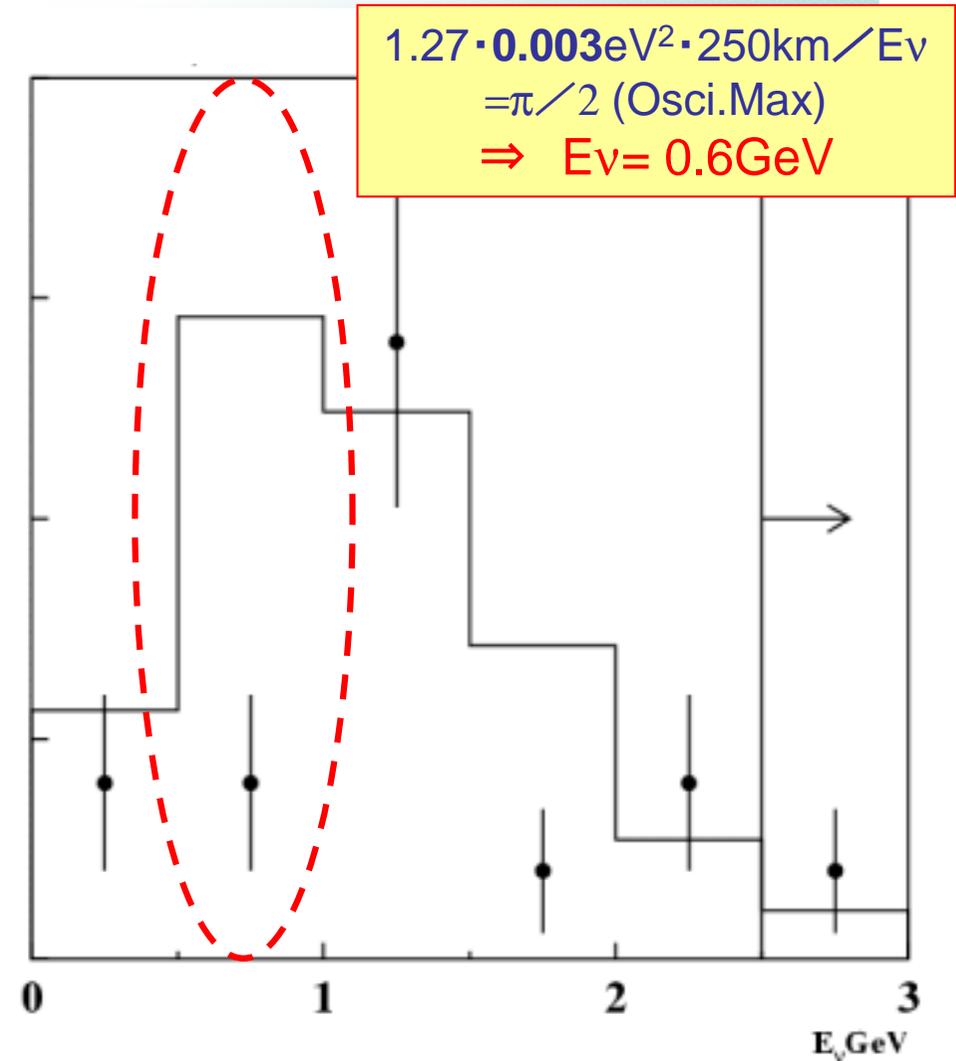
(Expected atm.ν B.G.  $\sim 10^{-3}/1.5\mu s$ )

## Event v.s. POT dist.



KS test probability: 43%

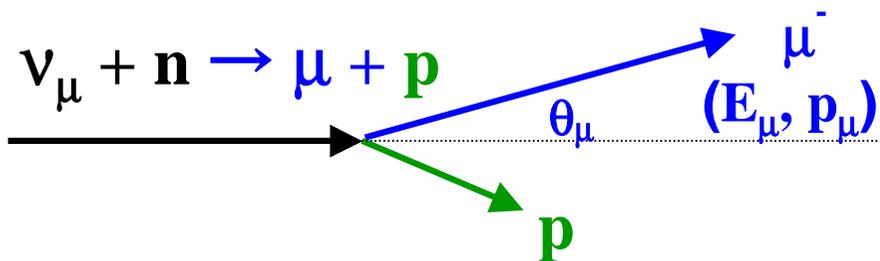
## Evrec of FCFV-1R $\mu$ like



29 1R events with HC250kA

# Neutrino Energy Reconstruction

CC quasi elastic (QE)



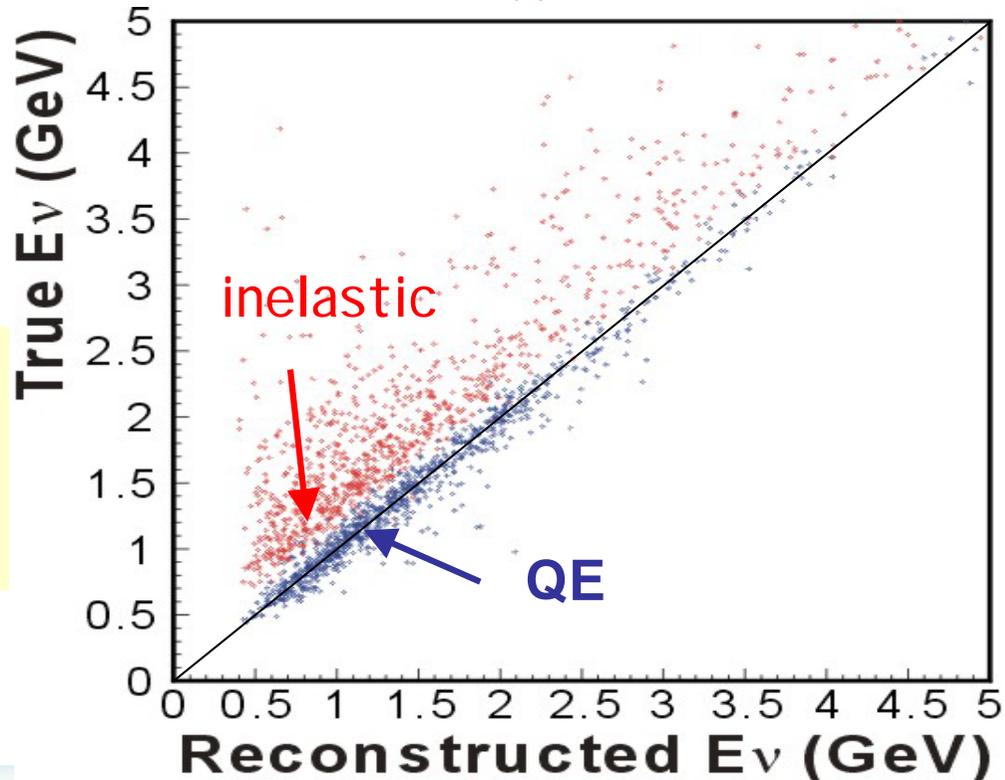
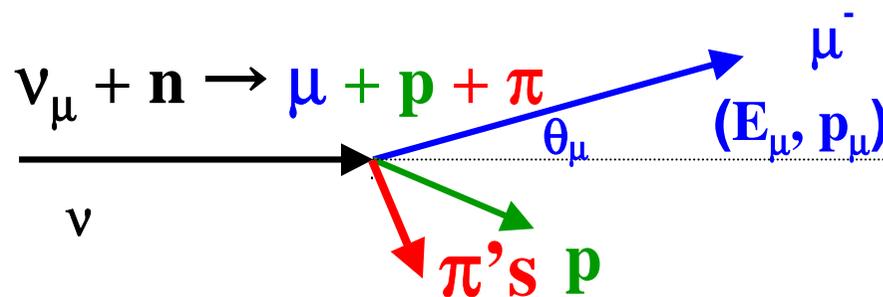
$$E_\nu = \frac{m_N E_\mu - m_\mu^2 / 2}{m_N - E_\mu + p_\mu \cos \theta_\mu}$$

$$\text{Rate}(E_\nu, \text{Near}) \rightarrow \phi(E_\nu, \text{Near})$$

↑

$$\sigma(\text{QE}), \sigma(\text{nonQE})$$

CC inelastic



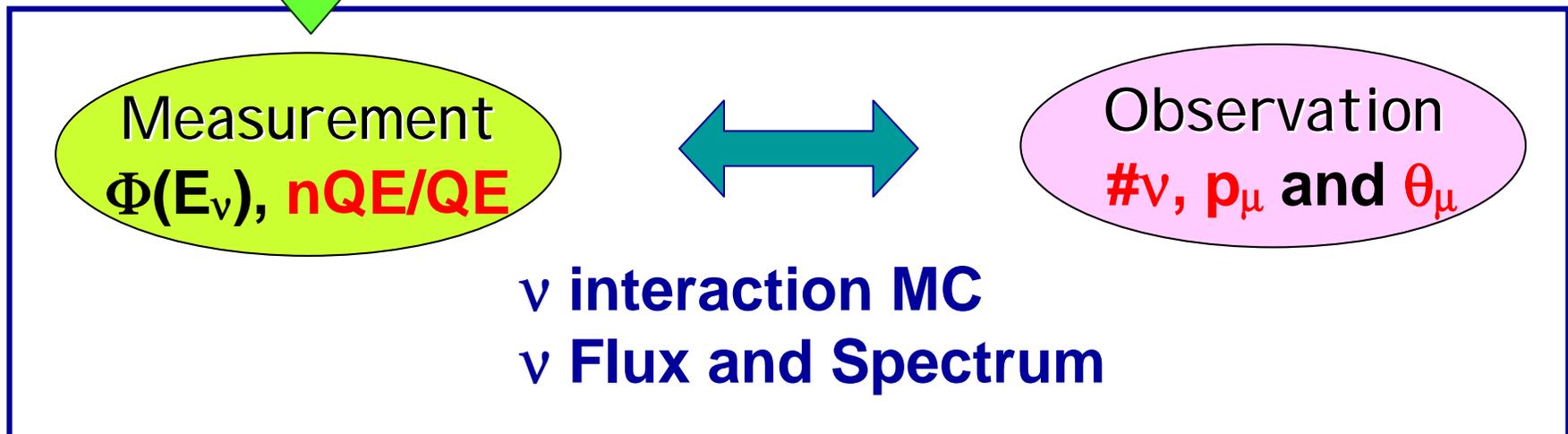
# Oscillation Analysis

Super-K



Far/Near Ratio  
(beam MC &  $\pi$  mon.)

