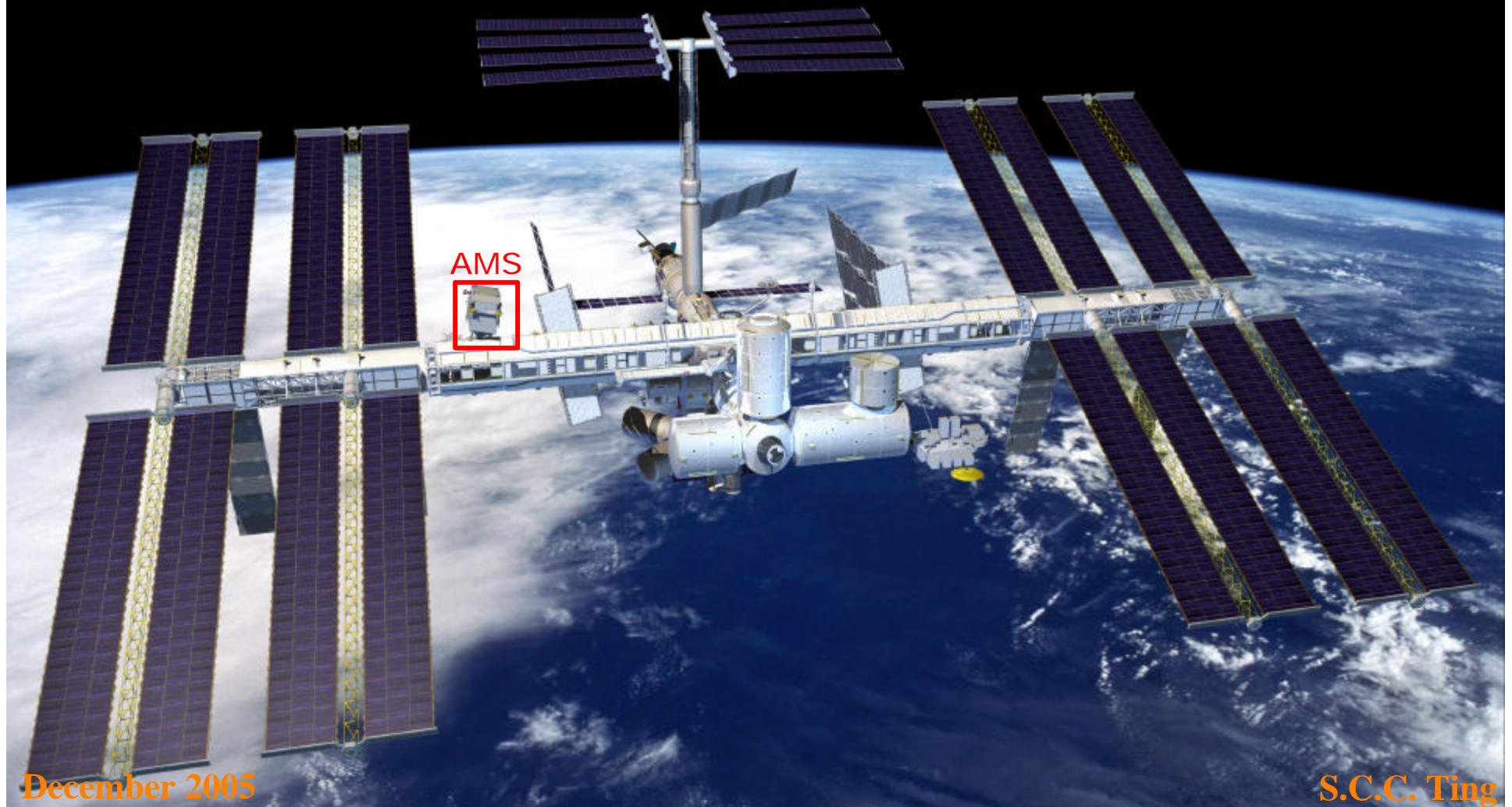


The Alpha Magnetic Spectrometer Experiment (AMS)



December 2005

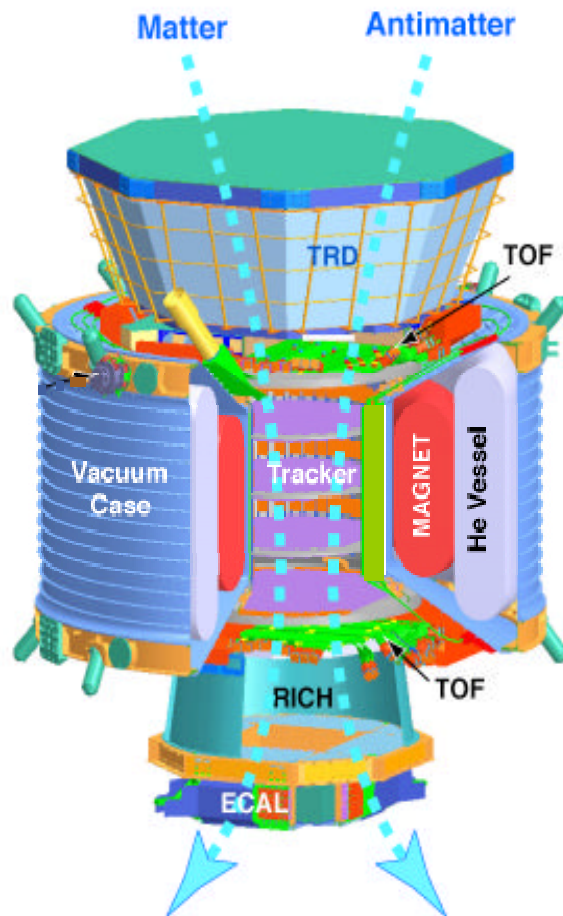
S.C.C. Ting

AMS: A TeV (10^{12} eV) Magnetic Spectrometer in Space

3m x 3m x 3m, 7t, $\sim 0.5\text{m}^2$ sr

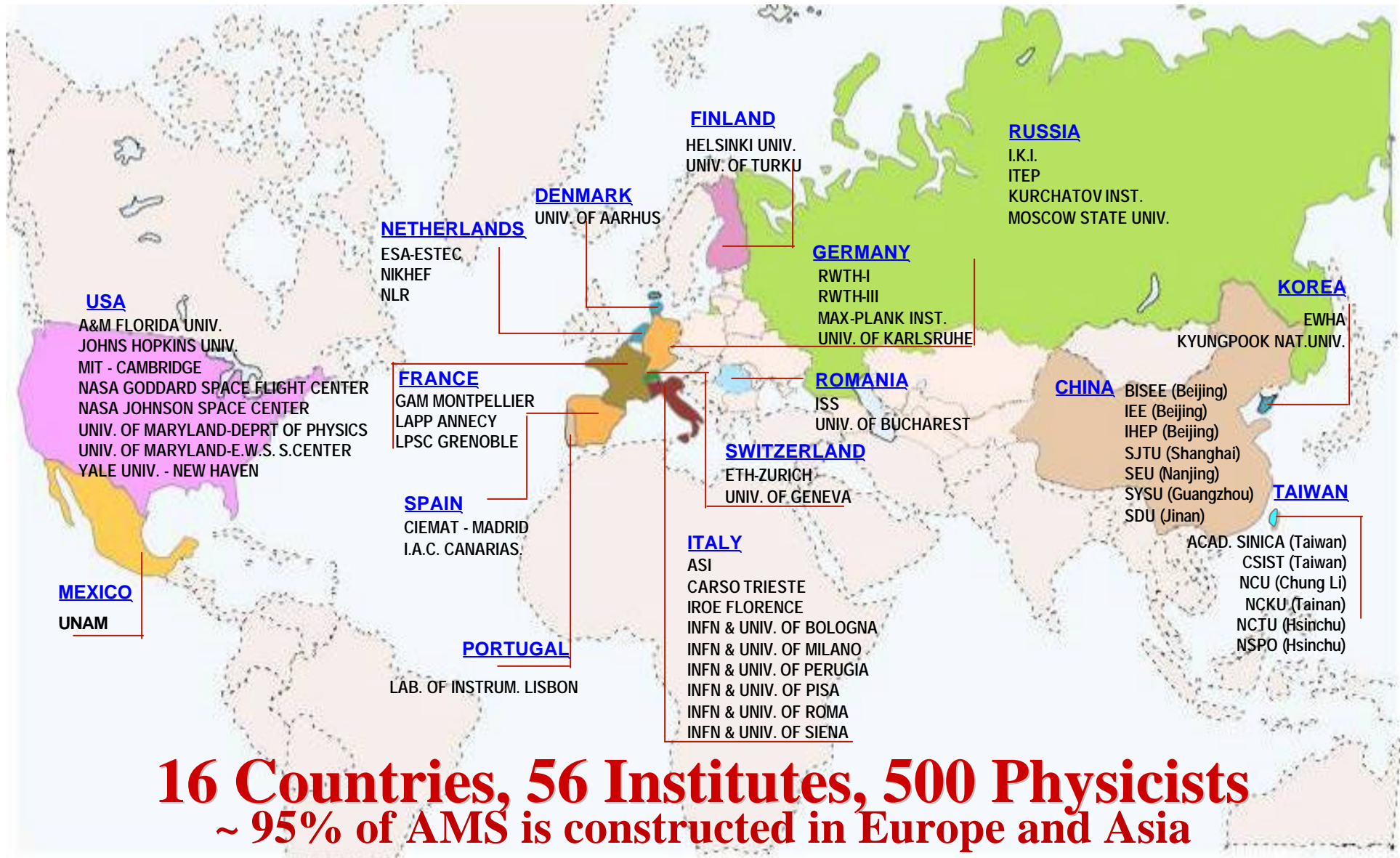
The AMS detector has been under construction for 10 years.

300,000 channels of electronics $\Delta t = 100$ ps, $\Delta x = 10\mu\text{m}$



0.3 TeV	e^-	P	He	C	Fe	γ
TRD						
TOF						
Tracker (magnet on)						
RICH						
ECAL						

International Participation in AMS

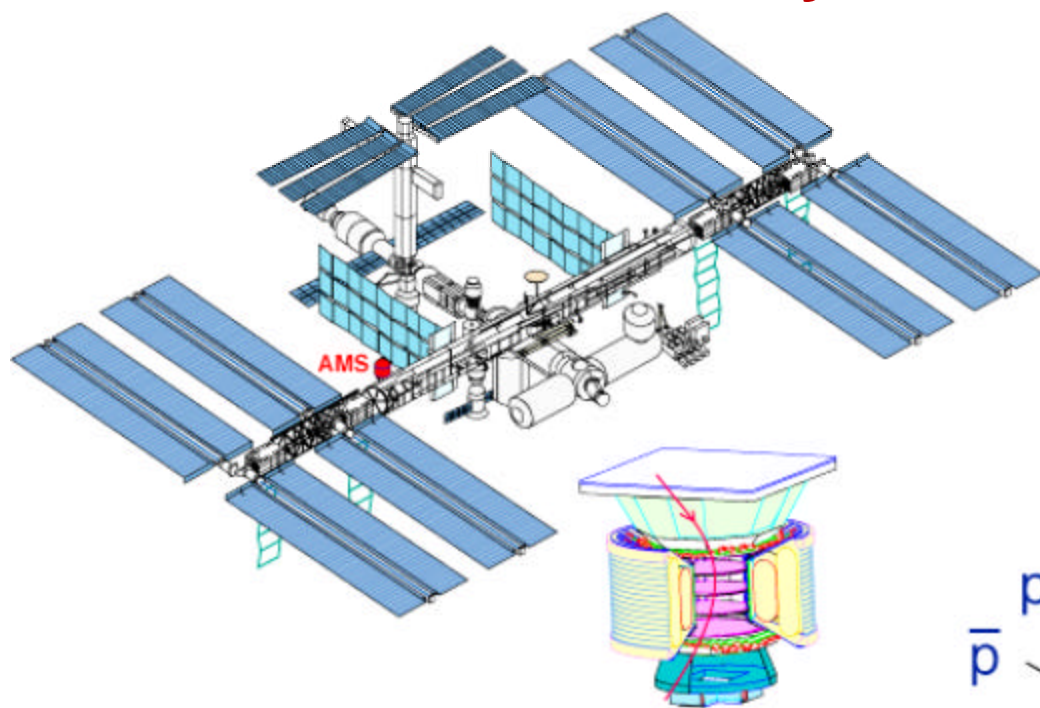


Italian National participation in AMS-02



- Universities
- Space Industry

Physics of AMS



Cosmic rays

1- Neutral component:

Light rays and neutrinos

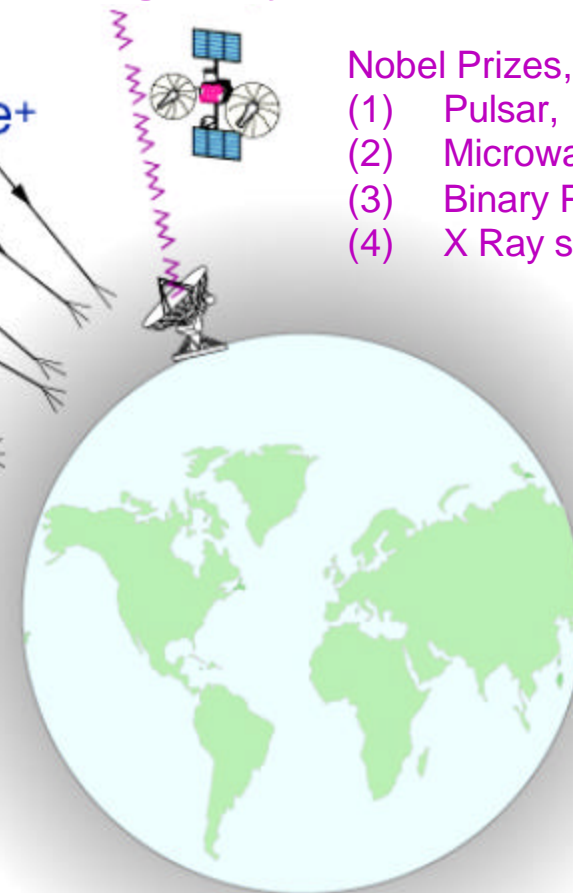
Nobel Prizes,
(1) Pulsar,
(2) Microwave,
(3) Binary Pulsars,
(4) X Ray sources

2- Charged component:

He, Be, C, Fe

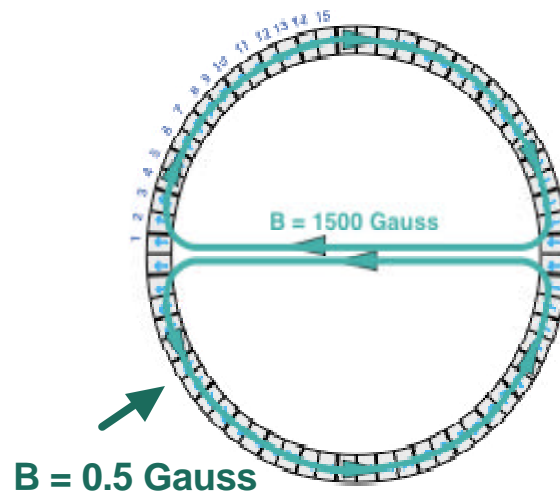
$\bar{\text{He}}$,

AMS will perform accurate measurements of
energetic charged cosmic rays (0.2 GeV/n to 2 TeV/n)
and high energy gamma rays (0.5 GeV to 300 GeV)



**There has never been a large superconducting magnet in space
due to the extremely difficult technical challenges**

AMS-01

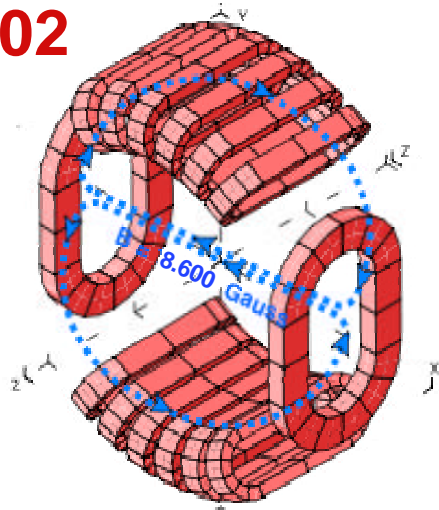


STEP ONE:

A Permanent Magnet to fly on STS-91

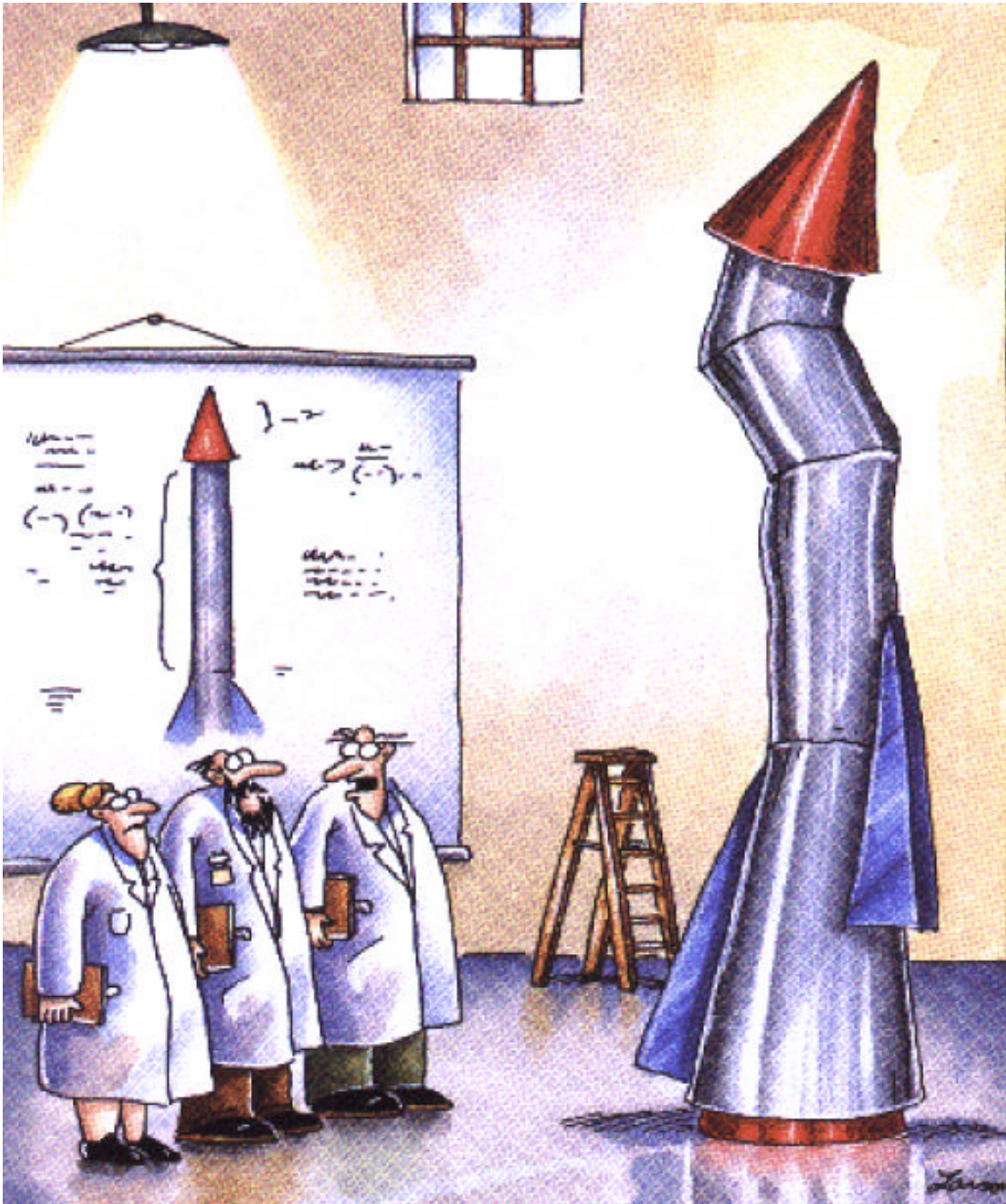
- 1- Stable: no influence from earth magnetic field**
- 2- Safety for the astronauts:
Minimum field leakage out of the magnet**
- 3- Low weight: no iron**

AMS-02



STEP TWO:

**A Superconducting Magnet for ISS
with the same field arrangement**



*“Its time we face reality,
my friends....*

*We’re not exactly
rocket scientists.”*

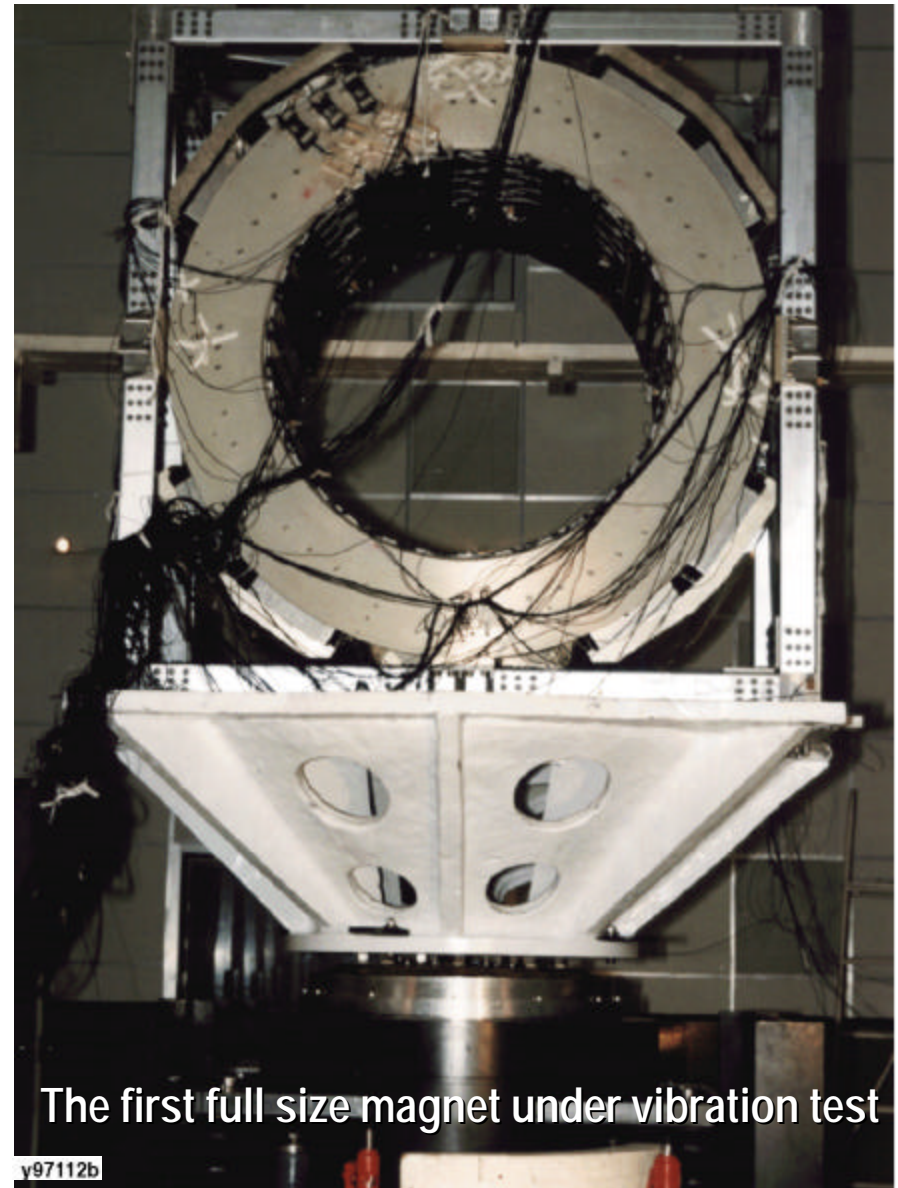
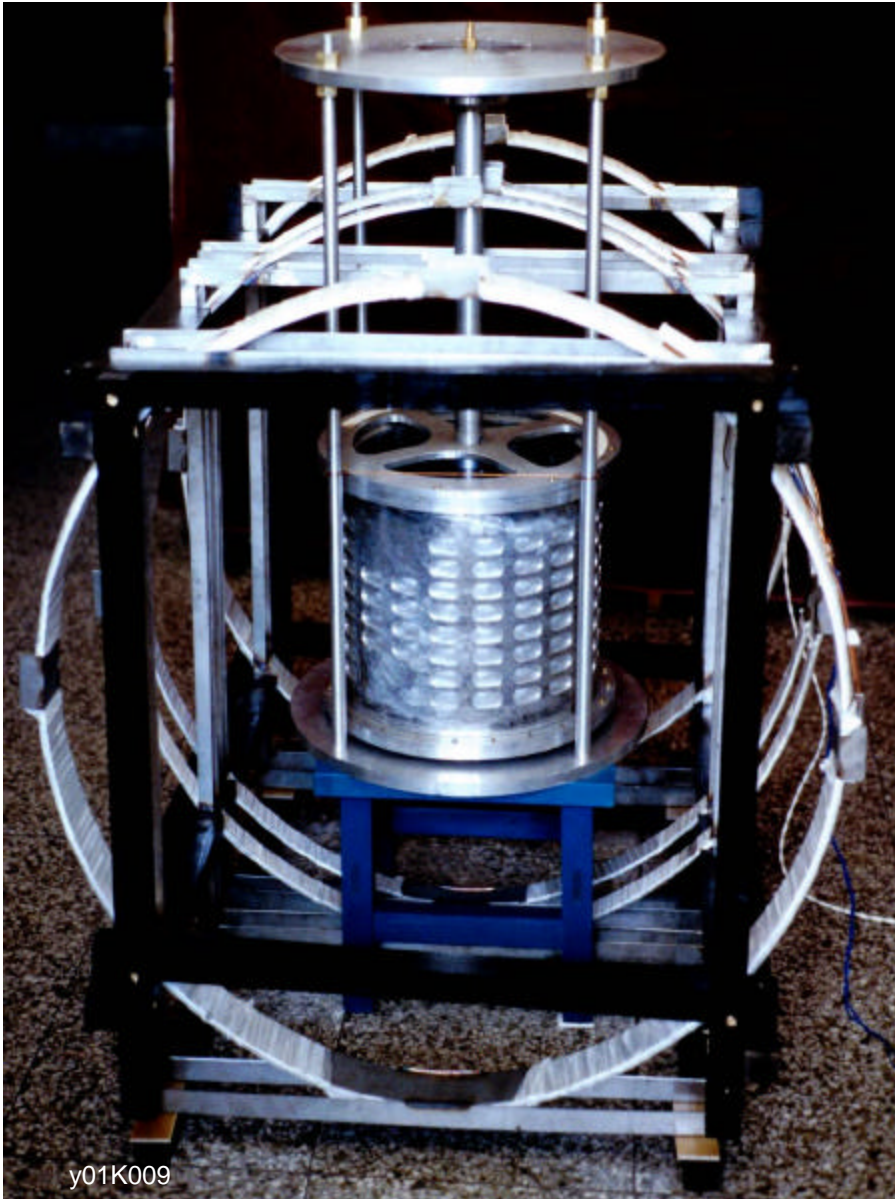
AMS-01: STS-91

10 Magnets were made

+

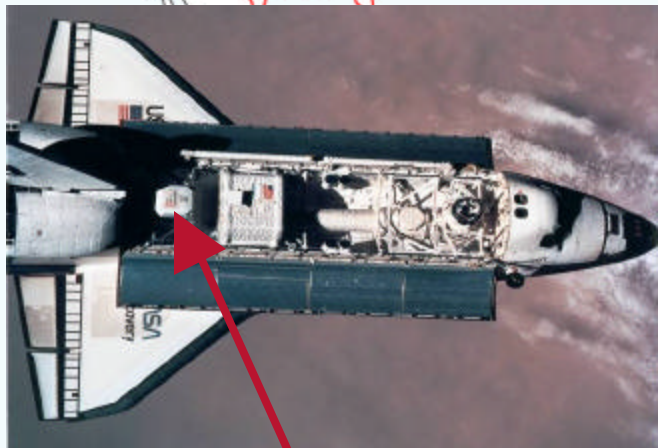
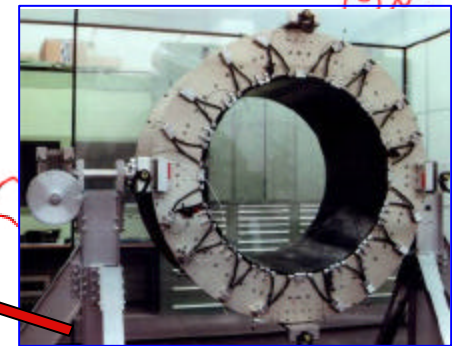
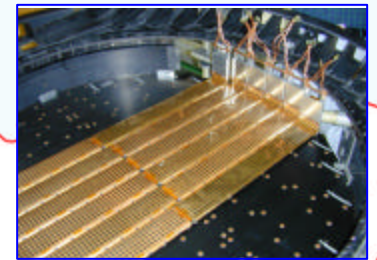
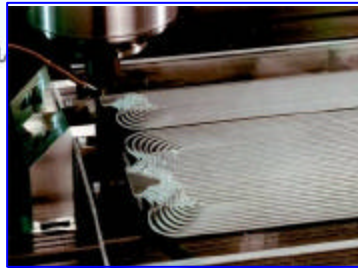
3 (full size) for Space Qualification,
test to Destruction and Flight.