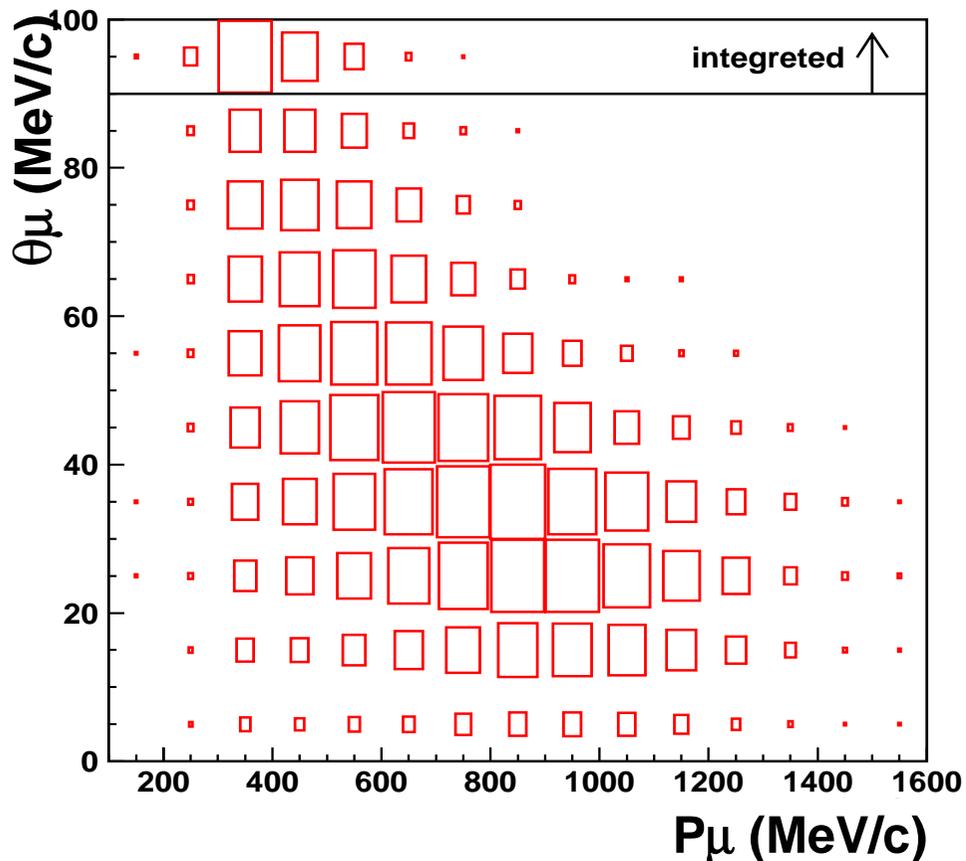
FD @ KEK



# FD @ KEK

$$(p_\mu, \theta_\mu) \rightarrow \phi(E\nu), nQE/QE$$

**(1) KT 1R $\mu$ -like ~36,000 events**



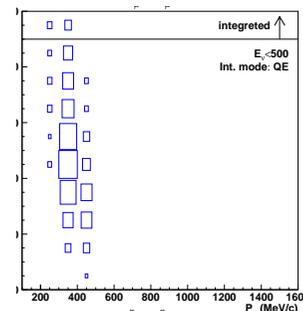
- ▶  **$\nu$  flux  $\phi_{near}(E\nu)$  (8 bins)**
- ▶  **$\nu$  interaction (nQE/QE)**

Ev

QE (MC)

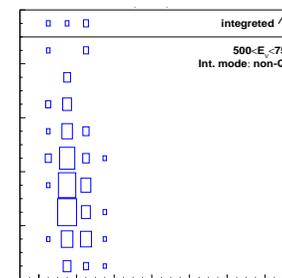
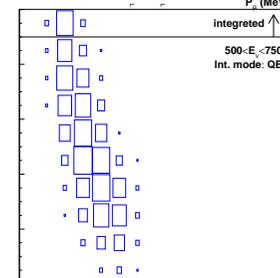
nQE(MC)

0-0.5 GeV

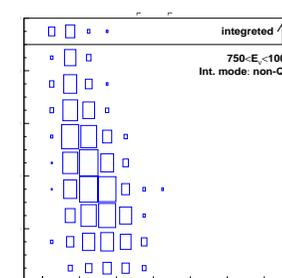
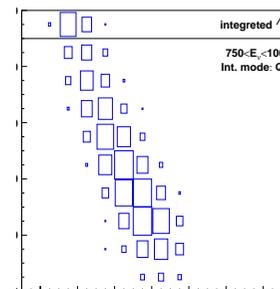


MC templates

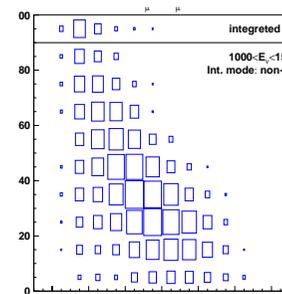
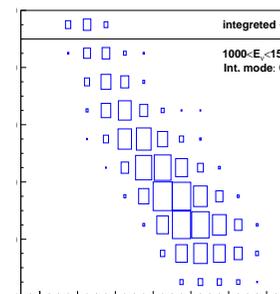
0.5-0.75 GeV



0.75-1.0 GeV



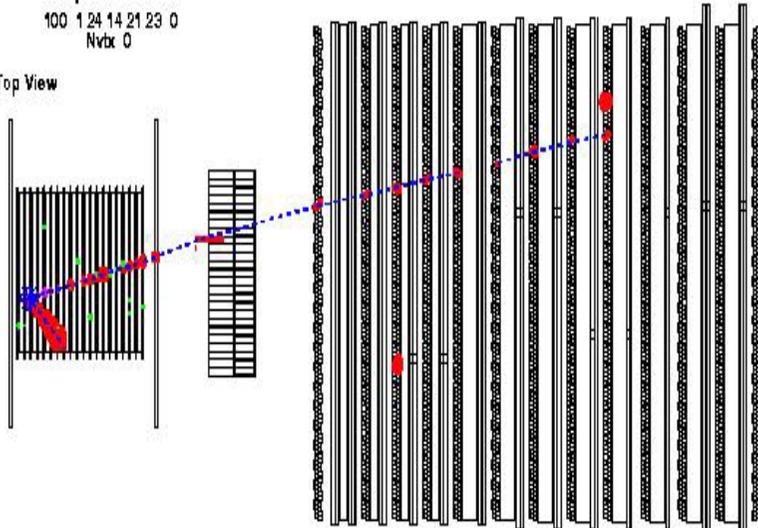
1.0-1.5 GeV



# QE and nQE in SciFi 2-track events

Run 2279 Spill 18568 TRGID 1  
100 1 24 14 21 23 0  
Nvtx 0

Top View



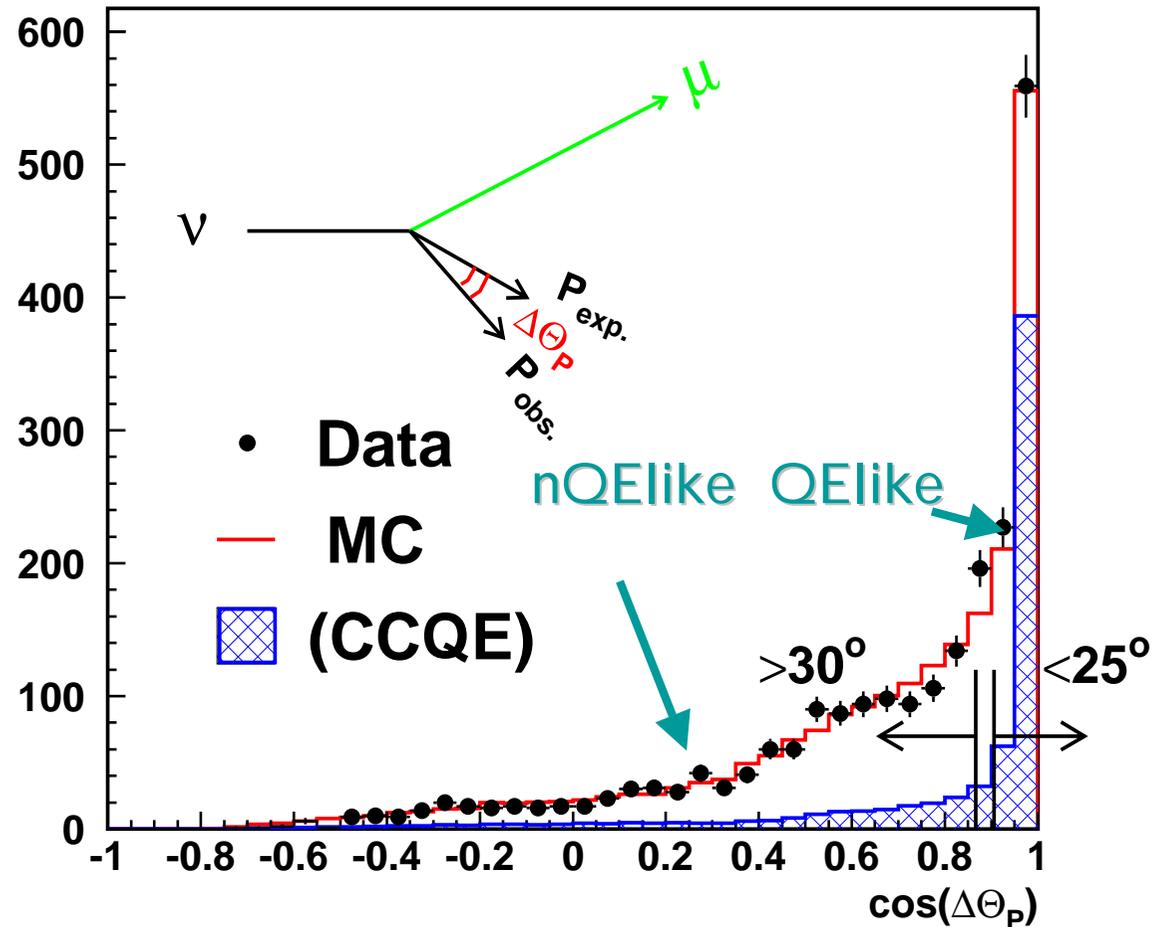
- (2) 1-track  
5,951events/44points
- (3) 2-track ( $\Delta\theta_P \leq 25^\circ$ )  
761events/40points
- (4) 2-track ( $\Delta\theta_P \geq 30^\circ$ )  
1,291events/40points

124 data points

~ 8,000 events in total.