7 to study Field Calculation,
Field Leakage, Dipole moment

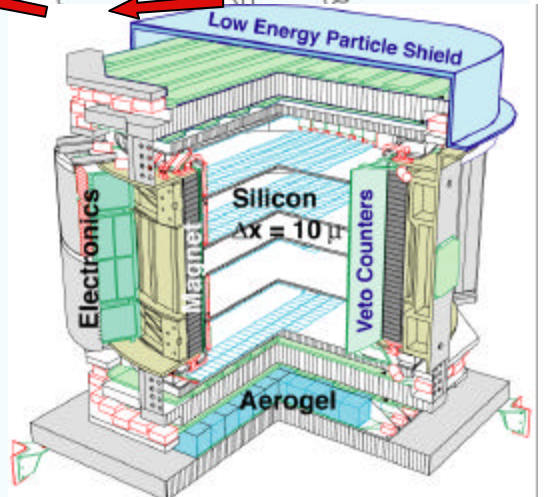


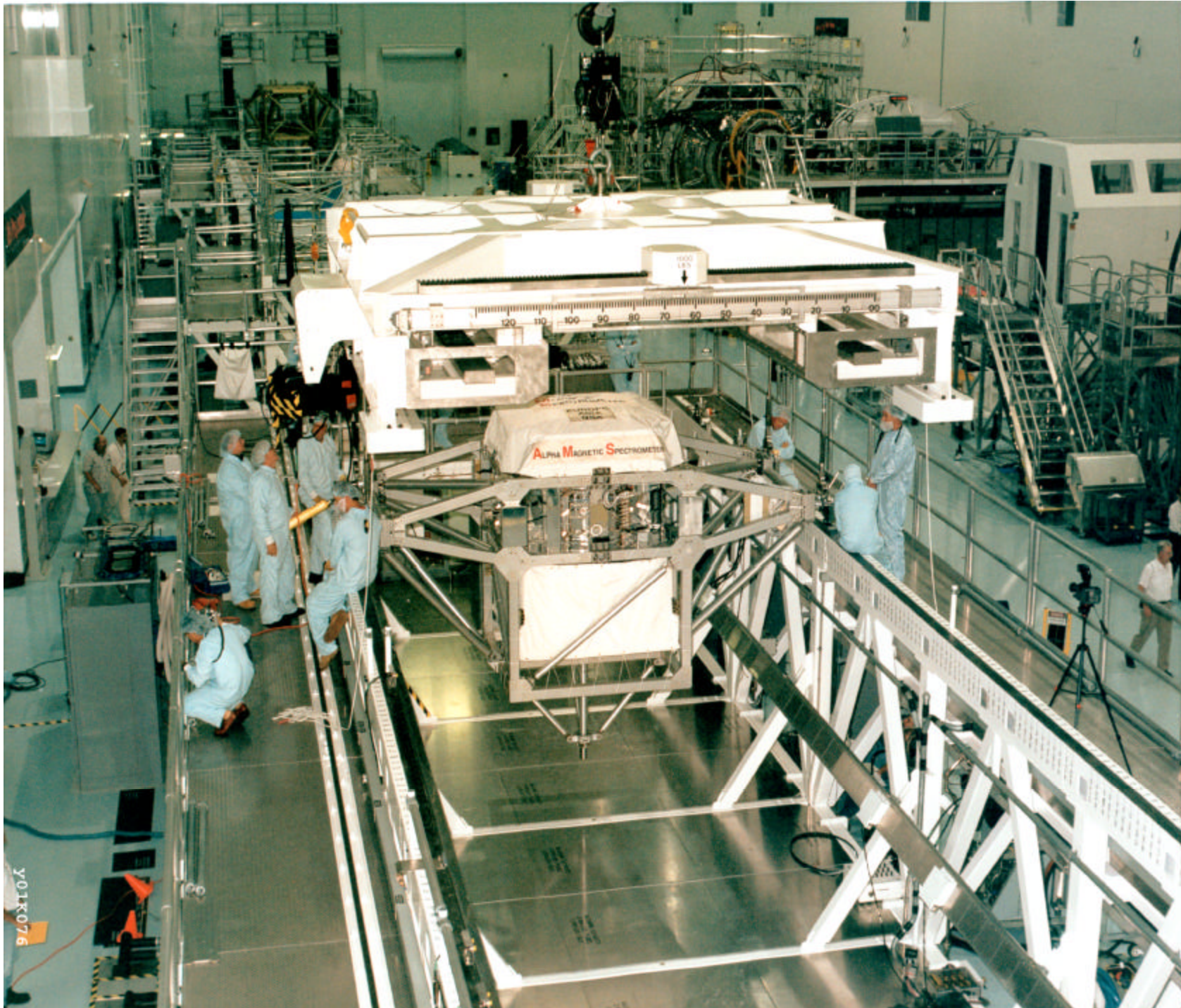
First flight AMS-01

Approval: April 1995, Assembly: December 1997, Flight: June 1998

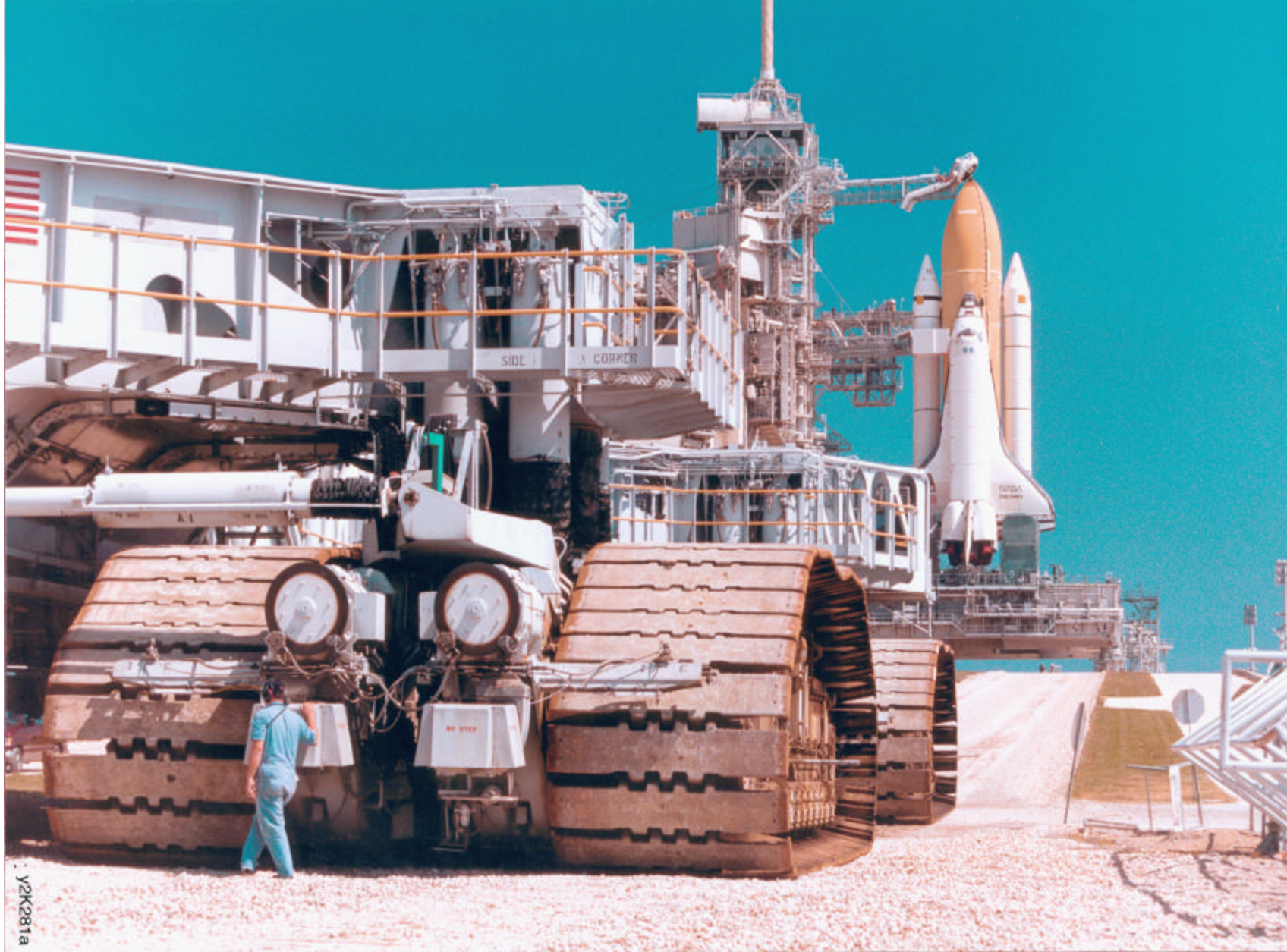


AMS

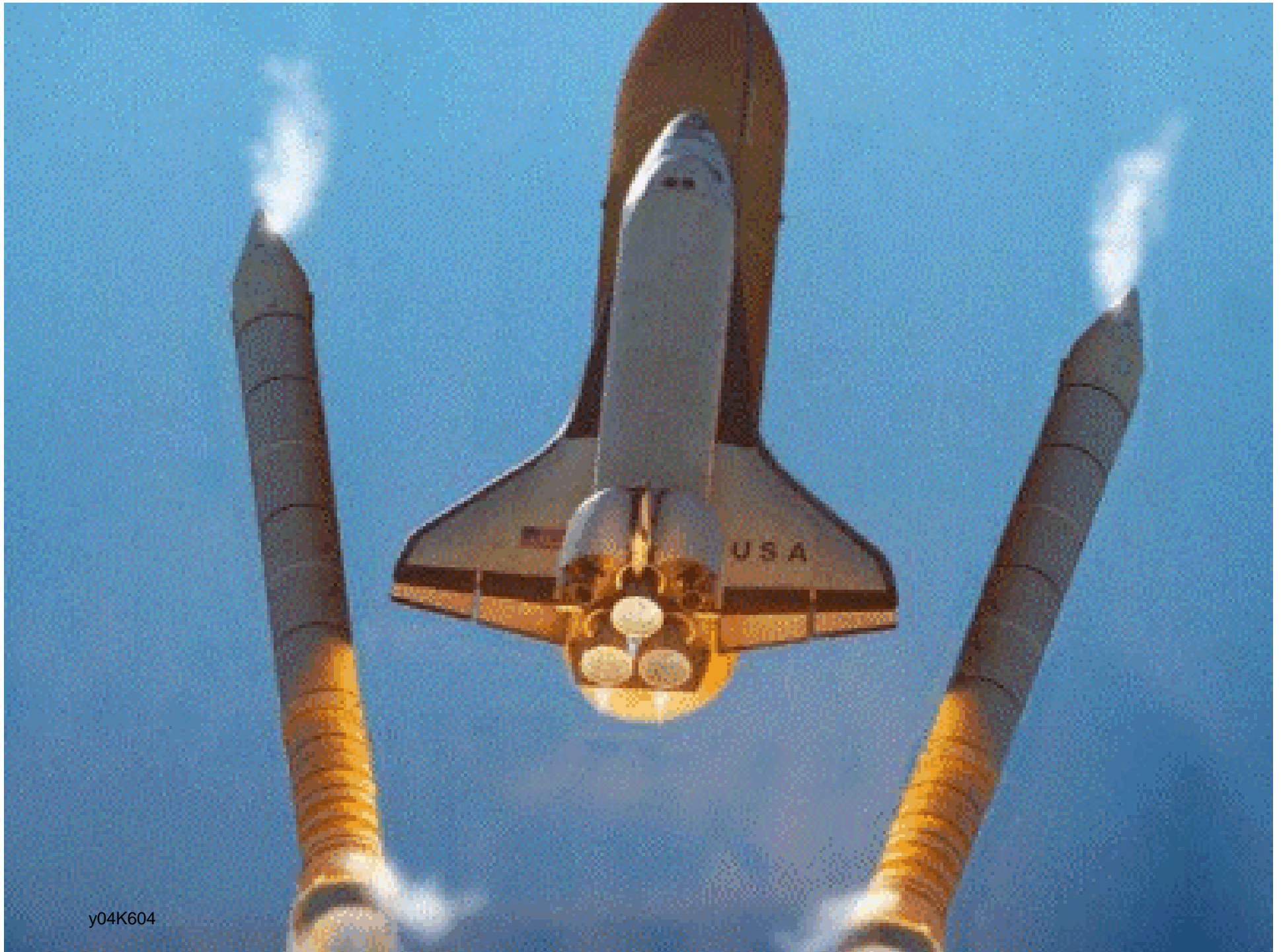












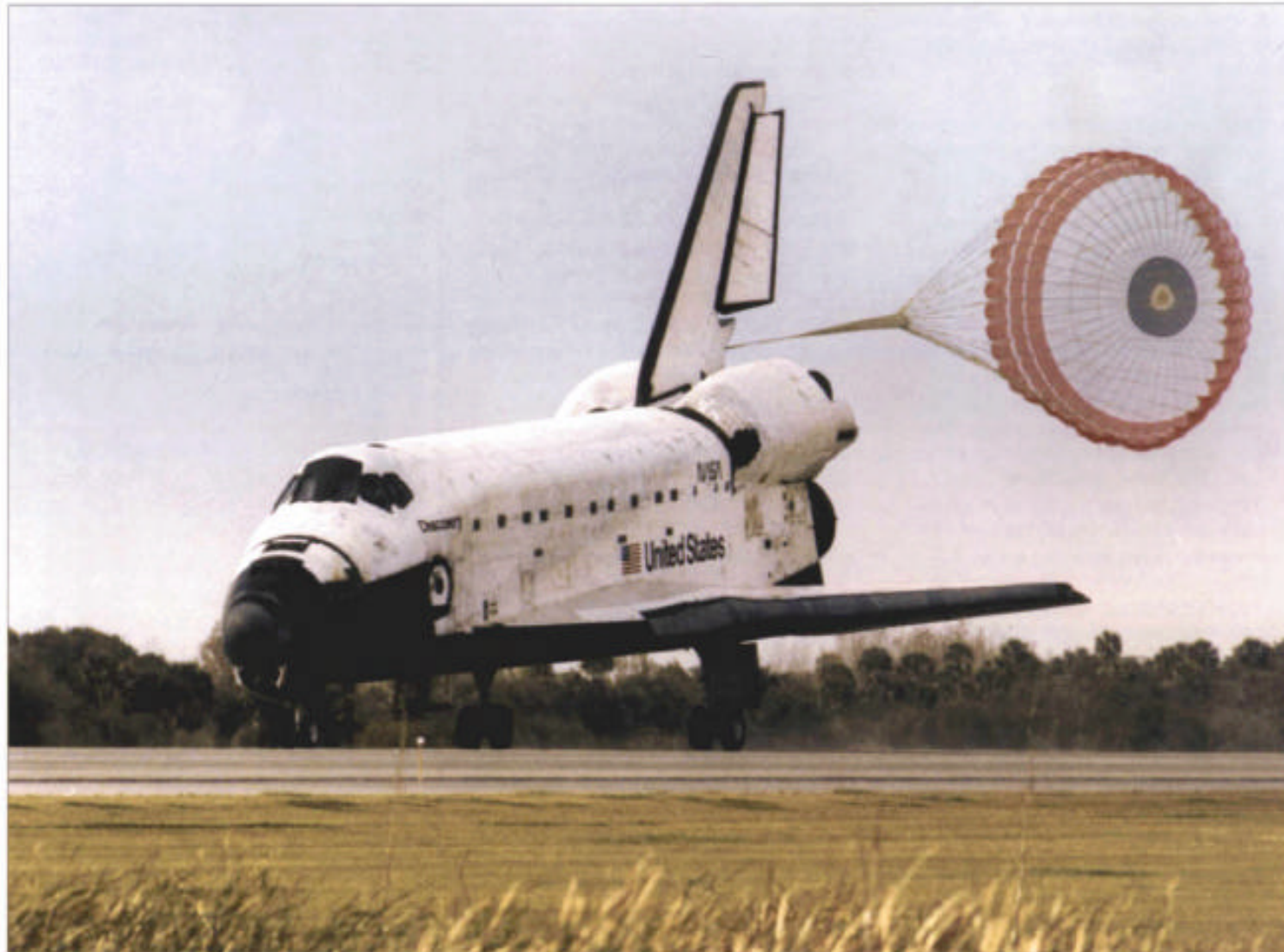
y04K604





National Aeronautics and
Space Administration

Space Shuttle *Discovery* Returns from Space

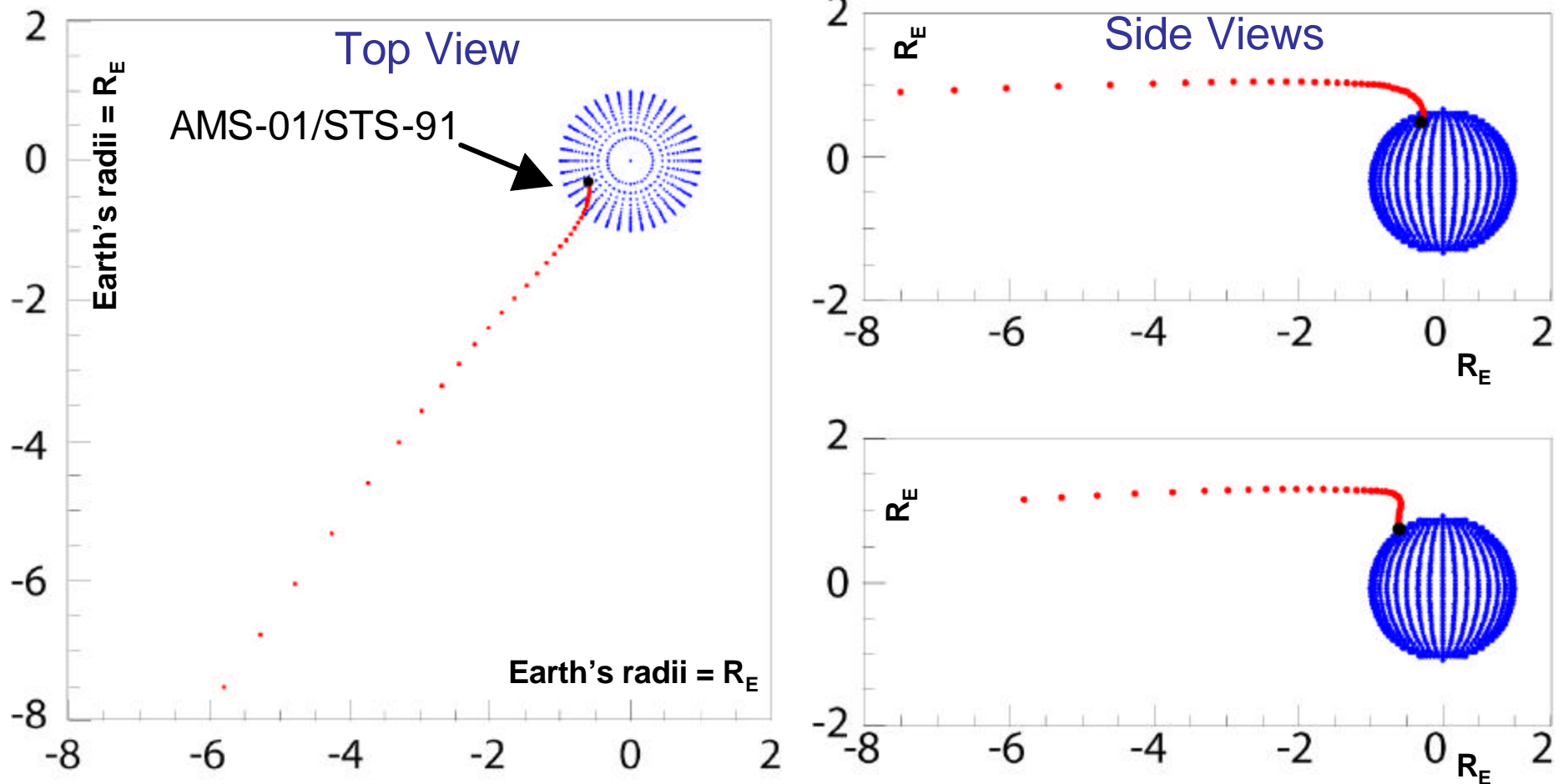


y97271c

AMS-01 Mission returns to K.S.C. - June 12 3 pm, 1998

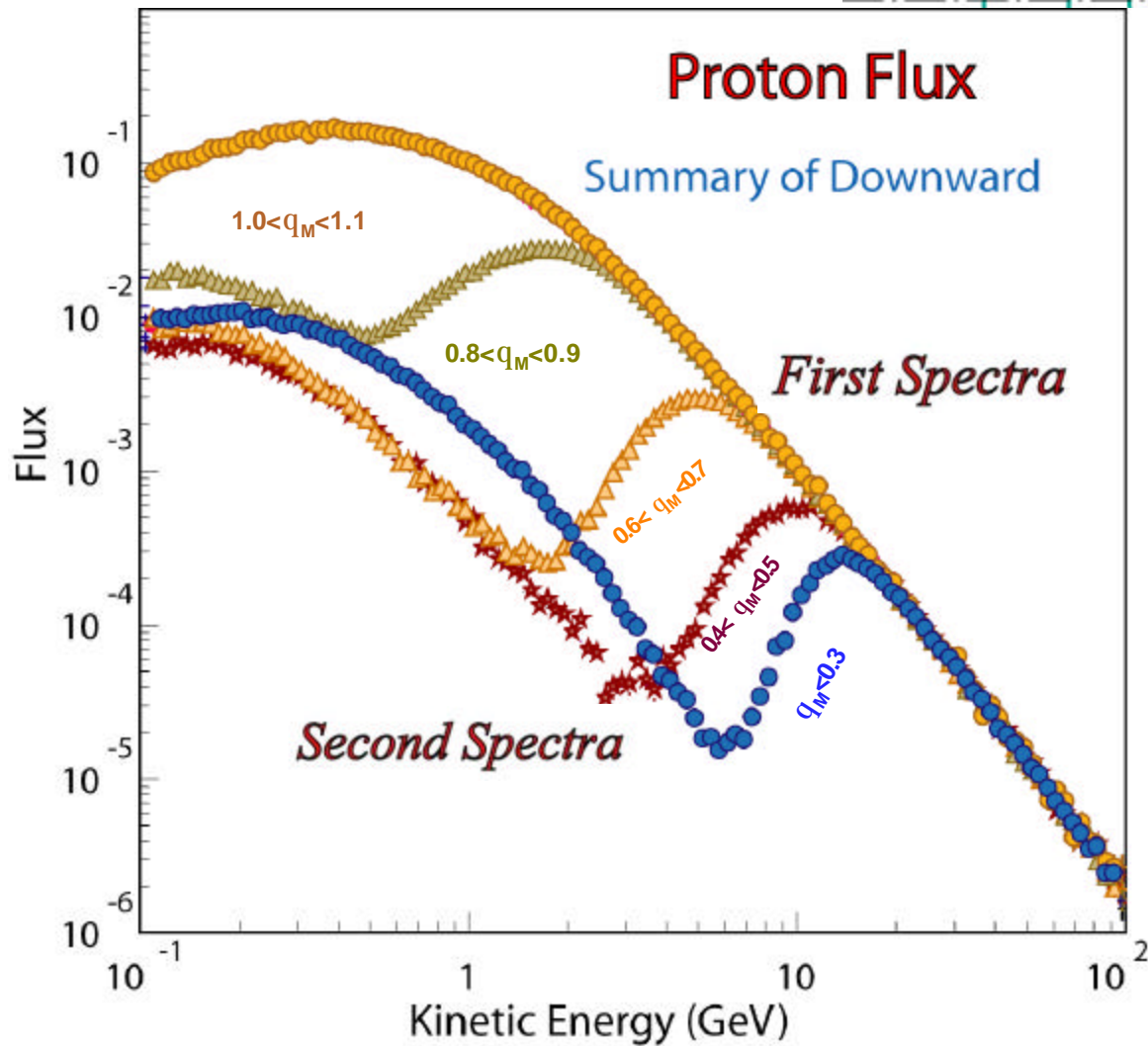
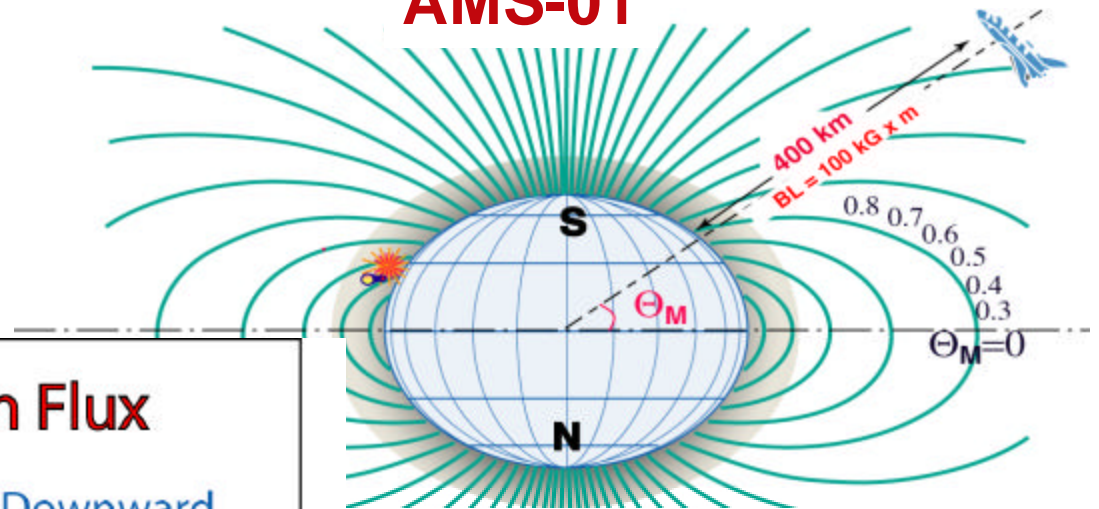
Tracing the origin of cosmic rays from precision measurements of incident momentum vector and location

Run/Event: 897171331/310063 06-Jun-98 22:33:48 Lat/Long: 50N/27W



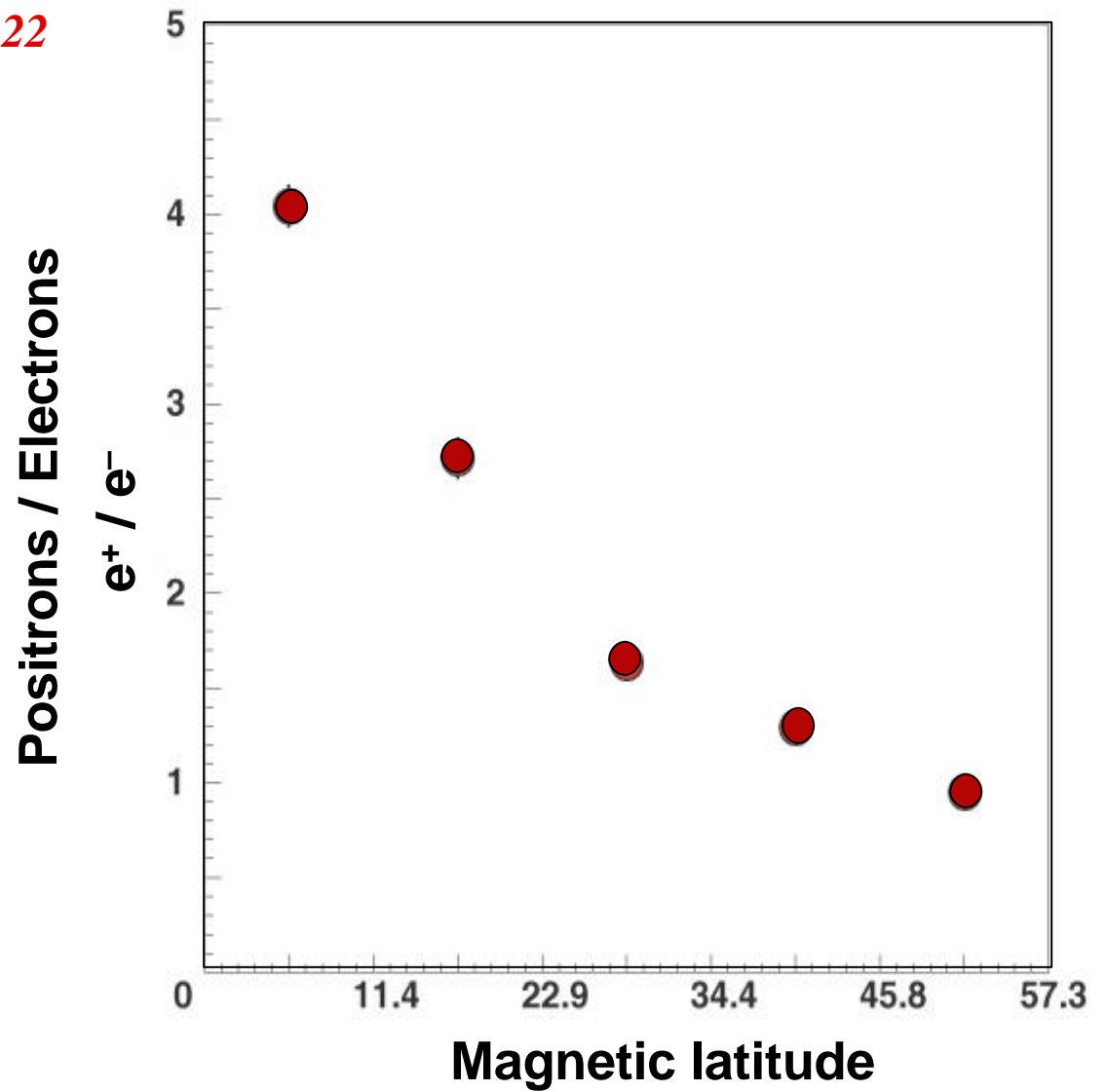
Primary proton, $E_K = 516$ MeV,
Measured at $R_E = 1.06$, traced back to $R_E > 8$

AMS-01



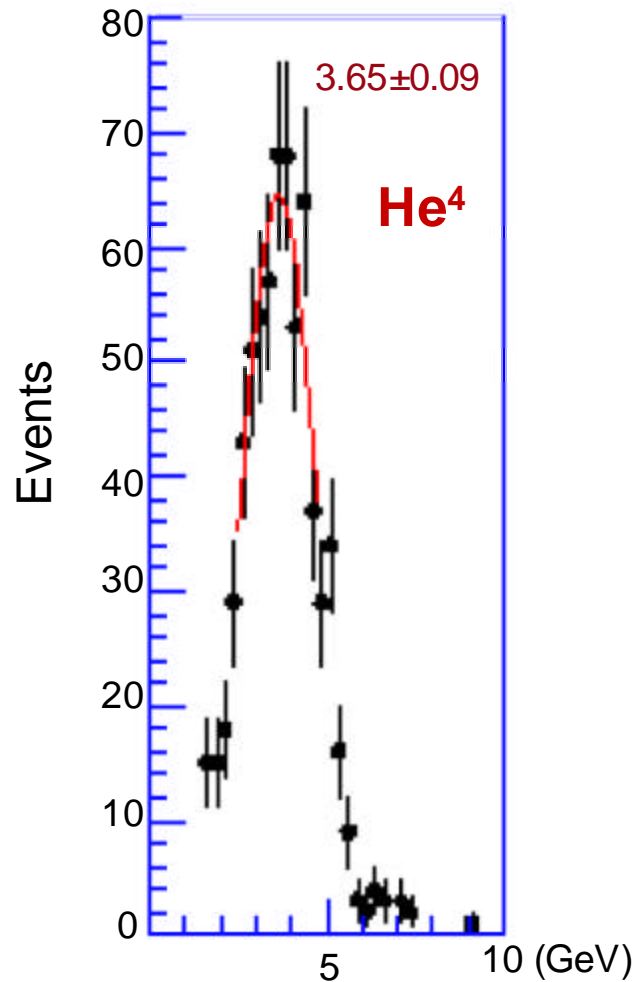
AMS paper: “Leptons in Near Earth Orbit”