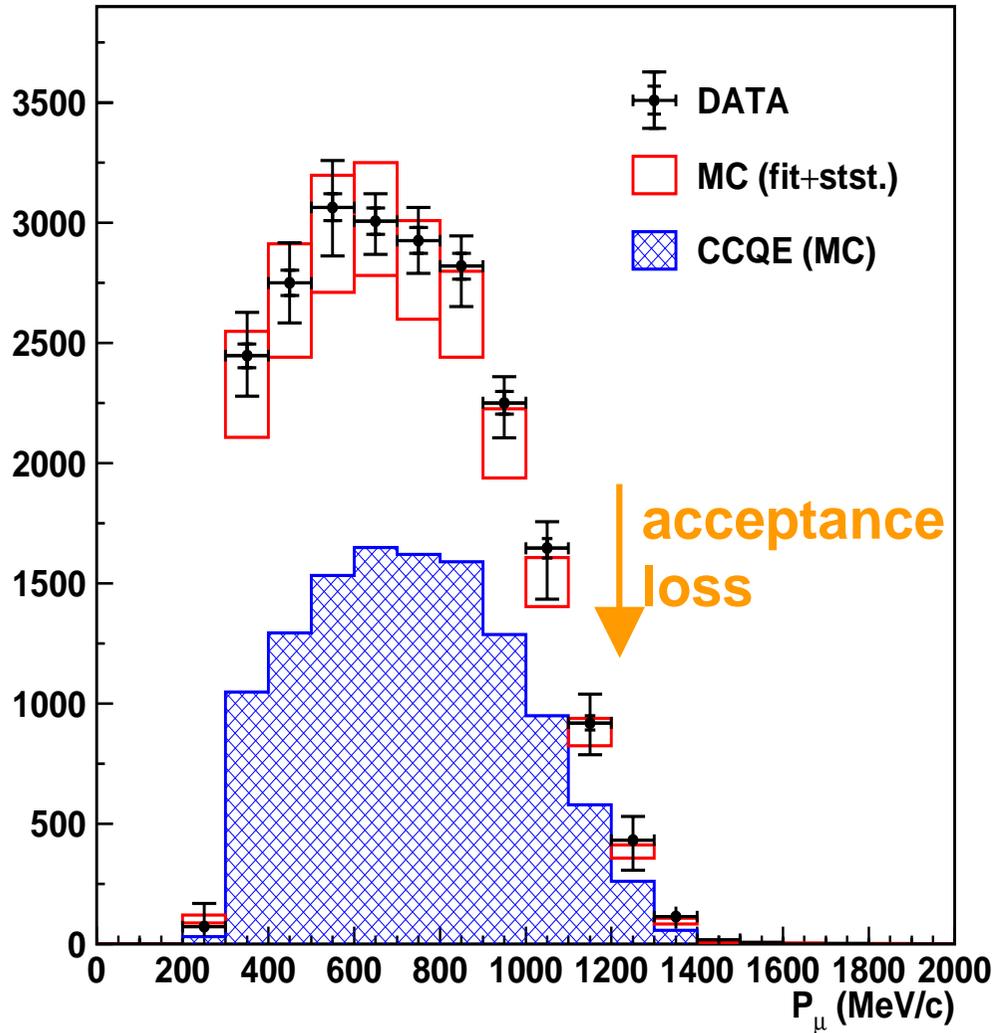
R(non-QE/QE) is constrained by (3) / (2)

$$\Delta R < 10\%$$

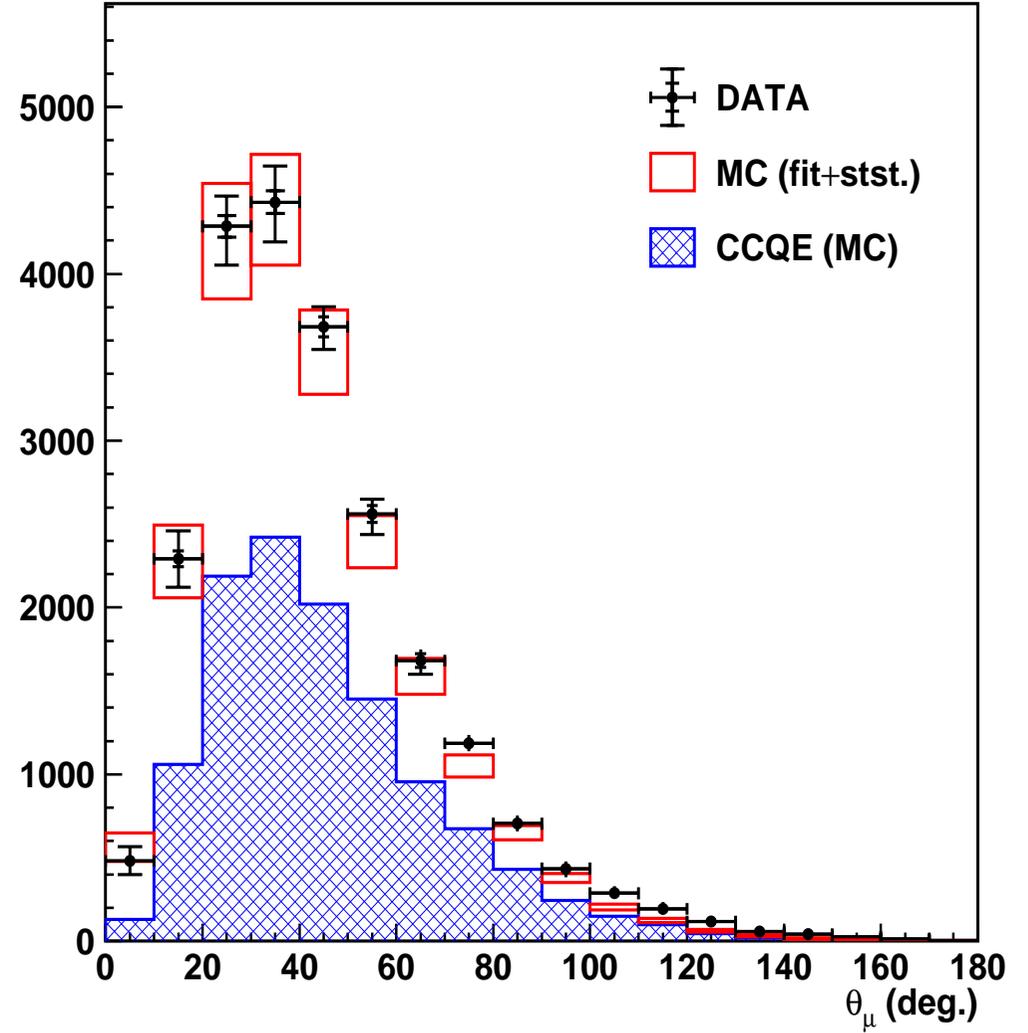


# (1) KT ( $p_\mu, \theta_\mu$ ) distribution using $\phi_{\text{fit}}, \text{QE}/n\text{QE}_{\text{fit}}$

$P_\mu$



$\theta_\mu$

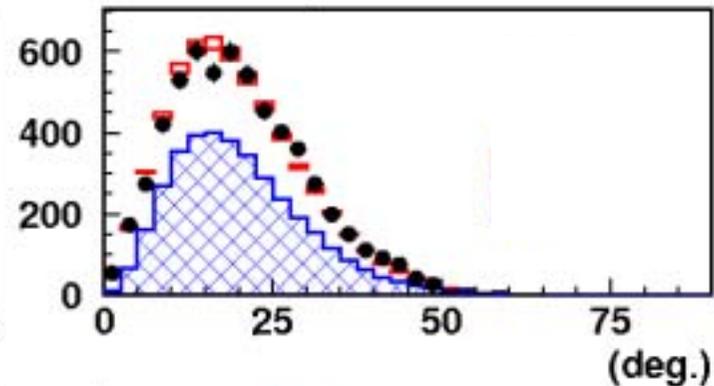
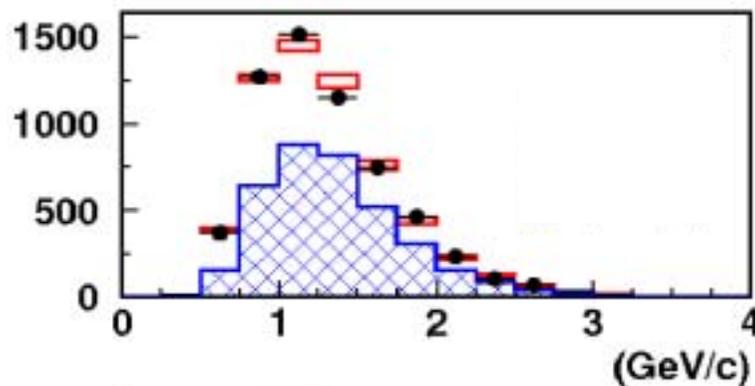


# SciFi ( $p_\mu, \theta_\mu$ ) distributions

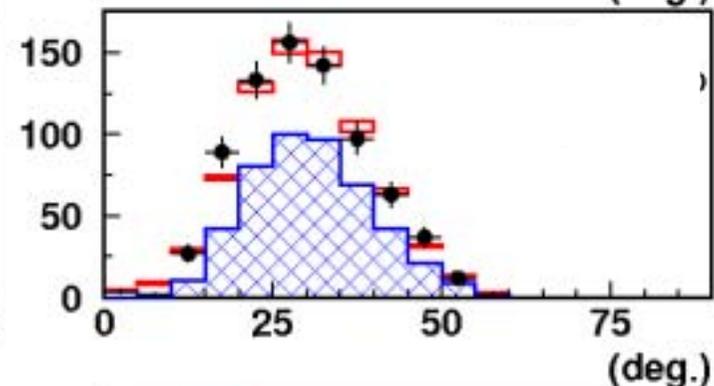
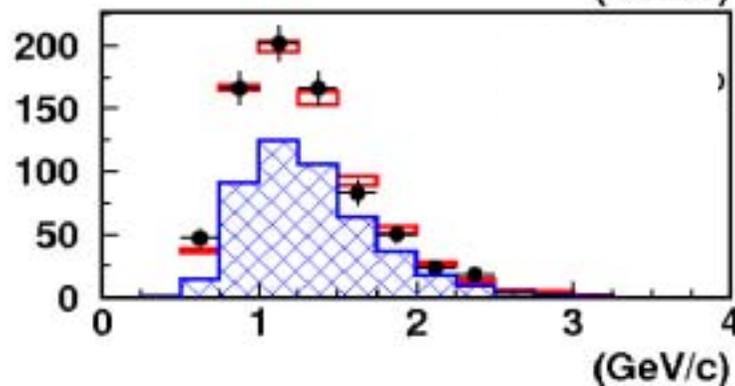
$P_\mu$

$\Theta_\mu$

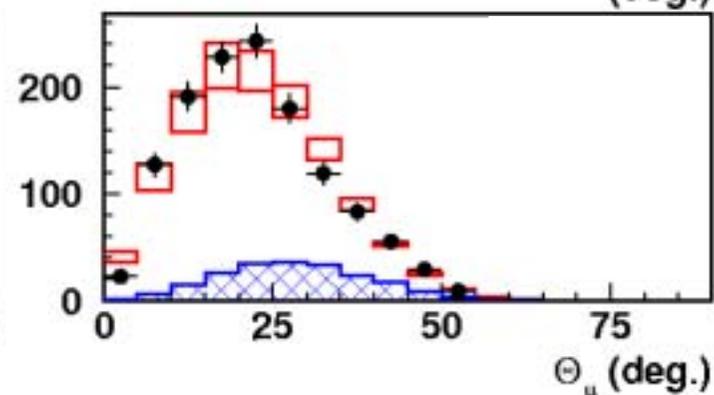
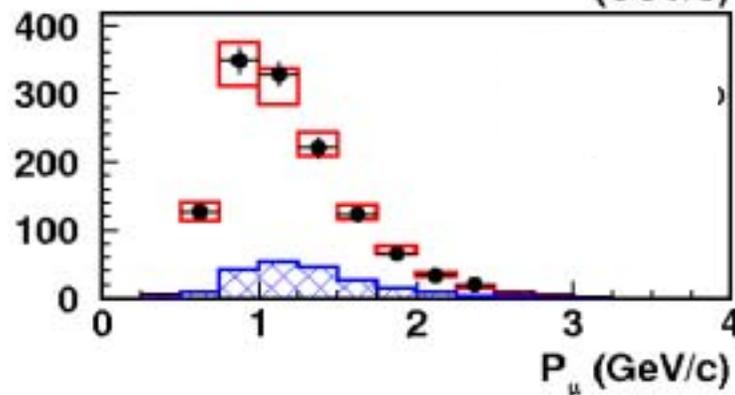
(2) 1-track



(3) 2-track  
 $\Delta\Theta_P \leq 25^\circ$

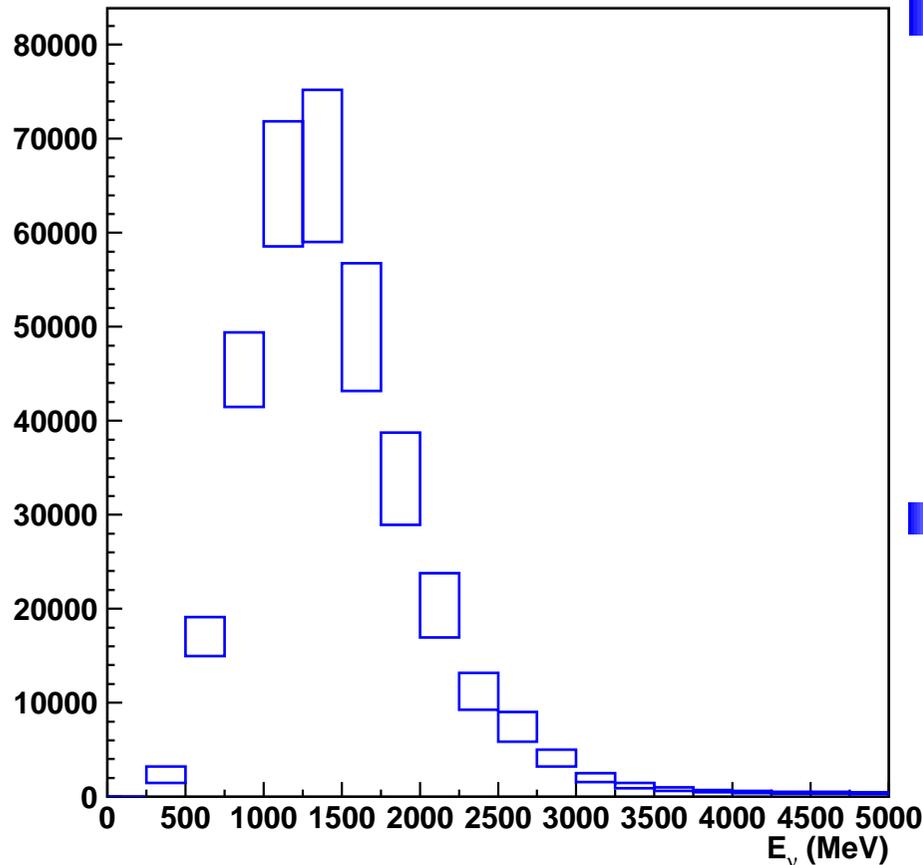


(4) 2-track  
 $\Delta\Theta_P \geq 30^\circ$



# Fit result of Neutrino Flux at KEK Site

Fitted  $E_\nu$  Flux at Front Detector



$E_\nu$  (MeV)

■  $\chi^2=227$  for 194 *d.o.f*

- ◆  $\Phi_1$  ( $E_\nu < 500$ ) =  $1.3 \pm 0.5$
- ◆  $\Phi_2$  ( $500 \leq E_\nu < 750$ ) =  $1.02 \pm 0.12$
- ◆  $\Phi_3$  ( $750 \leq E_\nu < 1000$ ) =  $1.01 \pm 0.09$
- ◆  $\Phi_4$  ( $1500 \leq E_\nu < 2000$ ) =  $0.95 \pm 0.07$
- ◆  $\Phi_5$  ( $2000 \leq E_\nu < 2500$ ) =  $0.96 \pm 0.08$
- ◆  $\Phi_5$  ( $2500 \leq E_\nu < 3000$ ) =  $1.18 \pm 0.19$
- ◆  $\Phi_6$  ( $3000 \leq E_\nu$ ) =  $1.07 \pm 0.20$

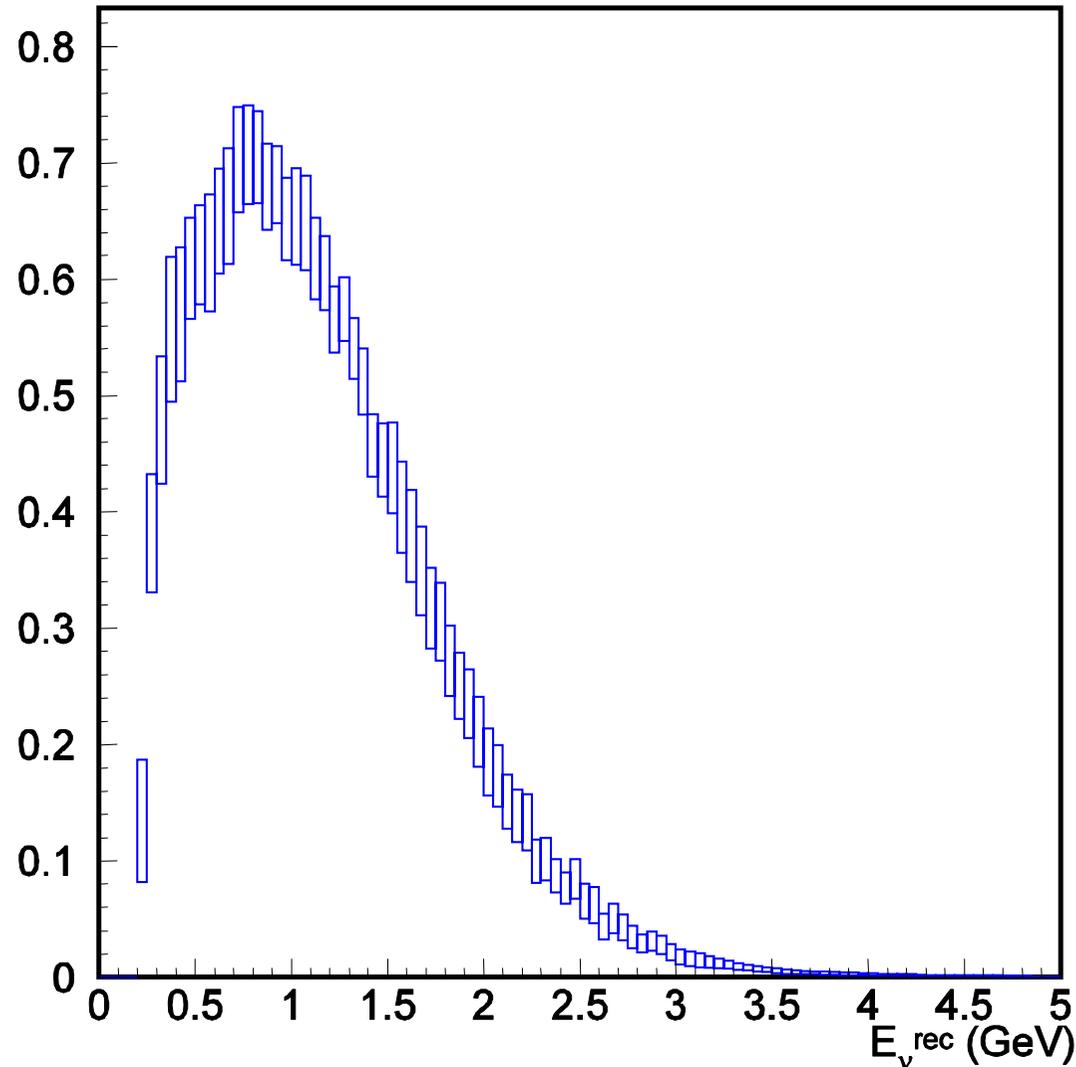
■  $nQE/QE = 0.93 \pm 0.20$  (0.06)

Error is assigned based on the disagreement between SciFi and KT results.