Physics Letters B 484 (2000) p. 10-22

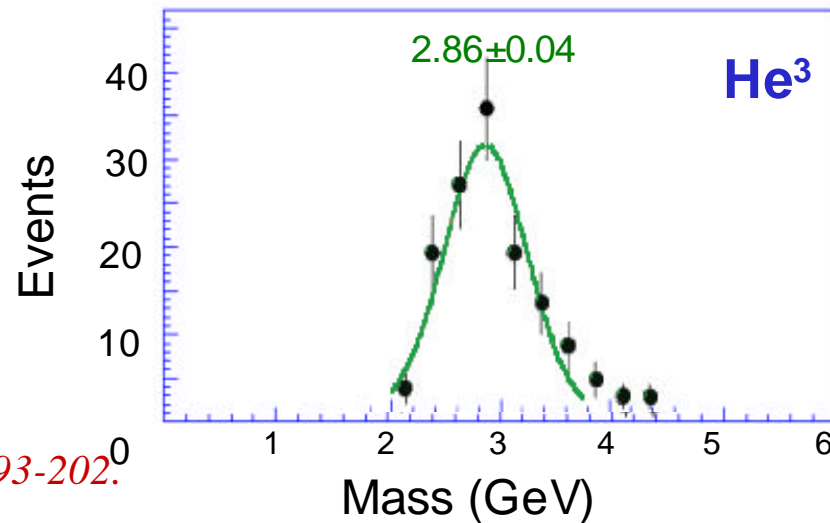
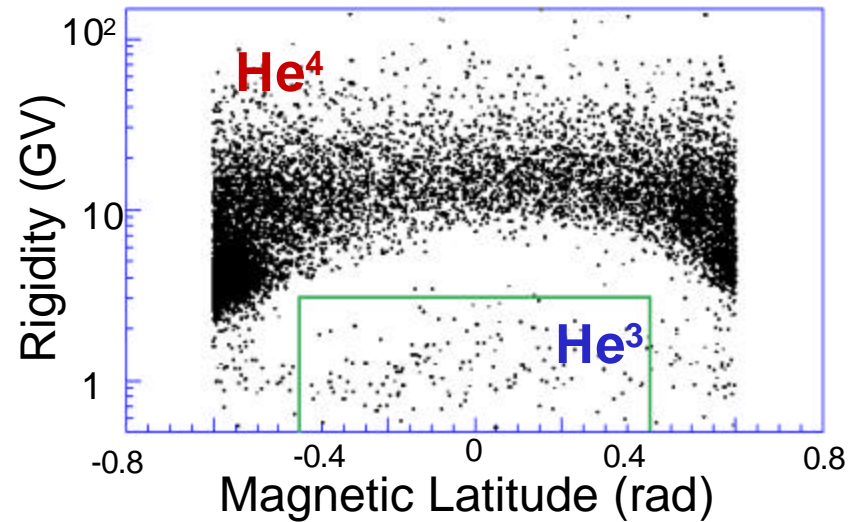


“Helium in Near Earth Orbit”

Helium spectrum

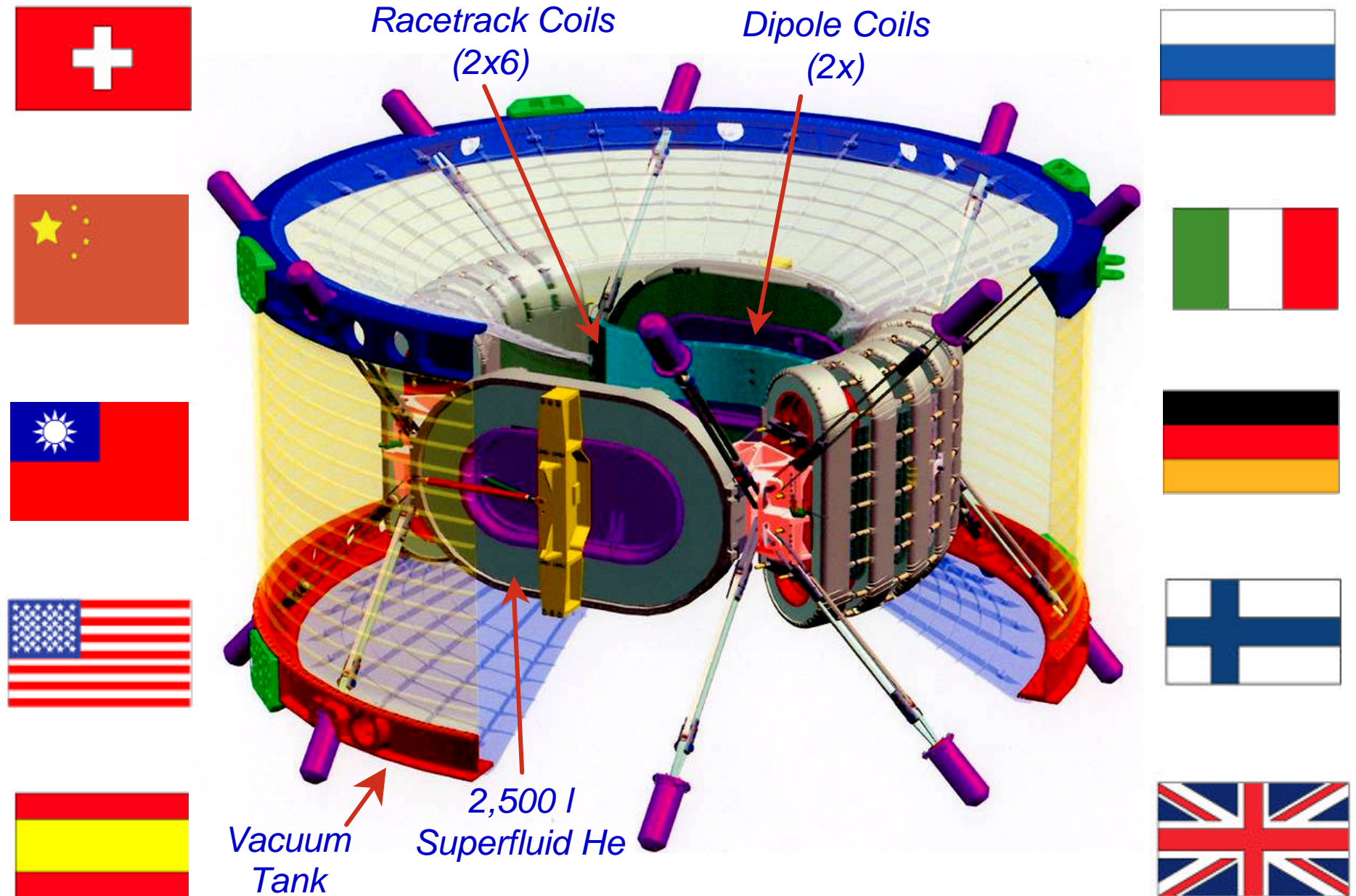


(Mass of $\text{He}^4 = 3.7 \text{ GeV}$; $\text{He}^3 = 2.8 \text{ GeV}$)



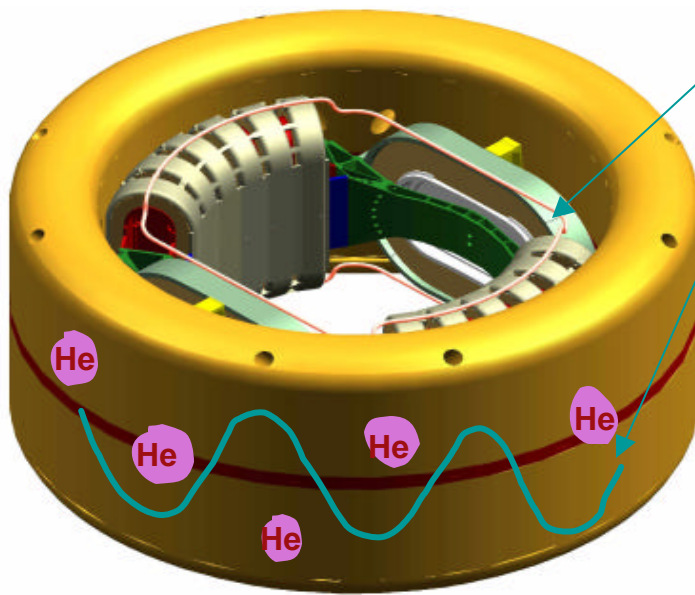
Physics Letters B vol.494 (3-4) (2000) p.193-202.

Construction of the AMS-02 Superconducting magnet



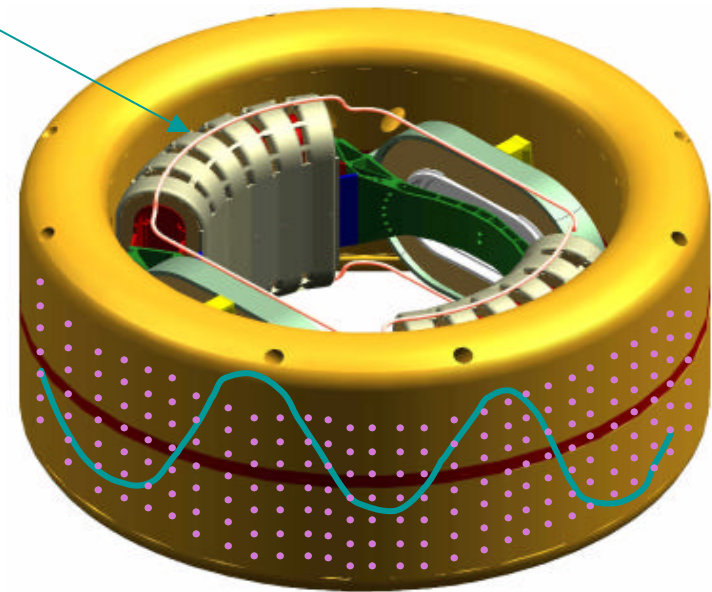
**For a magnet with long duration without refill and light weight,
use superfluid Helium**

Indirect cooling with cold heat exchanger



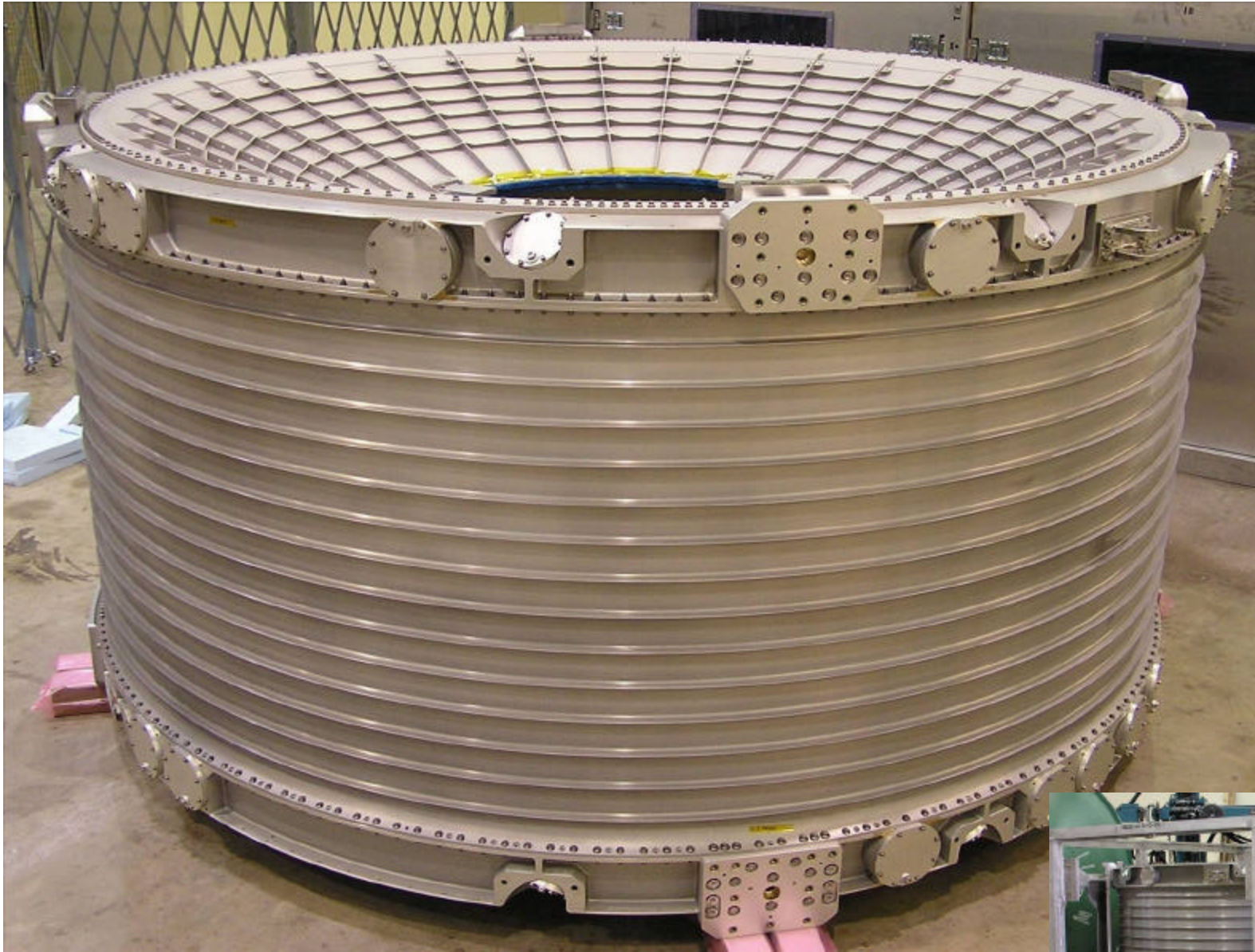
Normal liquid Helium:
4.2K

In Space: Cold Heat exchanger cannot
be uniformly cooled



Superfluid Helium:
1.8K
film flow

In Space: Cold Heat exchanger is
uniformly cooled

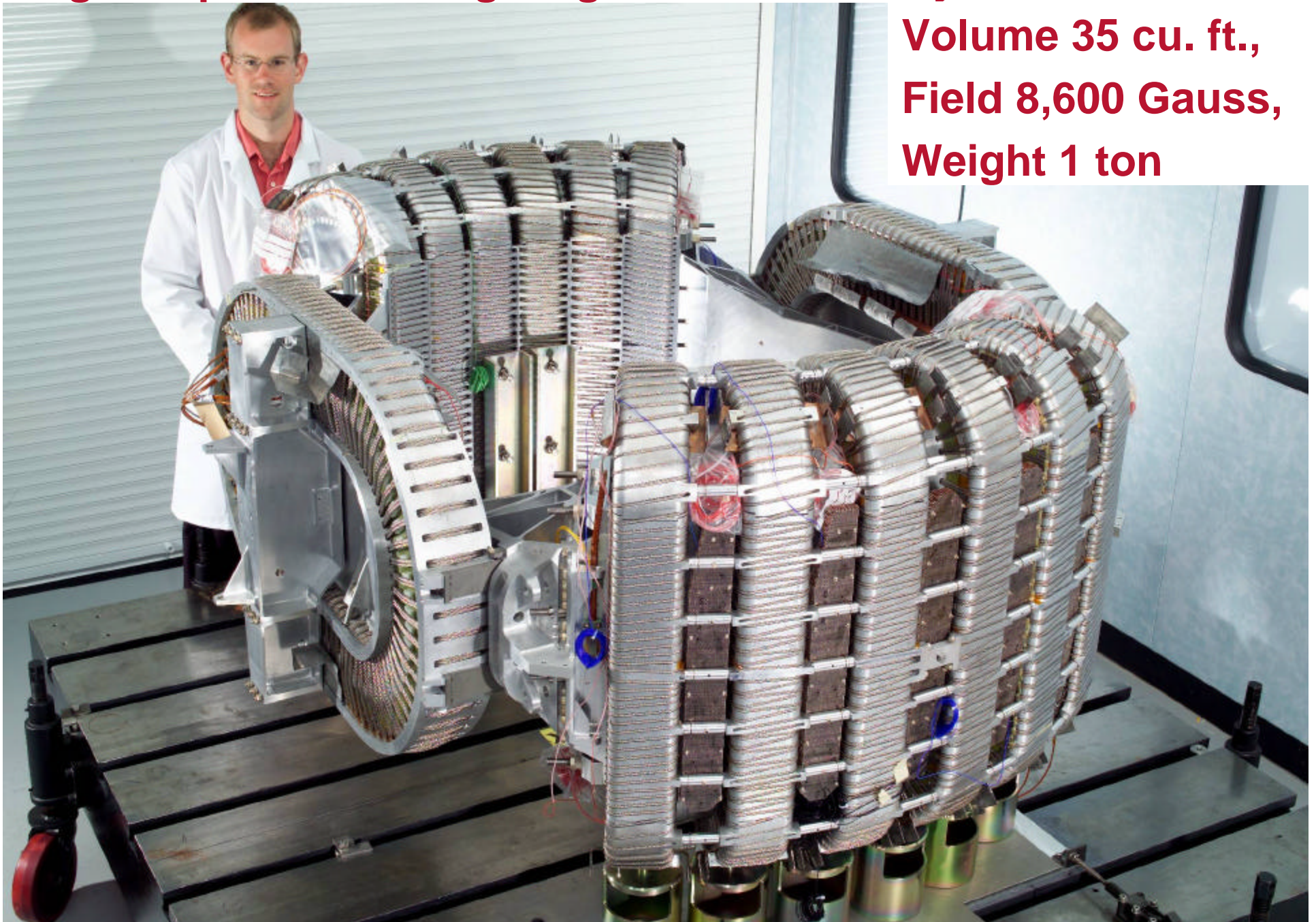


Flight Vacuum Case at JSC



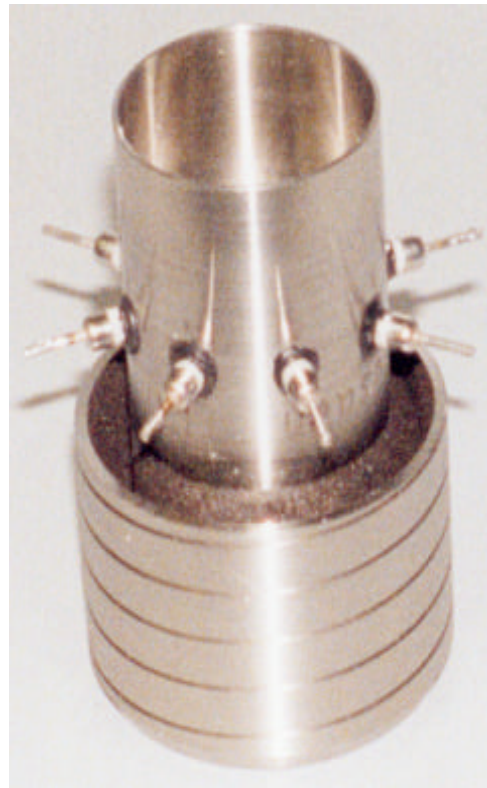
Flight superconducting magnet coils are fully assembled:

**Volume 35 cu. ft.,
Field 8,600 Gauss,
Weight 1 ton**



New Technology for superconducting magnet in space

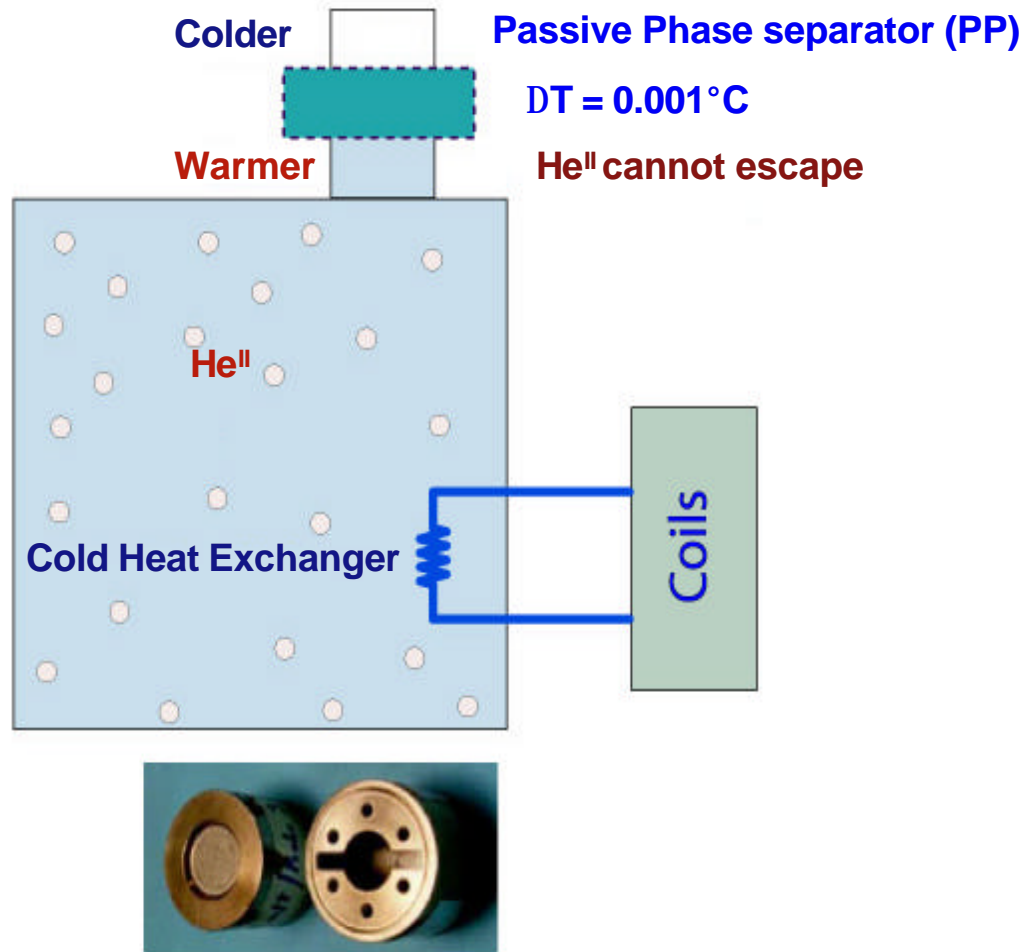
Thermal Mechanical Pump transfers superfluid helium at zero gravity
in strong magnetic field.