- ◆ KT only:  $nQE/QE=0.73$
- ◆ SciFi only:  $nQE/QE=1.09$

# Super-K: Expected $E_{\nu}^{\text{rec}}$ spectrum for $1R_{\mu}$

Initial  $1R_{\mu}$  spectrum w/ all syst. err. incl. Escale



## Maximum likelihood fit with $(\sin^2 2\theta, \Delta m^2)$

### Likelihood

$$L_{tot} = L_{norm} \cdot L_{shape} \cdot L_{syst}$$

### Term for # of FCFV events

$$L_{norm} = \text{Poisson}(N_{obs}, N_{exp'ed}(\Delta m^2, \sin^2 2\theta, f_{syst}))$$

$N_{obs}$ : Observed number of FCFV events (56)

$$N_{exp'ed}(\Delta m^2, \sin^2 2\theta, f_{syst}) = \frac{N_{obs}}{N_{KT}^{MC}(\Delta m^2, \sin^2 2\theta, f_{syst})} \cdot \frac{N_{SK}^{MC}(\Delta m^2, \sin^2 2\theta, f_{syst})}{N_{KT}^{MC}(f_{syst})}$$

(exp'ed # of FCFV events)

### Term for $E_{\nu}^{rec}$ distribution for $1R\mu$ events

$$L_{shape} \equiv \prod_{i=1}^{29} P(E_i^{rec}; \Delta m^2, \sin^2 2\theta, f_{syst})$$

$P$ : normalized  $E_{\nu}^{rec}$  distribution for  $1R\mu$  events estimated by MC simulation

# Systematic parameters

$$\begin{aligned} L_{syst} \equiv & \exp\left(-\Delta f_{\Phi,nQE}^T \cdot M_{FD}^{-1} \cdot \Delta f_{\Phi,nQE} / 2\right) \\ & \times \exp\left(-\Delta f_{F/N}^T \cdot M_{F/N}^{-1} \cdot \Delta f_{F/N} / 2\right) \\ & \times \exp\left(-\Delta f_{\varepsilon SK}^T \cdot M_{\varepsilon SK}^{-1} \cdot \Delta f_{\varepsilon SK} / 2\right) \\ & \times \exp\left(-\Delta f_{n6}^2 / 2\sigma_{n6}^2\right) \\ & \times \exp\left(-\Delta f_{n11}^2 / 2\sigma_{n11}^2\right) \\ & \times \exp\left(-\Delta f_{Esk}^2 / 2\sigma_{Esk}^2\right) \end{aligned}$$

$$\Delta f = f - f_{cent}$$

- $M_{FD}$ ,  $M_{\pi}$ ,  $M_{\varepsilon SK}$ : error matrices(spec+nQE/QE, far/near,  $\varepsilon_{SK}$ )
- $\sigma_{n6}$ : norm. err. for Jun99 (=+0.80–0.68 evts)
- $\sigma_{n11}$ : norm. err. for Nov99~ (=5.34% dominated by KT/SK fid. vol. err.)
- $\sigma_{Esk}$ : SK energy scale error (3%)

# Expected number of FCFV events w/o oscillation

Generate many sets of random numbers for  $f_{\text{syst}}$  which distribute according to the error matrices

Calculate

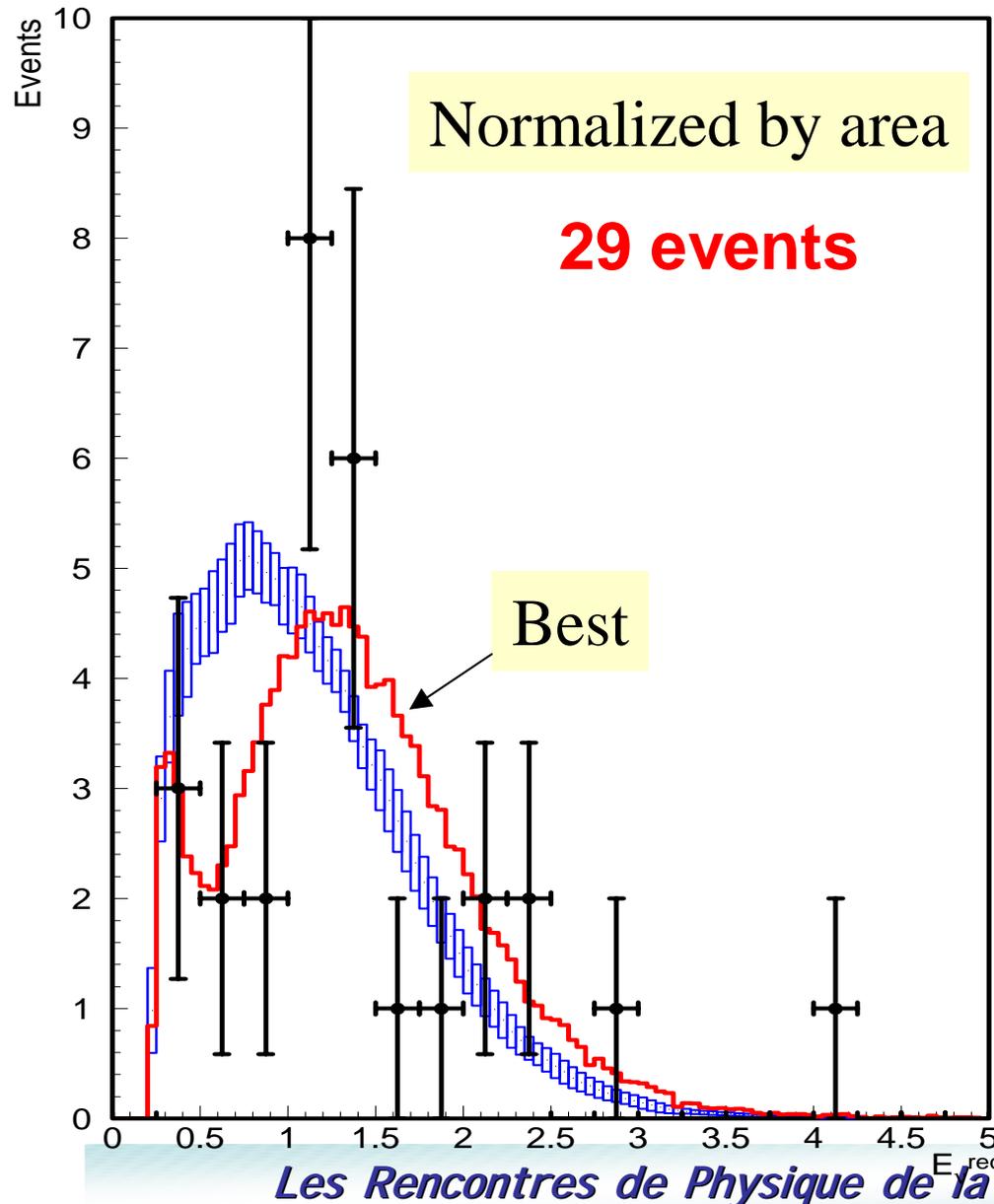
$$N_{\text{exp'ed}}(\Delta m^2, \sin^2 2\theta, f_{\text{syst}})$$

for each set w/  $\sin^2 2\theta=0$

**80.1<sup>+6.2</sup><sub>-5.4</sub> events**

Jun99	Total	+1.0% -0.85%
Nov99~	Spectrum	+0.56% -0.63%
	nQE/QE	+0.47% -1.1%
	Far/Near	+4.9% -5.0%
	Norm	5.0%
<b>Total</b>		<b>+7.7%</b> <b>-6.7%</b>

# Best fit 1R $\mu$ spectrum & N<sub>sk</sub>



**Best fit point ( $\sin^2 2\theta$ ,  $\Delta m^2$ )**  
**= (1.0,  $2.8 \times 10^{-3} \text{eV}^2$ )**

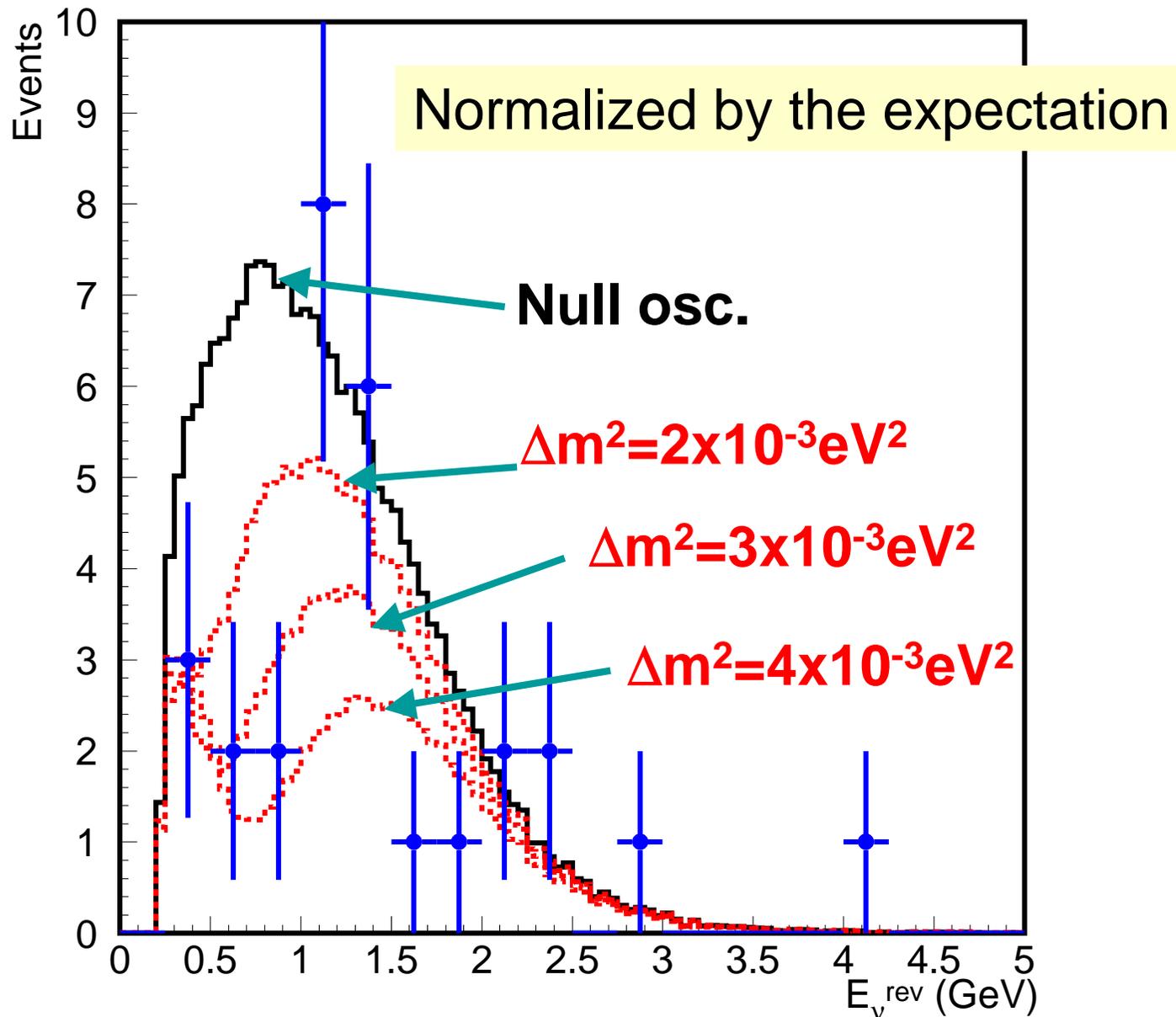
KS test prob.(shape): 79%

**$N_{SK} = 54.2$  (Obs.=56)**

**Very good agreement**  
**Shape &  $N_{SK}$**

# 1R $\mu$ spectrum with $\Delta m^2$

1R $\mu$  Observation and Default Spectrum



# Result

## Null Oscillation Probability

	analysis-1	analysis-2
$N_{SK}$ only	1.3%	0.7%
Shape only	15.7%	14.3%
$N_{SK}$ +Shape	0.7%	0.4%

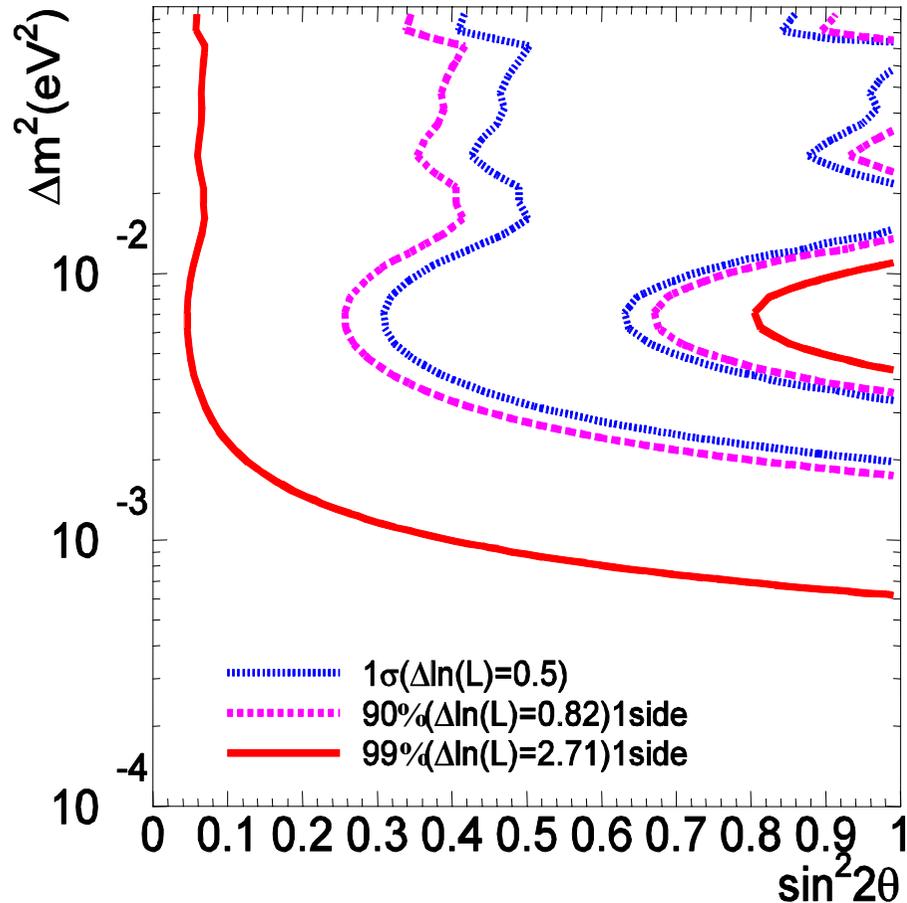
## Best fit ( $\sin^2 2\theta$ , $\Delta m^2$ )

Shape only	(1.0, $3.0 \times 10^{-3} eV^2$ )	(1.0, $3.2 \times 10^{-3} eV^2$ )
$N_{SK}$ +Shape	(1.0, $2.8 \times 10^{-3} eV^2$ )	(1.0, $2.7 \times 10^{-3} eV^2$ )

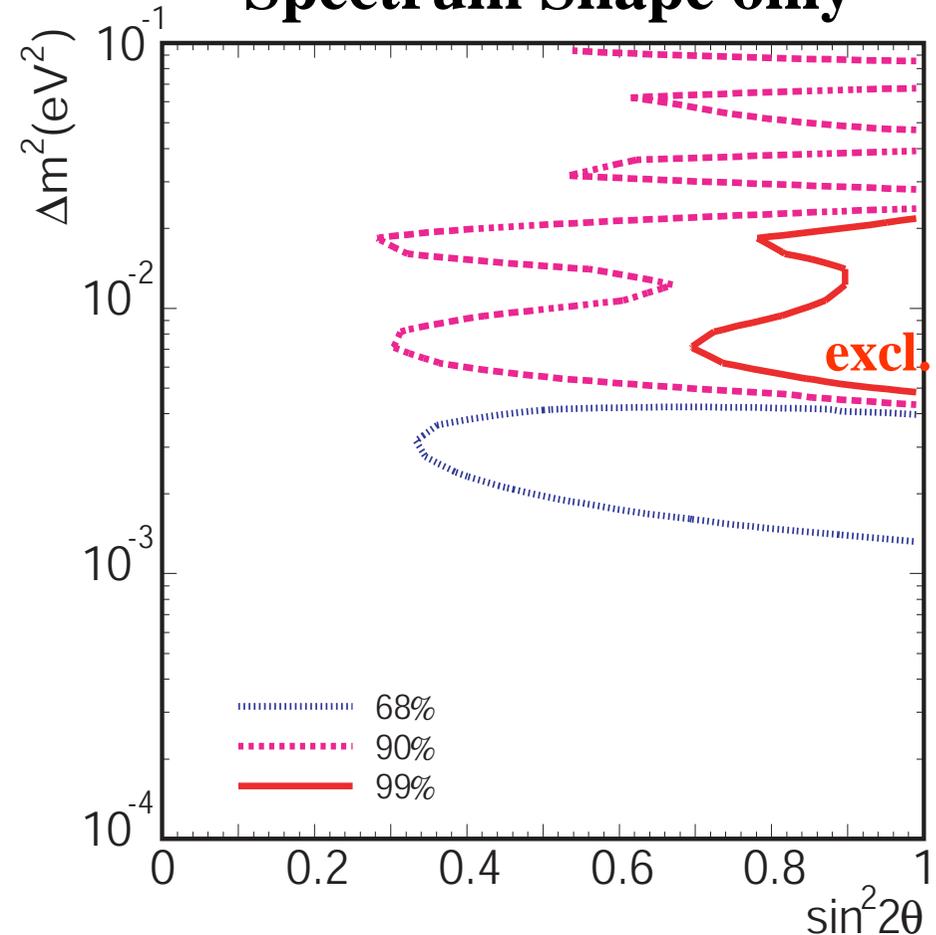
Both Shape and  $N_{SK}$  +Shape indicate consistent parameter region