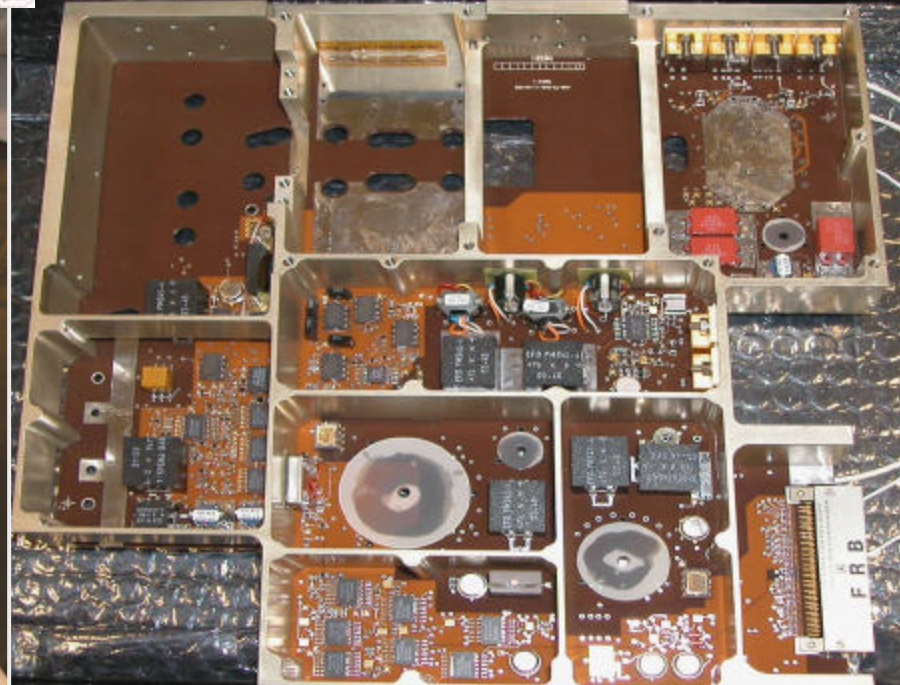
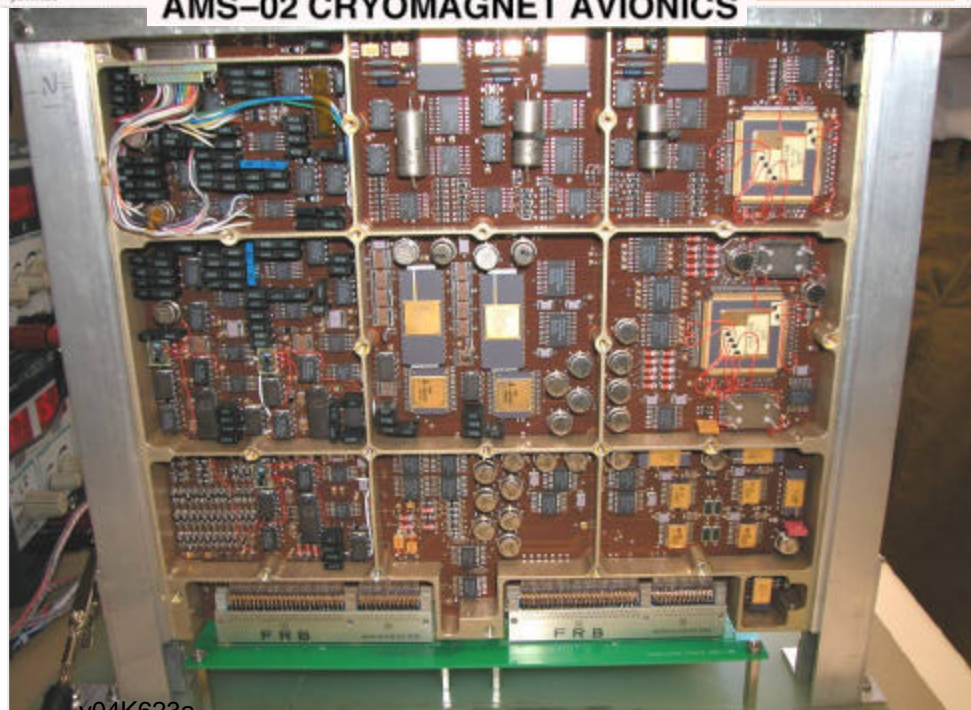
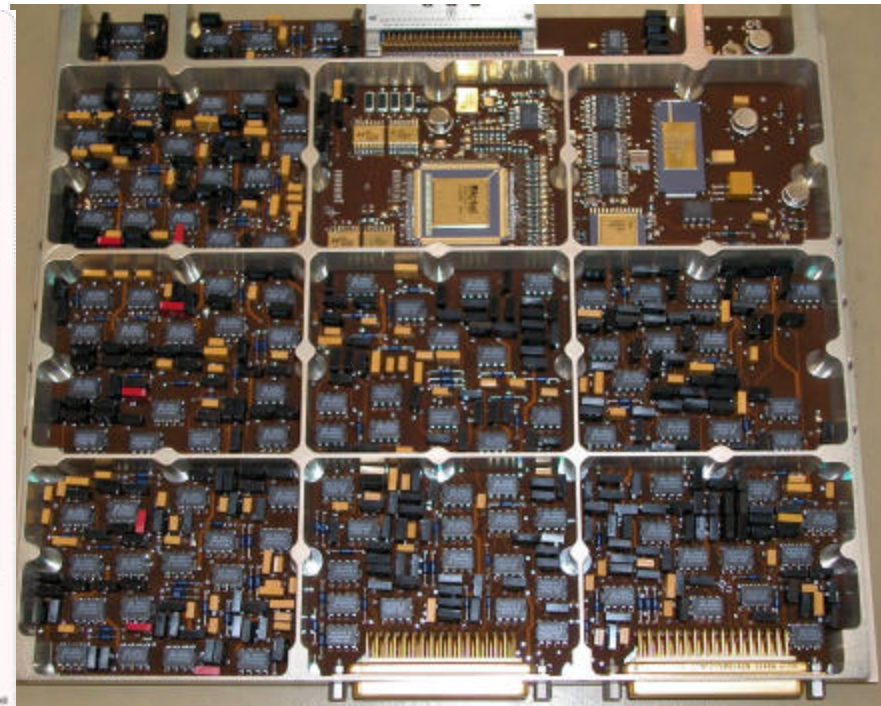
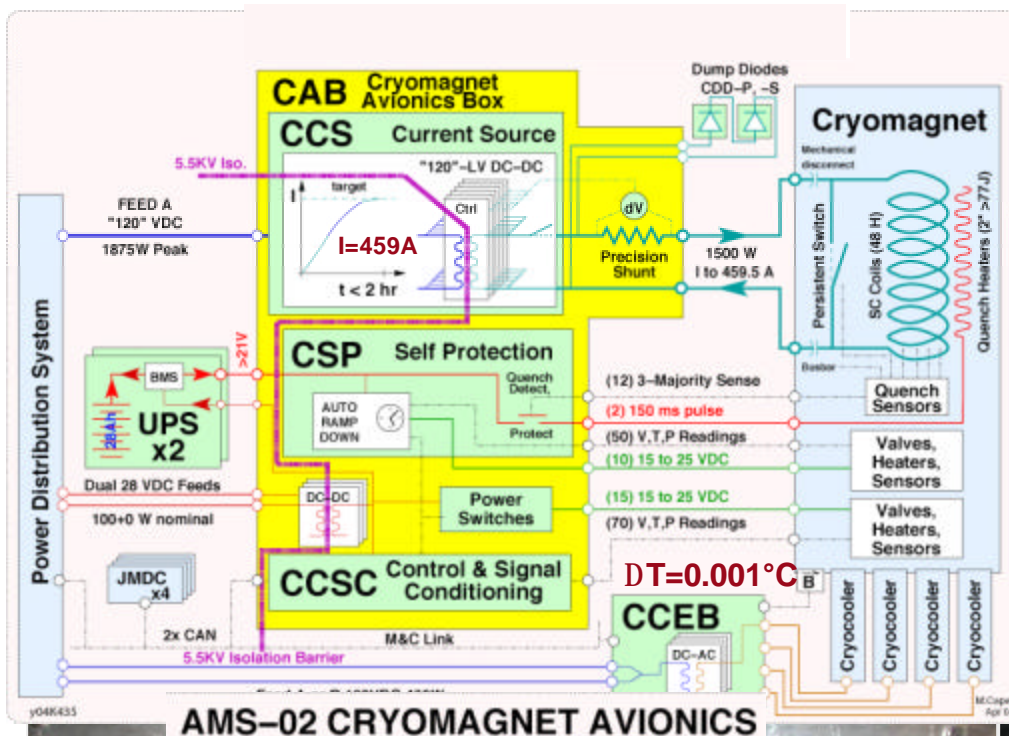
Heater $\xrightarrow[\text{colder}]{\text{warmer}}$ He² $\Delta = 0.01^\circ\text{C}$

New Technology for superconducting magnet in space

A passive phase separator in a strong magnetic field at 1.8K and zero gravity to remove helium gas from superfluid He tank



Flight Hardware

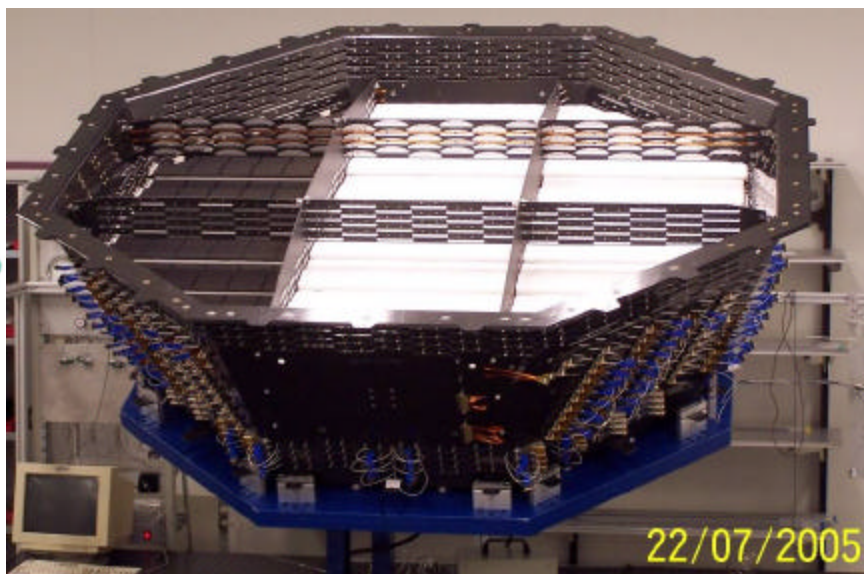
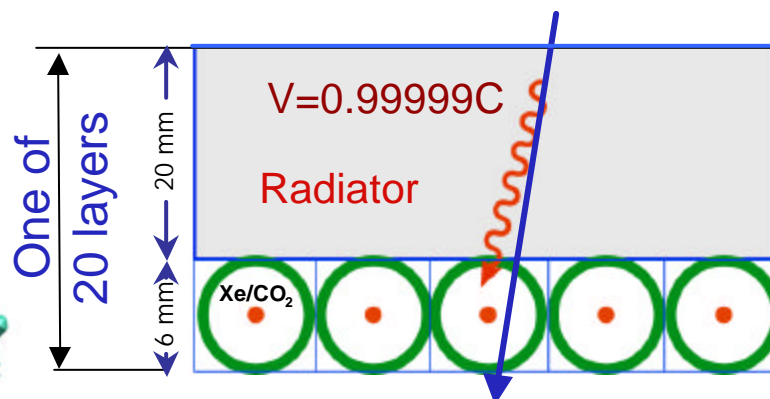
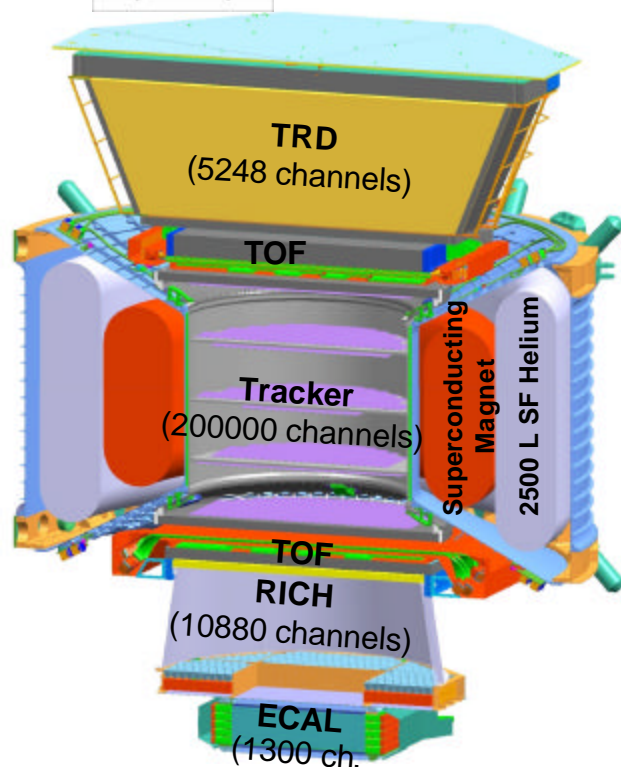


y04K623a



Transition Radiation Detector (TRD)

Distinguishes electrons and positrons from all other particles

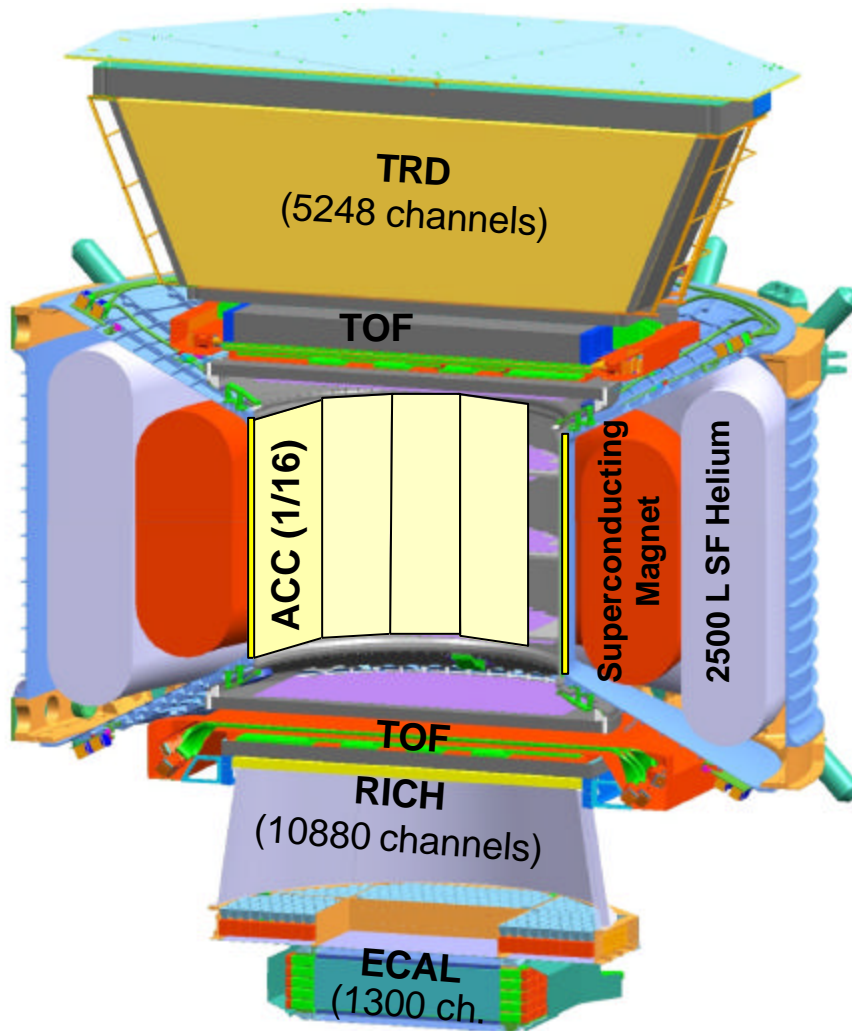


5248 tubes
 $L(\text{max}) = 1.8\text{m}$

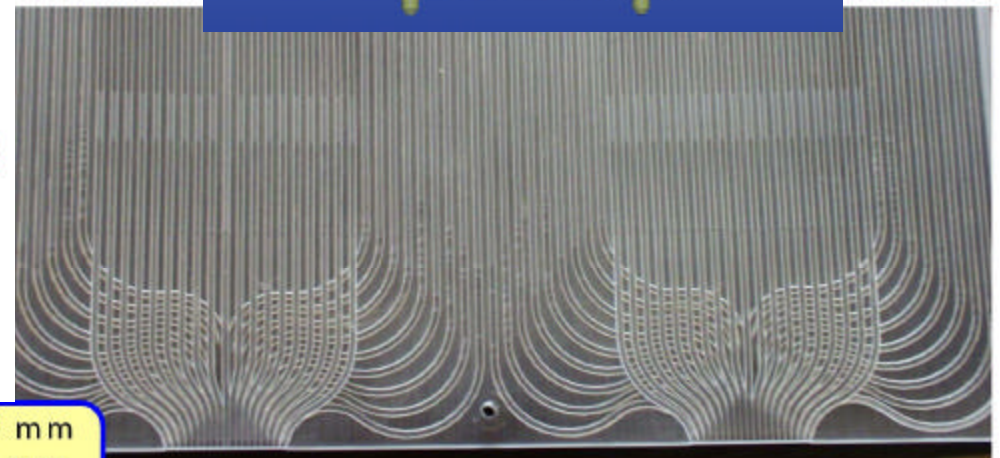
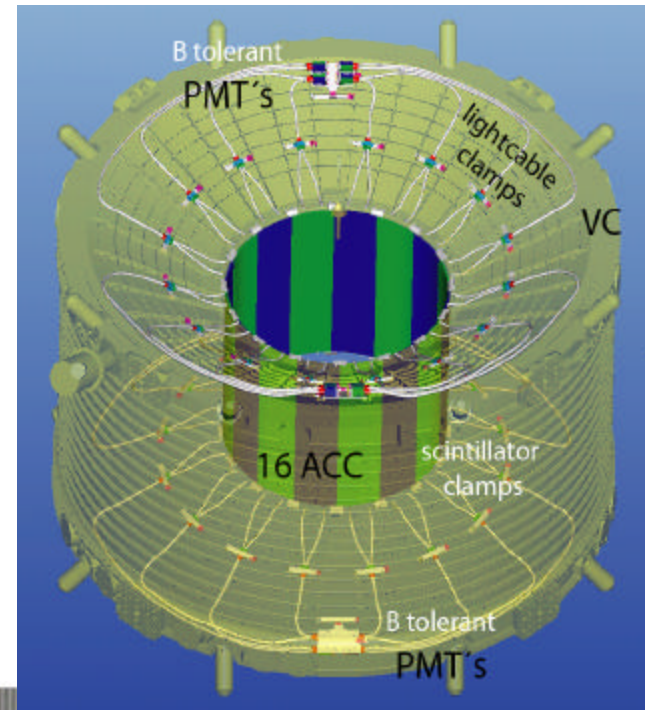
All Flight Hardware Modules produced. Assembly complete in 2005.

Anti-Coincidence Counters

Coordinator: W.Wallraff, RWTH-Aachen



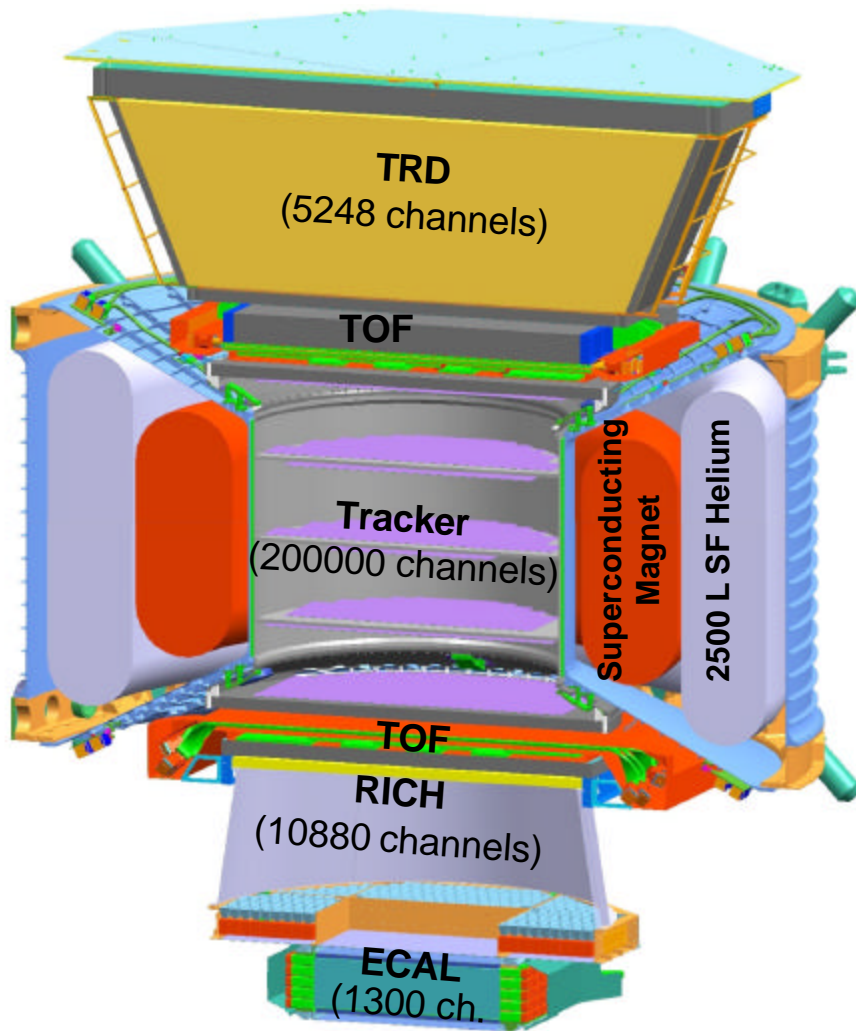
Thickness=8 mm
 Length=830 mm
 Width=220 mm
 Fiber pitch=2.9 mm
 Fiber diam.=1.0mm
 2×37 fibers



Precision milling of wave length fiber grooves
 to PMT

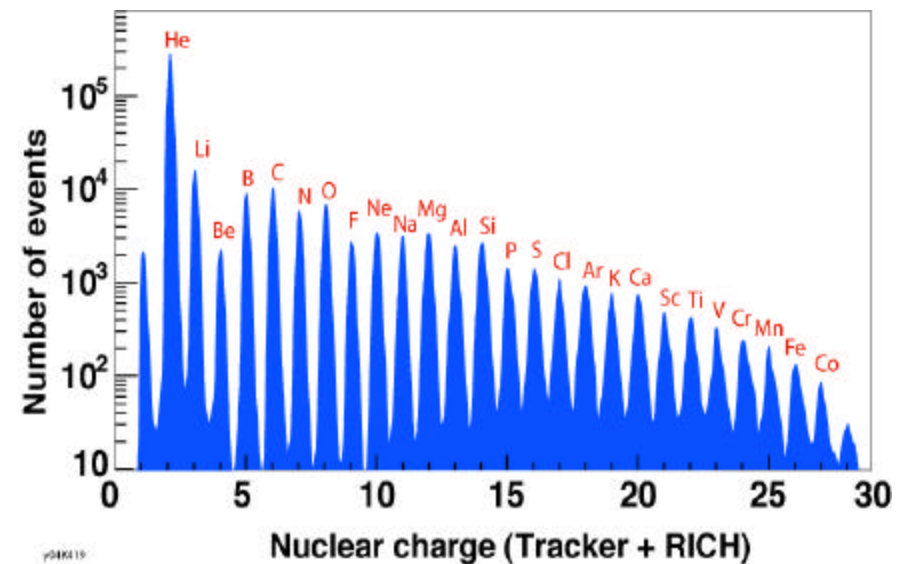
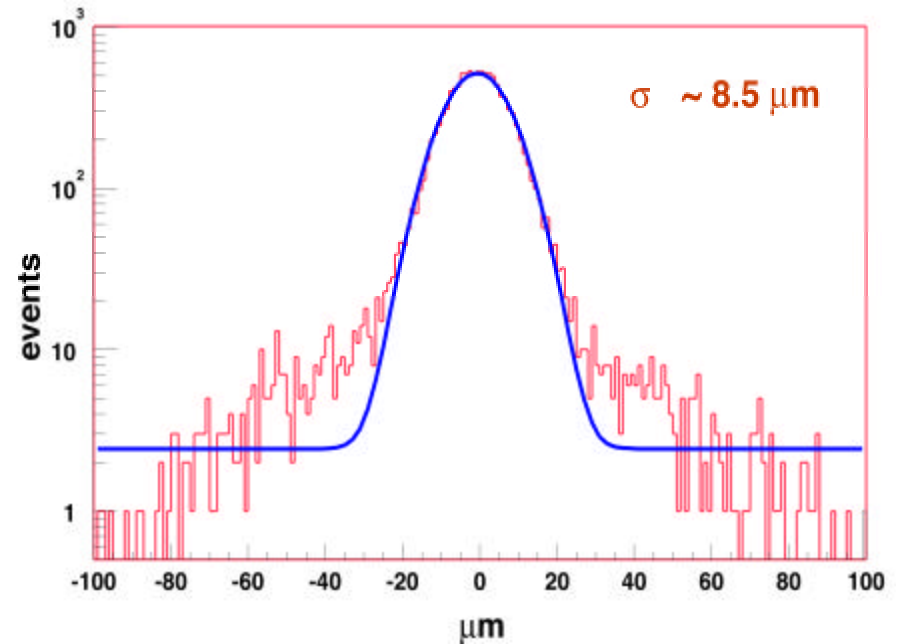
to PMT

Silicon Tracker



**8 planes 6.6 m²:
Largest Silicon Detector
with 10 mm accuracy**

Production completed. Test results from accelerator.



Silicon Tracker Flight Hardware

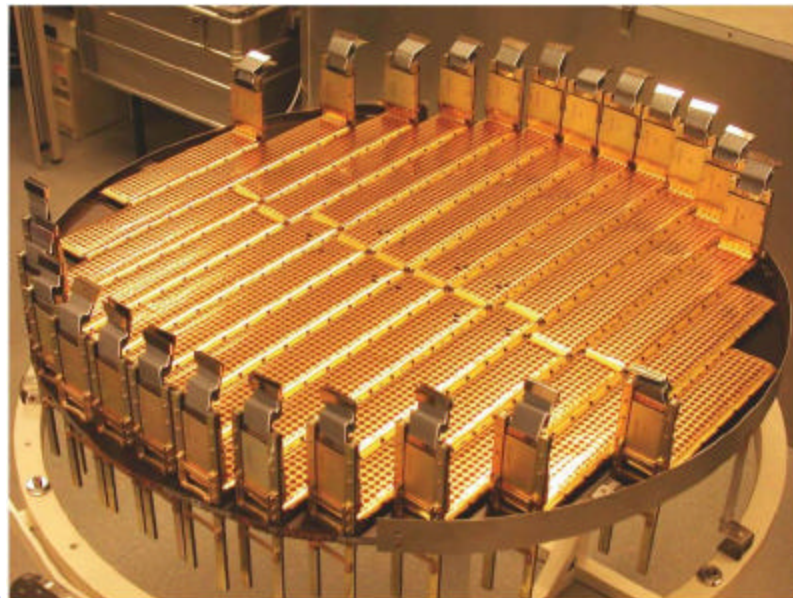
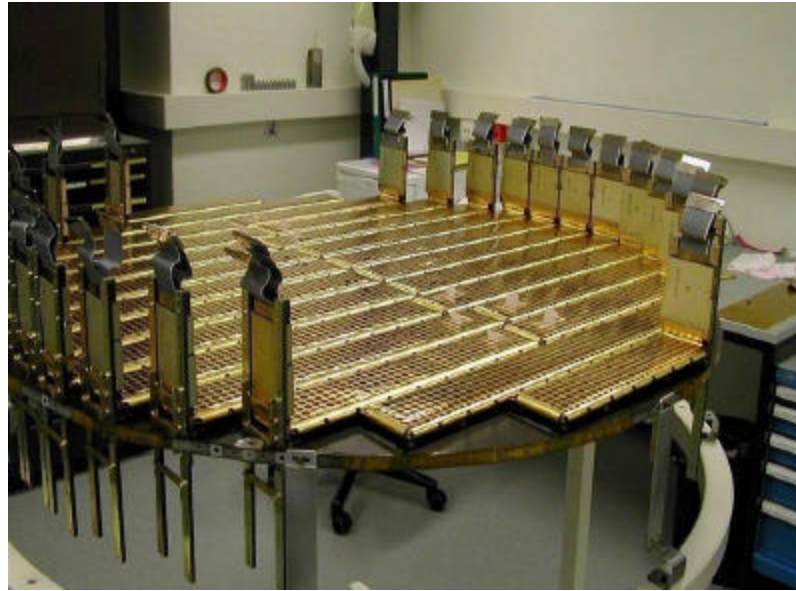
All 8 planes (200,000 channels) have been produced



Sensor positioning (Geneva)



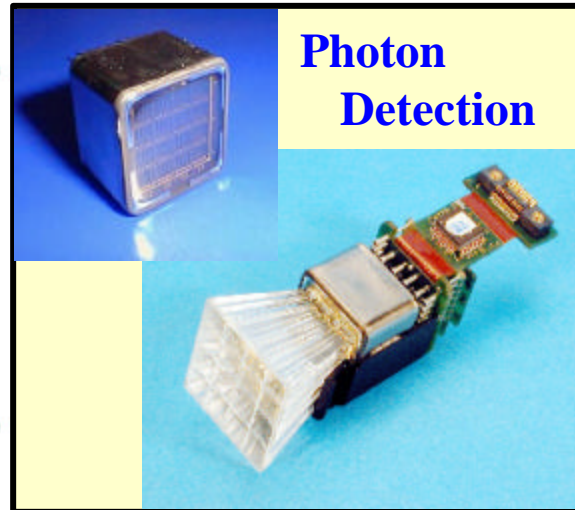
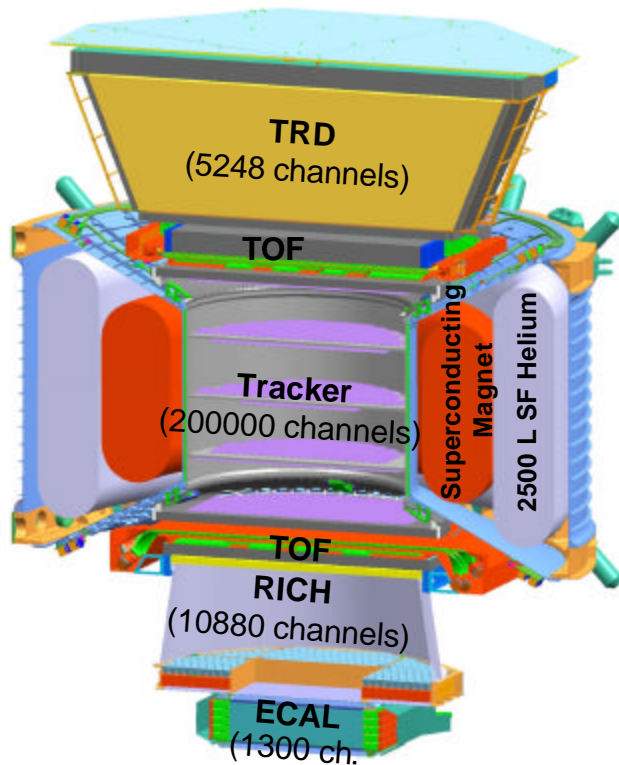
Ultrasonic wire bonding (ETH-Z)



Ring Imaging Cerenkov Counter (RICH)

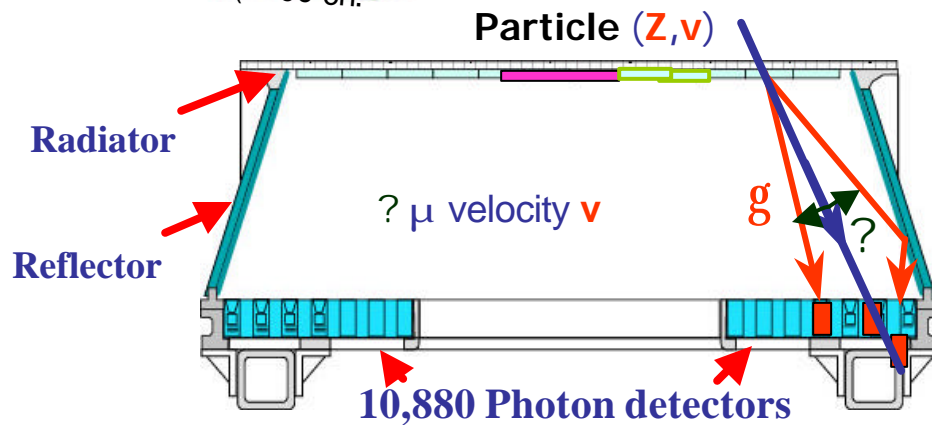
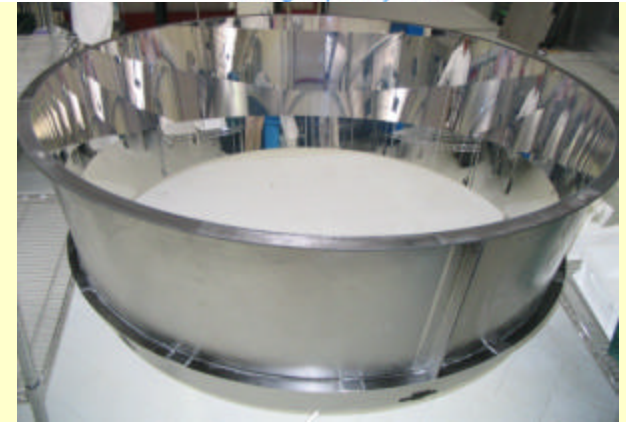
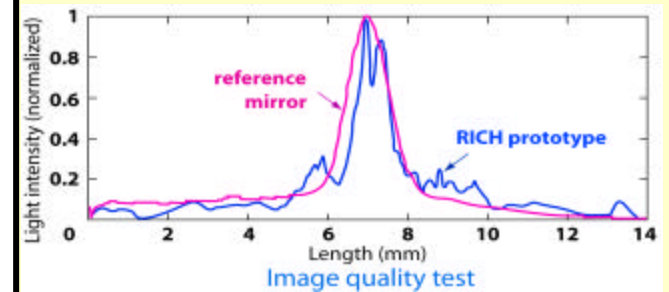


Coordinator: G.Laurenti, INFN-Bologna

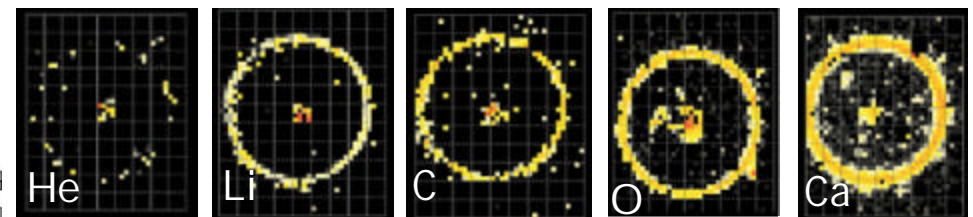


Photon Detection

Reflector

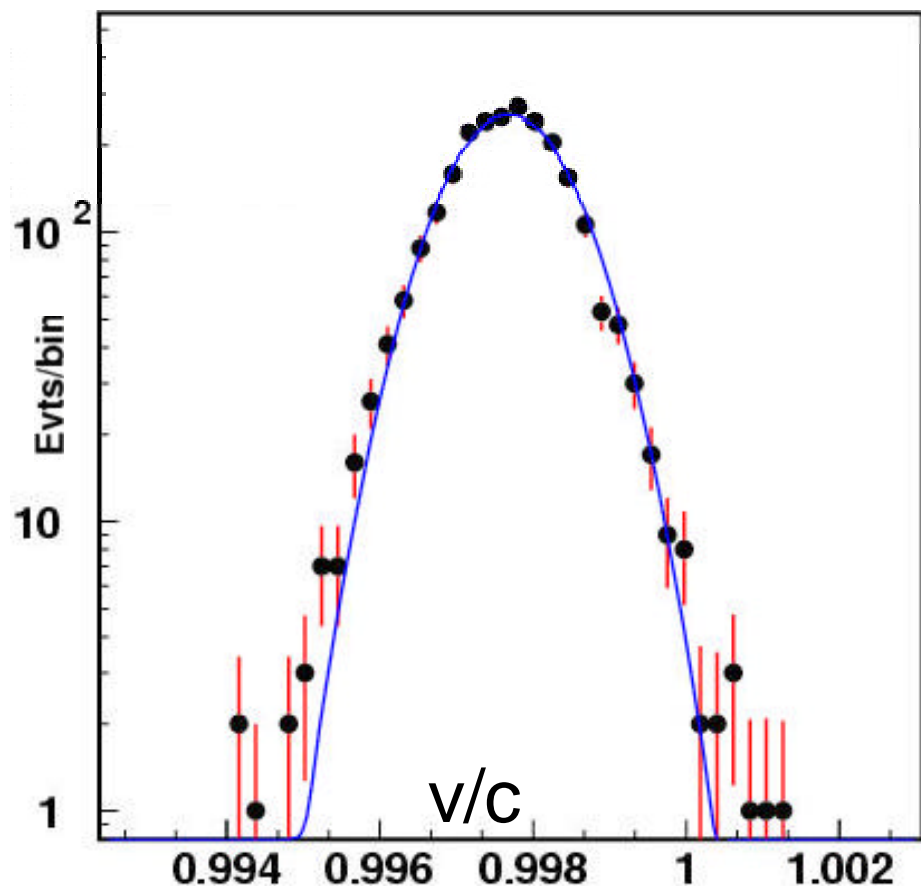


Test results at 158 GeV/n

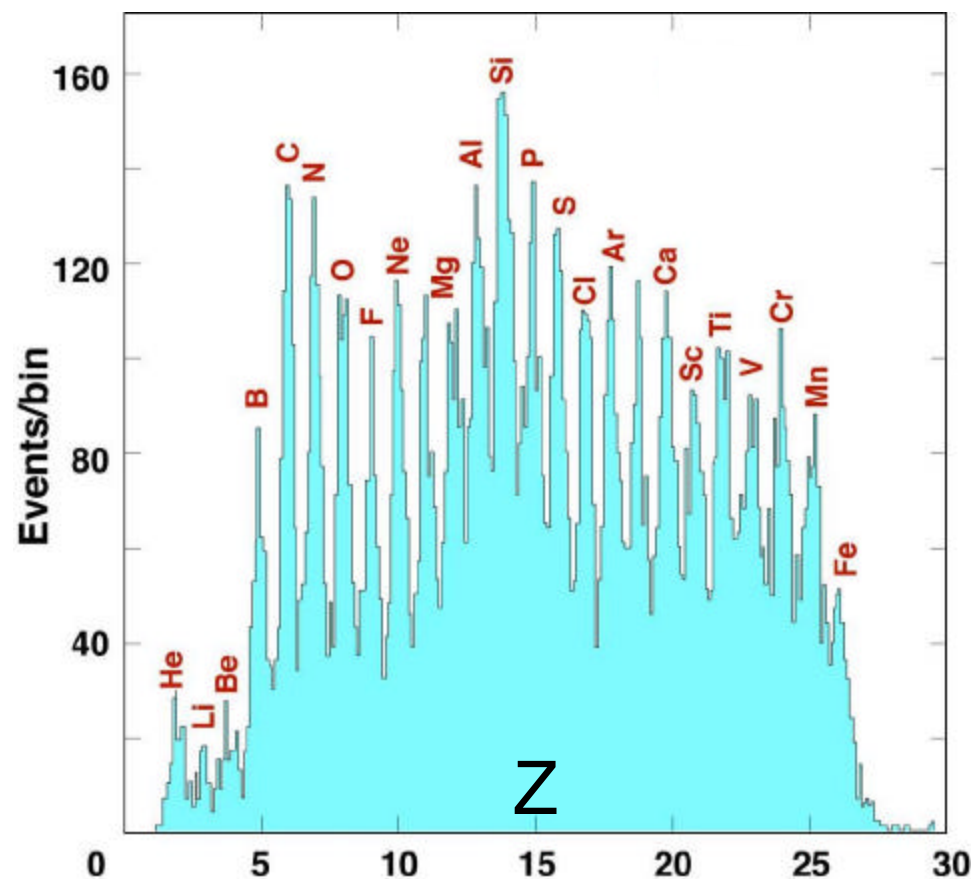


Light (g) intensity $\propto Z^2$

RICH Test Beam Results $E = 158 \text{ GeV/n}$



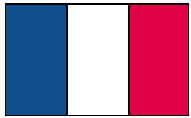
Velocity resolution = 0.1%



Charge measured
up to $Z = 26$ (Fe)