# Allowed regions

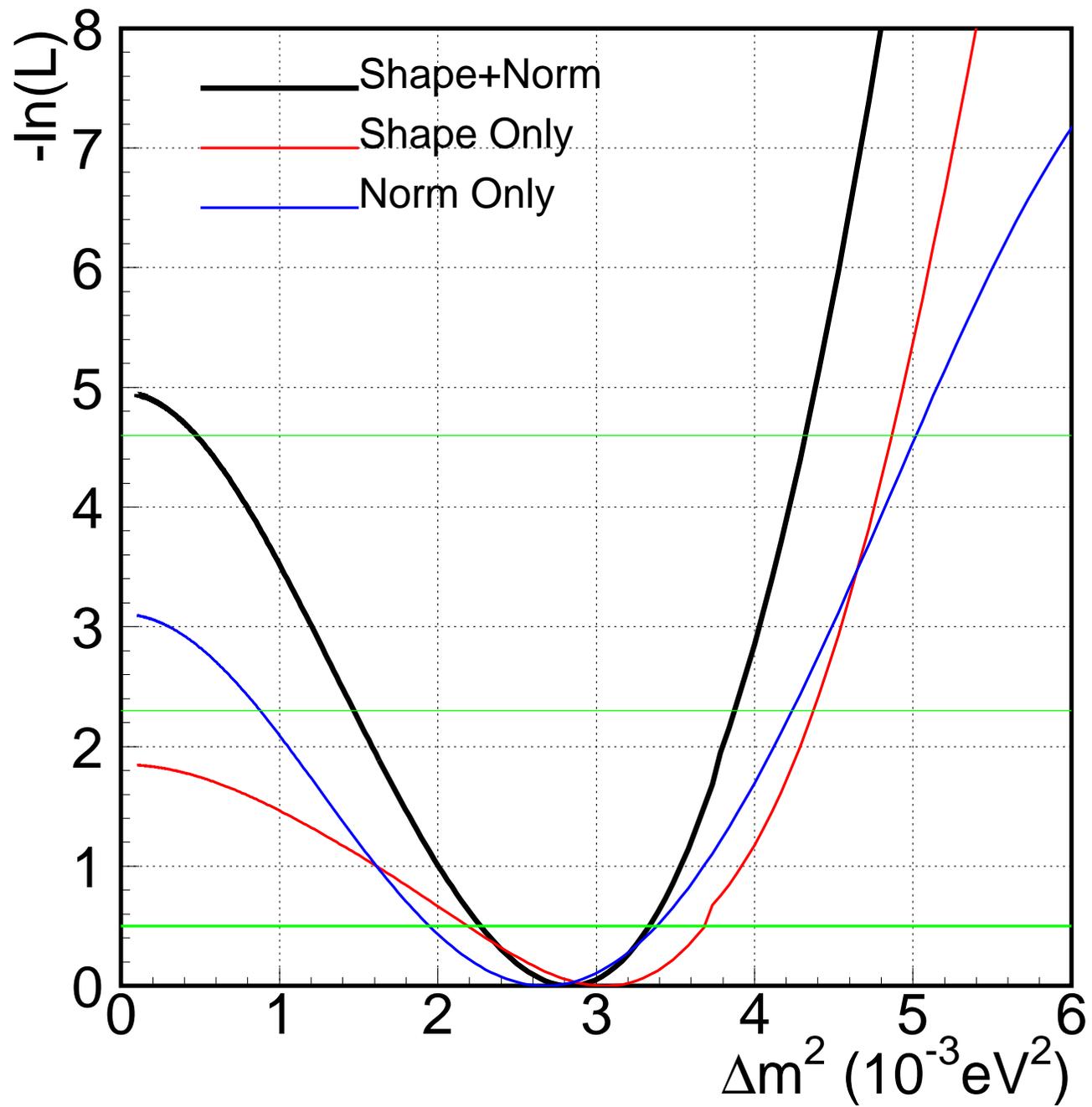
## Total no. of Events only



## Spectrum Shape only

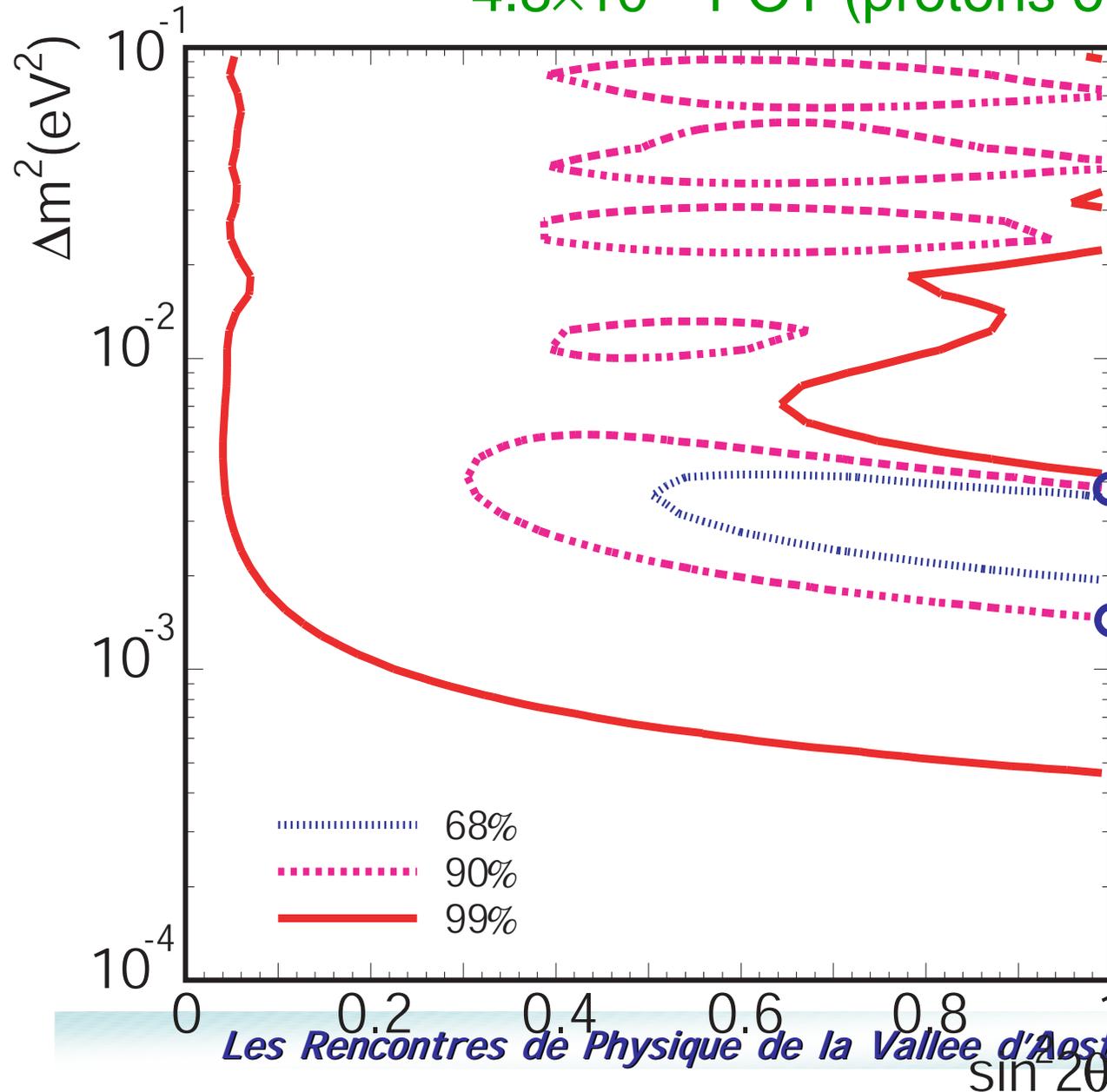


**Both indicate consistent  $\Delta m^2$  region**



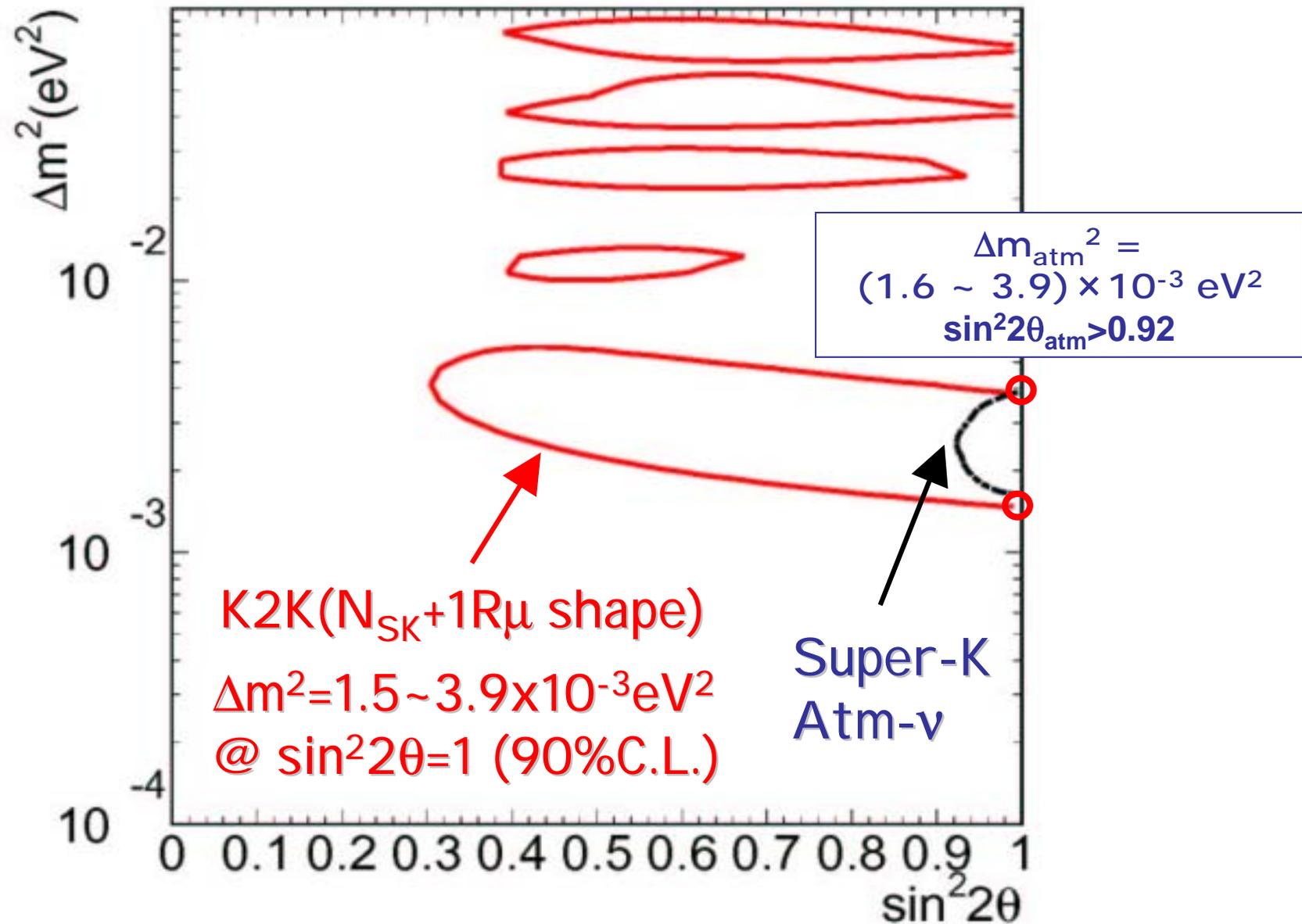
# Allowed region (Shape+Norm)