AMS-02 Calorimeter

Measuring energy of electrons and γ rays



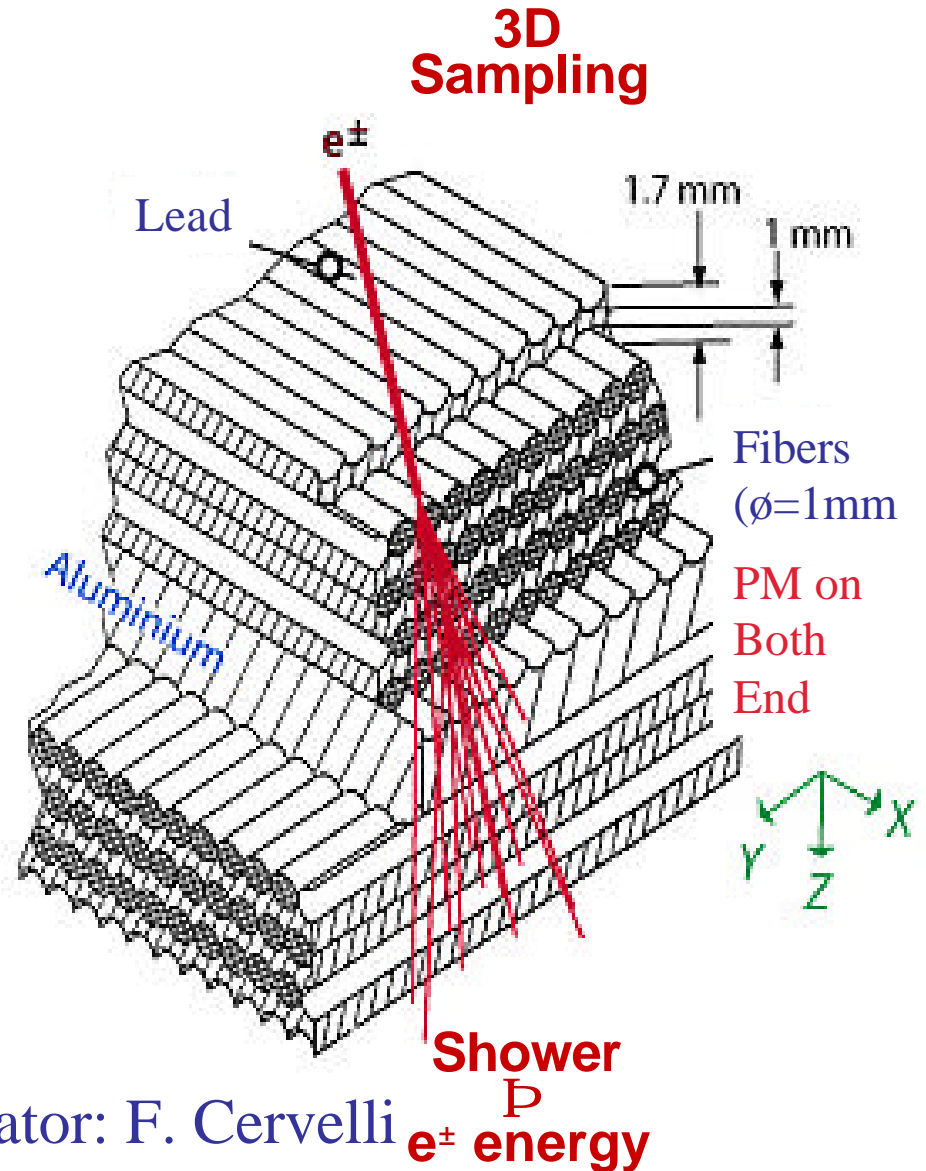
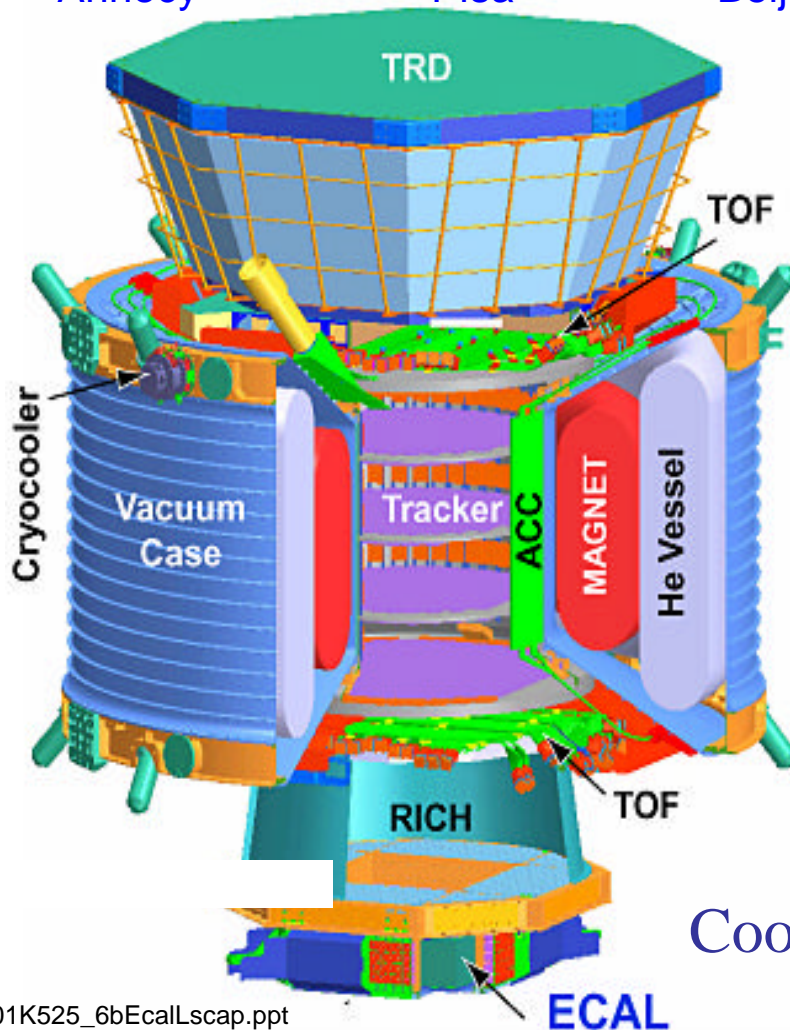
J.P. Vialle
LAPP
Annecy



F. Cervelli
INFN
Pisa



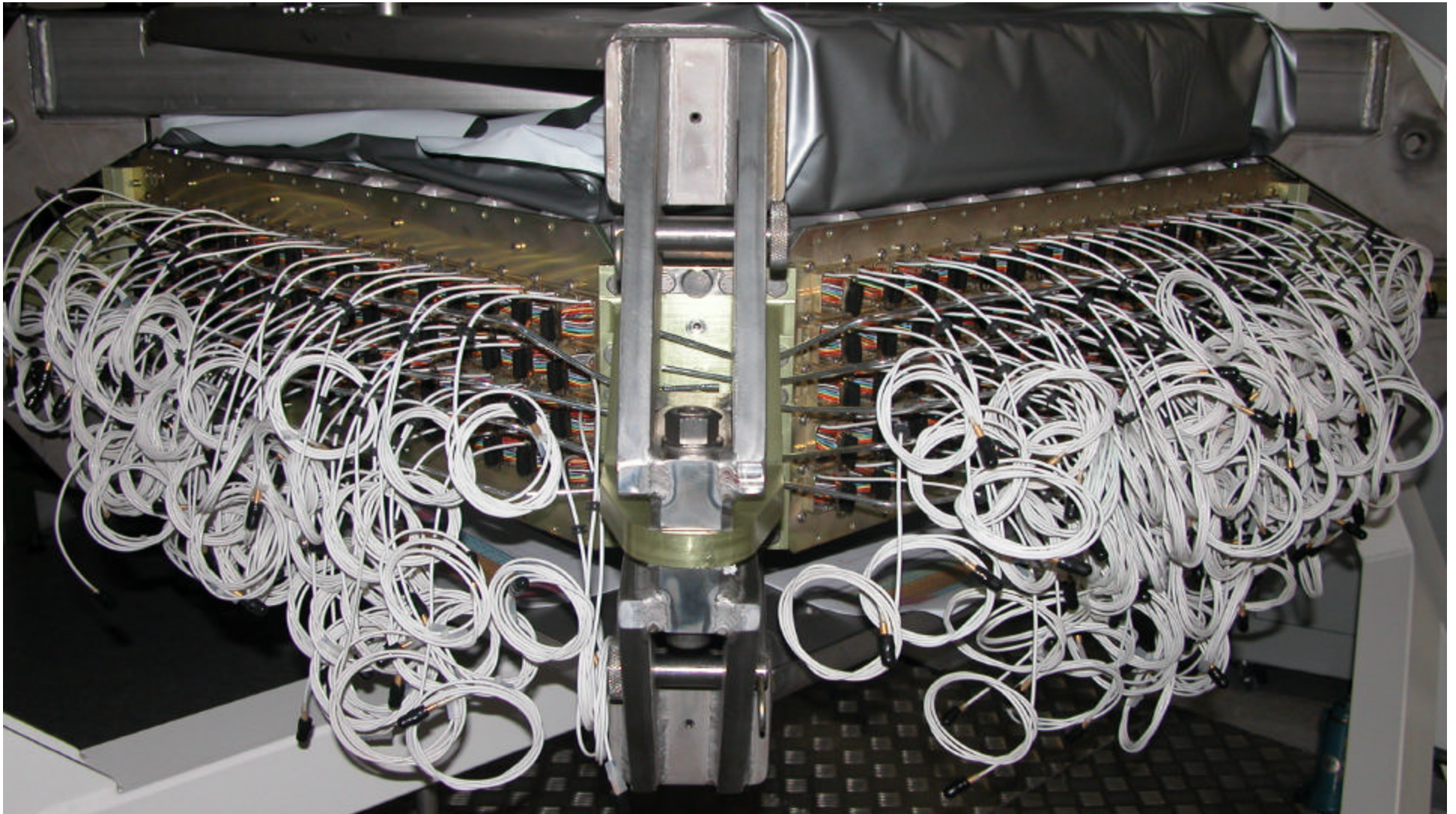
H.S. Chen
IHEP
Beijing



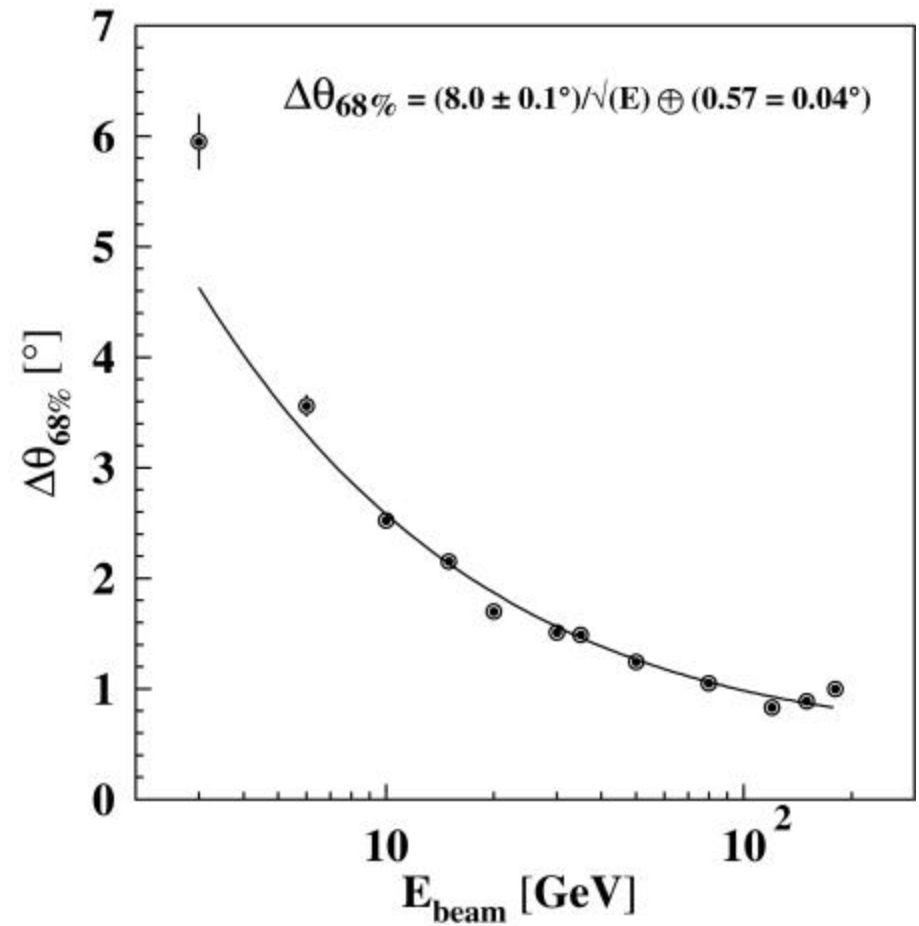
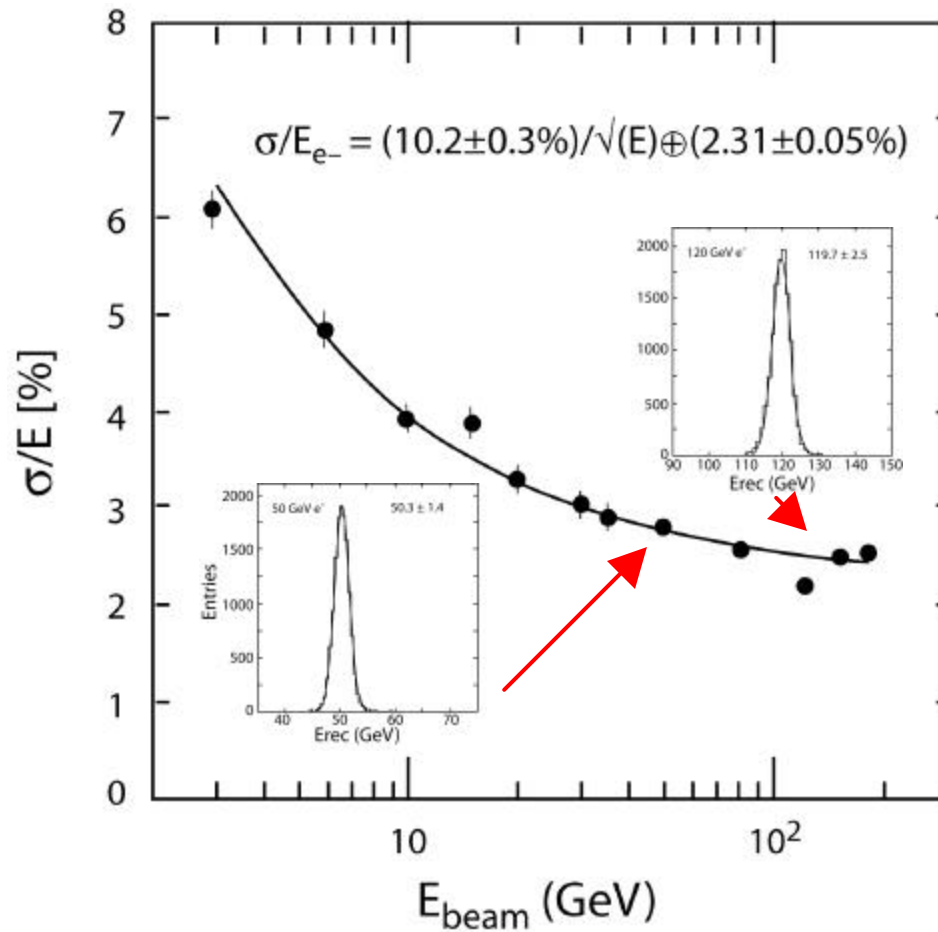
Coordinator: F. Cervelli

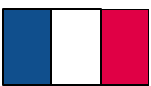
Shower
 P
 e^\pm energy

ECAL Flight Module

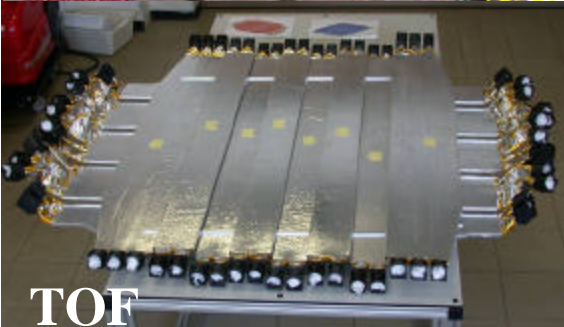


Energy and Angular Resolution of ECAL from test beam measurements

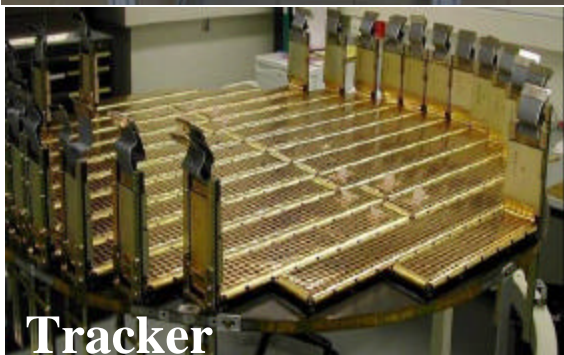




TRD



TOF

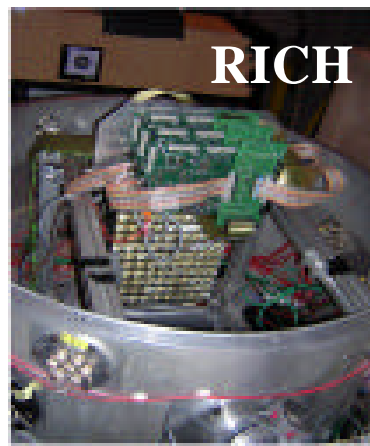
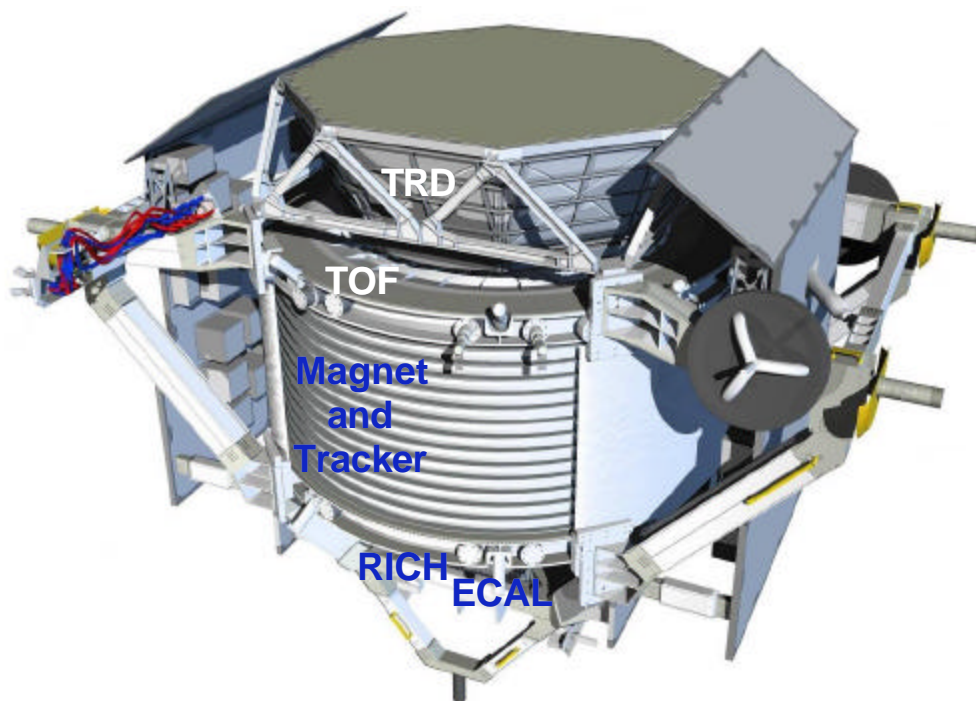