$4.8 \times 10^{19}$  POT (protons on target)



$\Delta m^2 =$   
 $1.5 \sim 3.9 \times 10^{-3} \text{eV}^2$   
@  $\sin^2 2\theta = 1$   
@ 90% CL

# 90%CL Allowed Regions of K2K and SK atm-ν

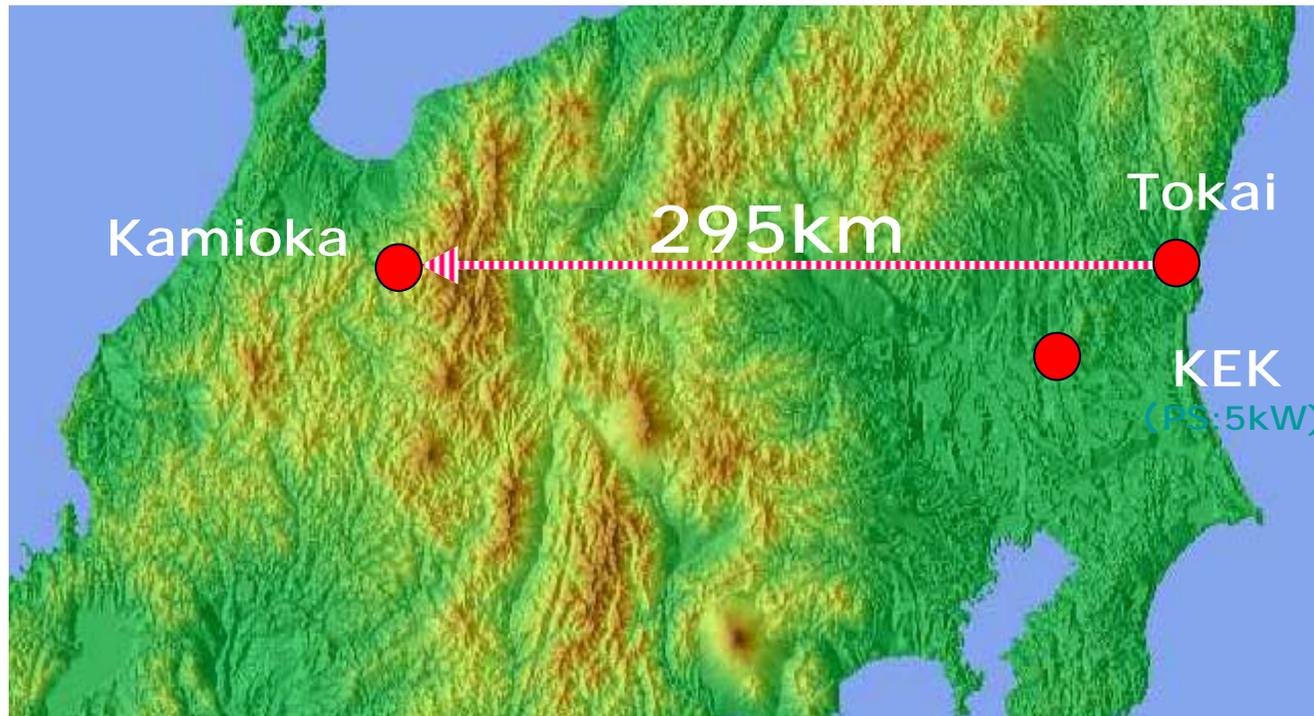




## Conclusion

- K2K Oscillation analysis on June99 ~July01 data(K2K-I)
  1. Null oscillation probability is **less than 1%**.
  2. Both SK rate reduction and  $E_\nu^{\text{rec}}$  shape indicate consistent oscillation parameters region.
  3.  $\Delta m^2 = 1.5 \sim 3.9 \times 10^{-3} \text{eV}^2$  for  $\sin^2 2\theta = 1$  @ 90%CL
  4.  $\sin^2 2\theta$ ,  $\Delta m^2$  are consistent with atmospheric neutrino results
- Data taking has been resumed successfully (2002/12/22~, **K2K-II**).
- Goal is to accumulate  $10^{20}$  protons on target, Twice as large as this data sample.

# JHF-Kamioka $\nu$ Project



■ *JHF 50GeV PS* → *Super-Kamiokande*

(0.75MW)

(22.5kt fid.)

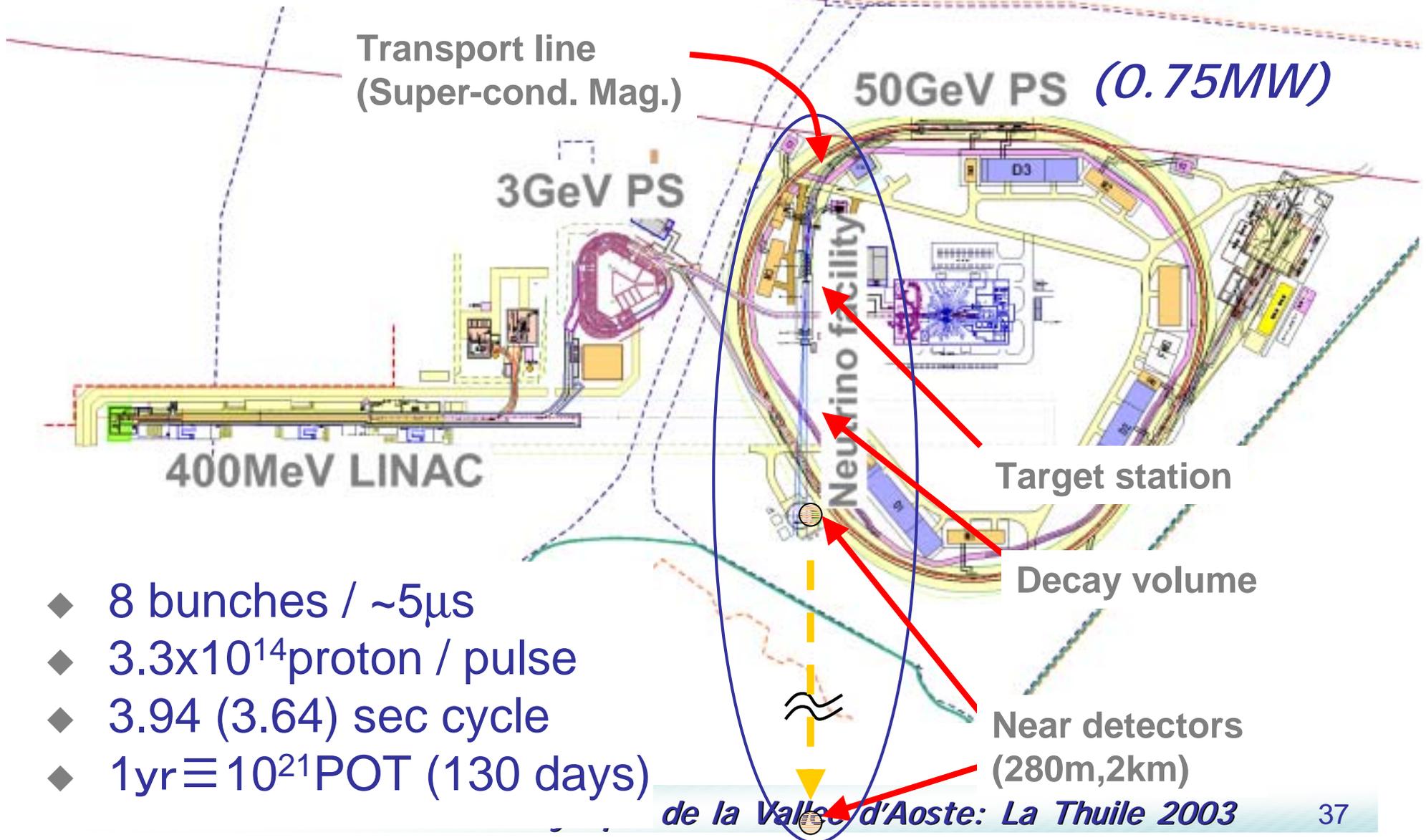
▶  $\times \sim 100$  of K2K

▶  $\nu_{\mu} \rightarrow \nu_x$  disapp. /  $\nu_{\mu} \rightarrow \nu_e$  app. / NC

# JHF Facility

Const.: 2001 ~ 2006 (approved)

(**v beam-line**: budget request submitted)



- ◆ 8 bunches /  $\sim 5\mu\text{s}$
- ◆  $3.3 \times 10^{14}$  proton / pulse
- ◆ 3.94 (3.64) sec cycle
- ◆  $1\text{yr} \equiv 10^{21}$  POT (130 days)

2002

2004

2006

2008

2010

 ■ **K2KII**

JHF- $\nu$  construction

physics run (OAB)

SK  
rebuild

SK-half

SK  
rebuild

SK-full