Tracker



ECAL

2005 - 2006 Assembly of all Flight Hardware onto the Superconducting Magnet



RICH

The 650 fast microprocessors



JHIF



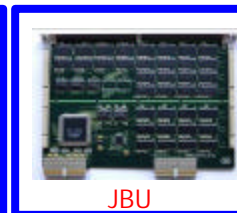
JSBC



JIM-CAN



JIM-HRDL

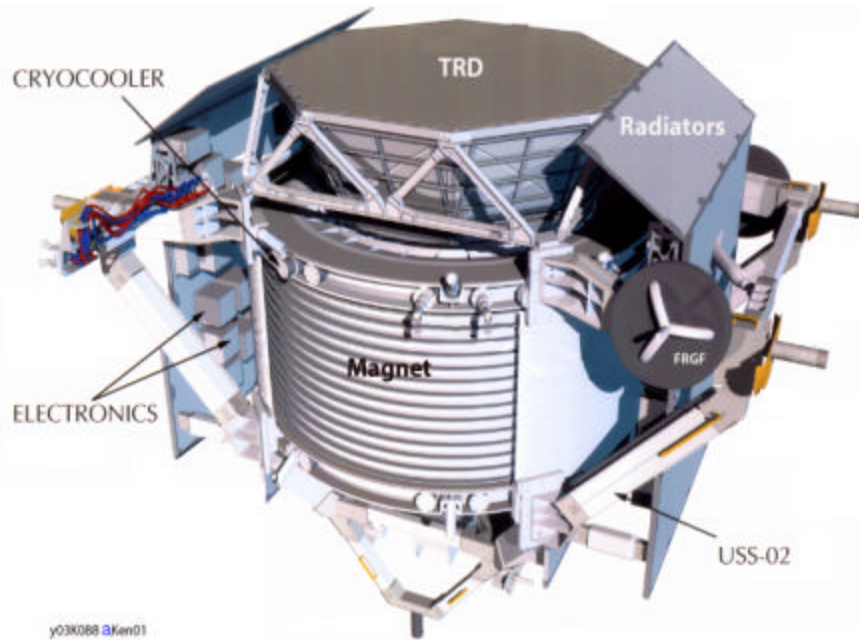


JBU



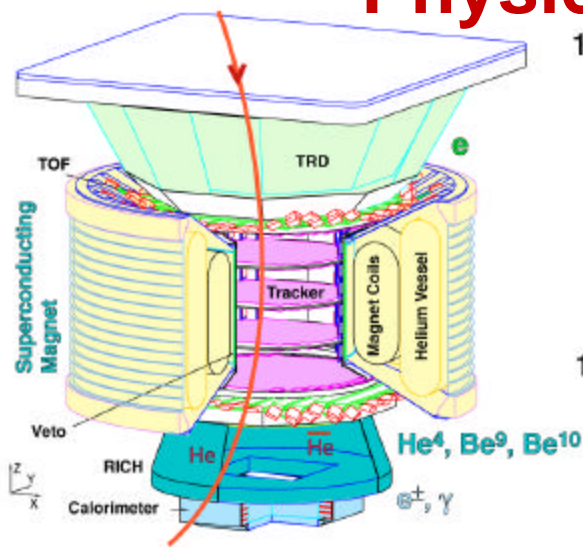
JIM-AMSW&1553

Thermal system

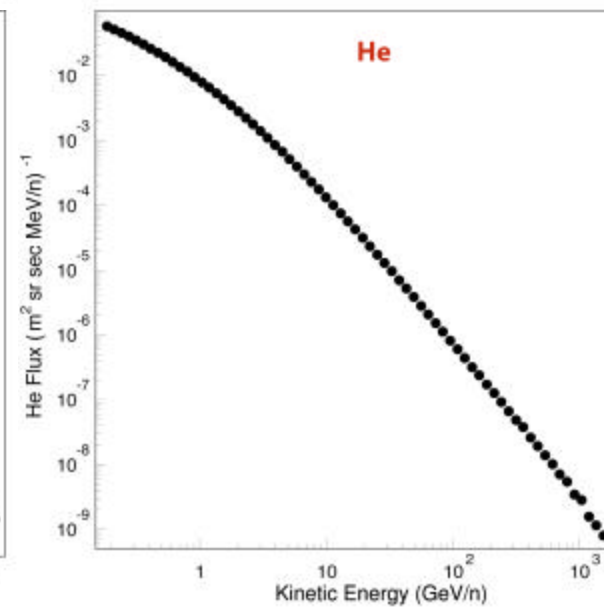
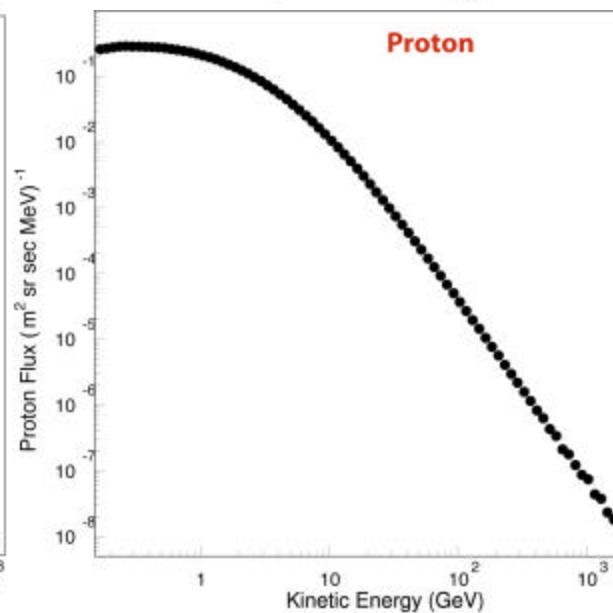
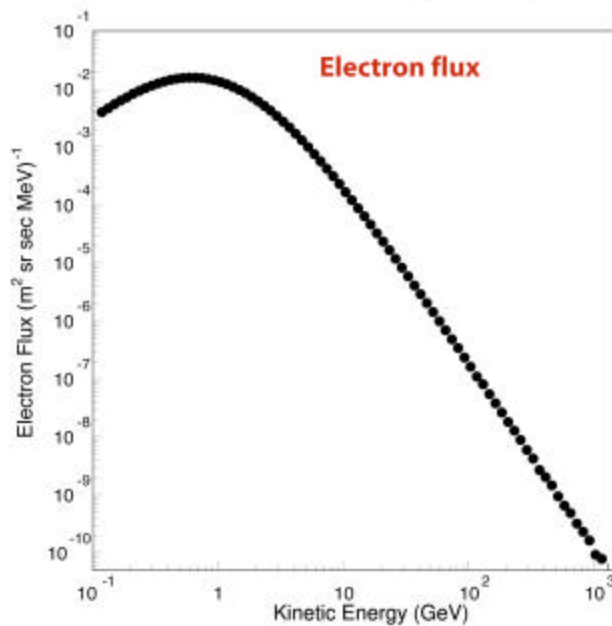
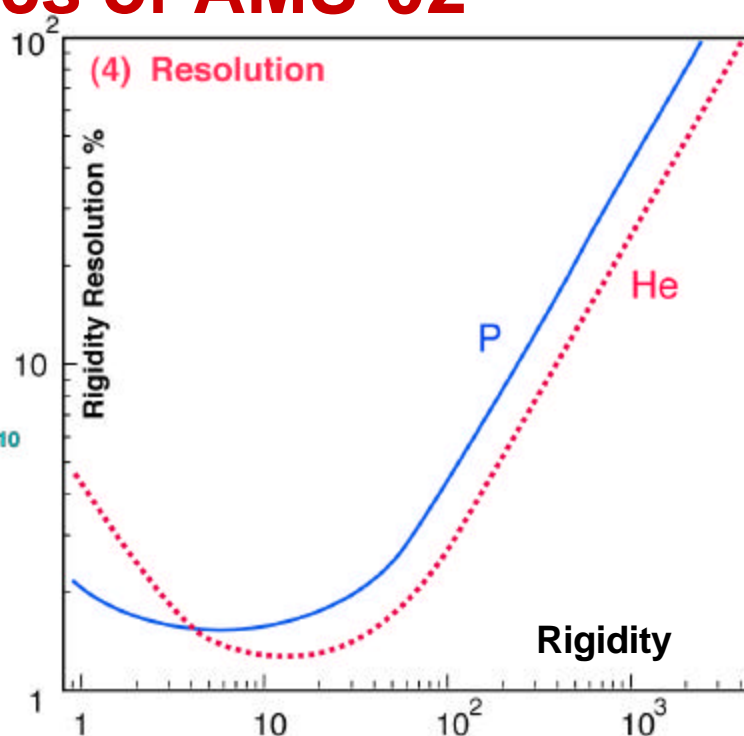


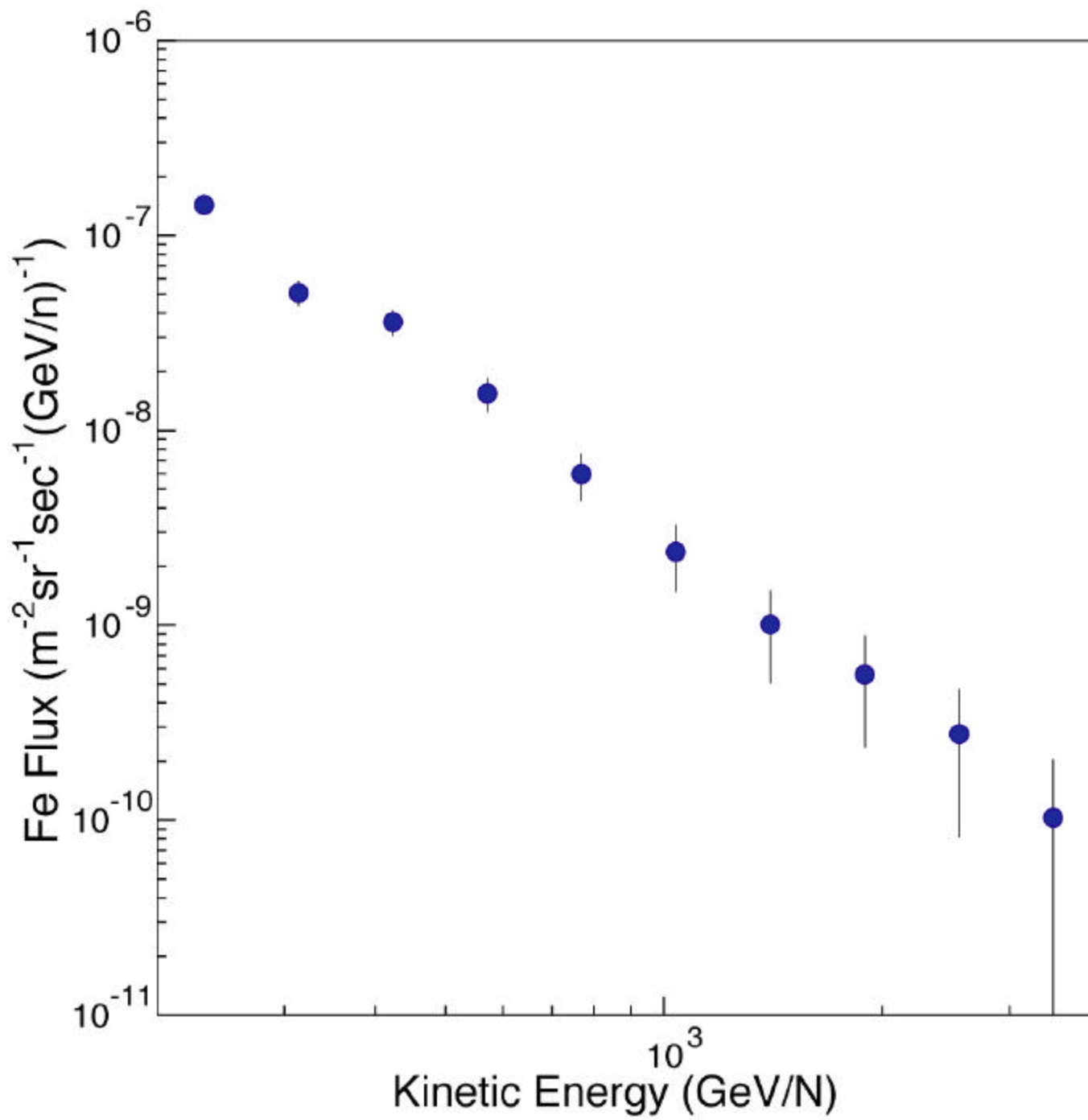
- Thermal-Vacuum Test at ESA (*Noordwijk*)

Physics of AMS-02



- (1) $0.5 \text{ m}^2 \text{ sr}$
- (2) $h/e^+ > 10^6$
- (3) $\Delta\beta/\beta = 0.1\%$

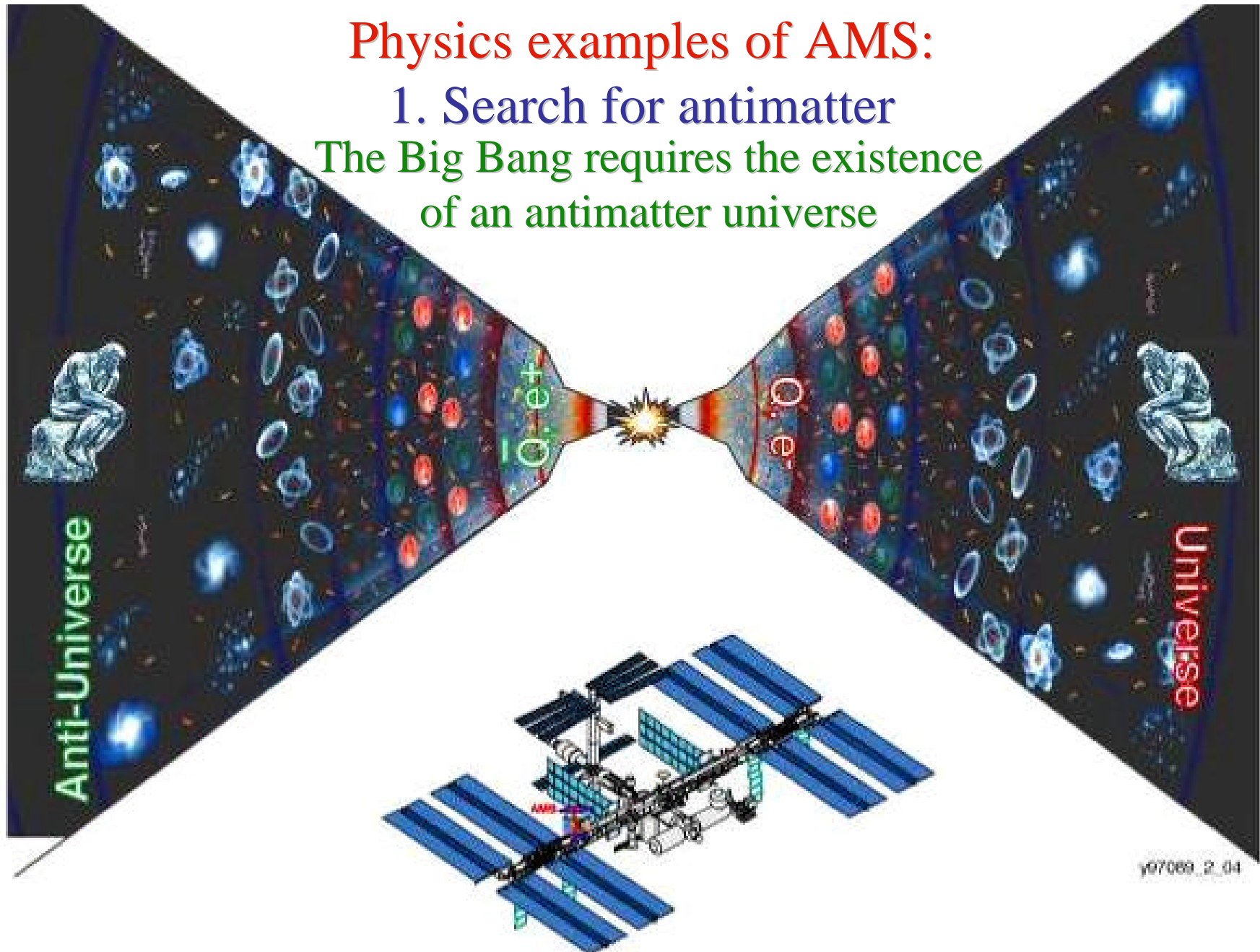




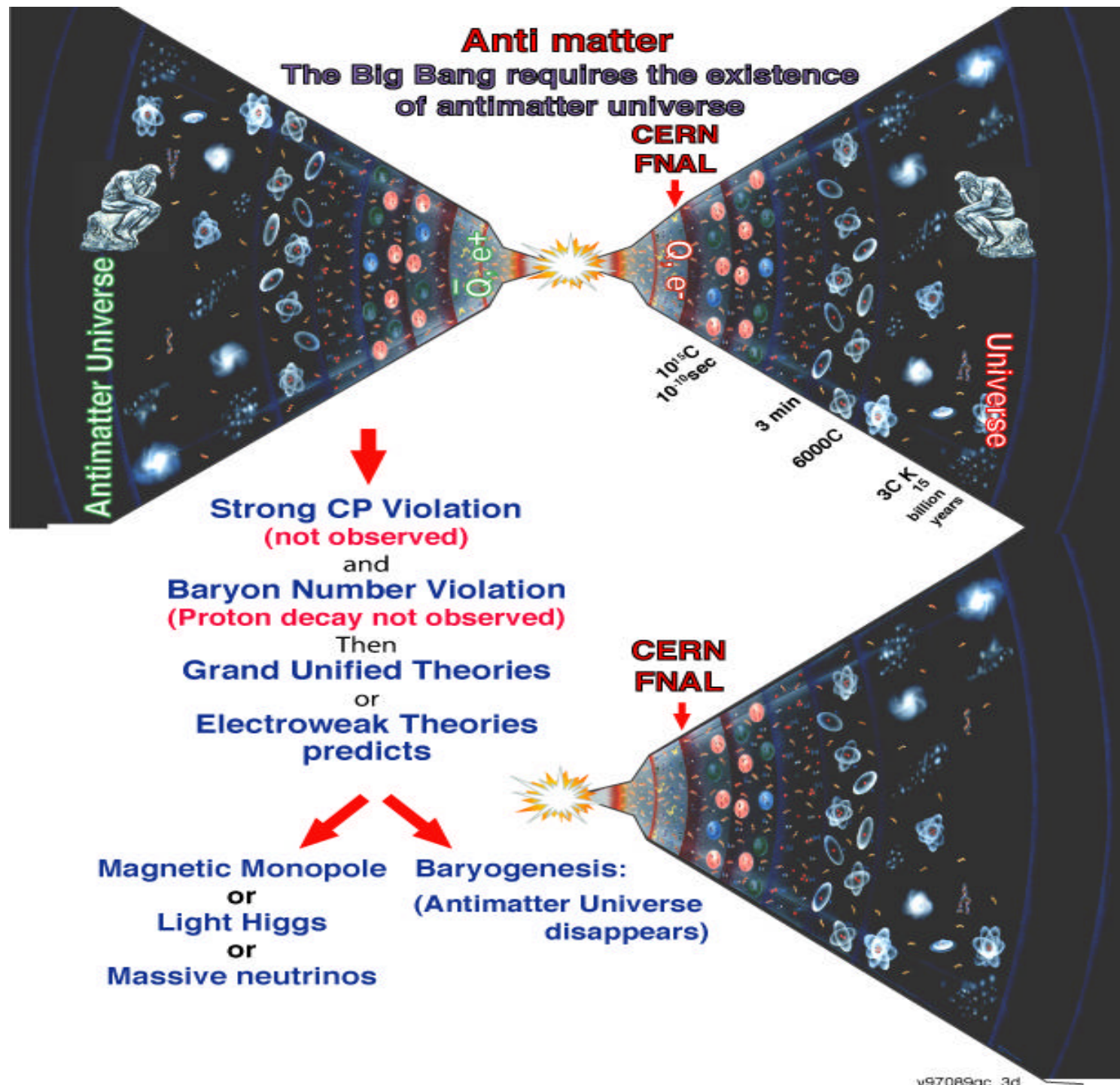
Physics examples of AMS:

1. Search for antimatter

The Big Bang requires the existence
of an antimatter universe



Cosmic antimatter cannot be detected on earth because matter and antimatter will annihilate each other in the atmosphere



Baryon Number Violation

No data from accelerator or proton decay has yet provided evidence for baryon number violation.

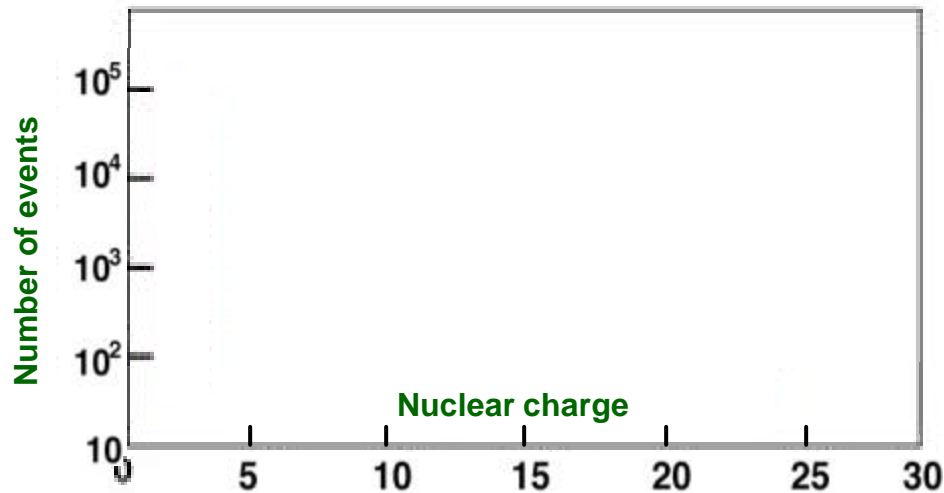
$$\tau > 1.6 \times 10^{33} \text{ Y (e}^+ \pi^0 \text{ mode)}$$

CP-violation

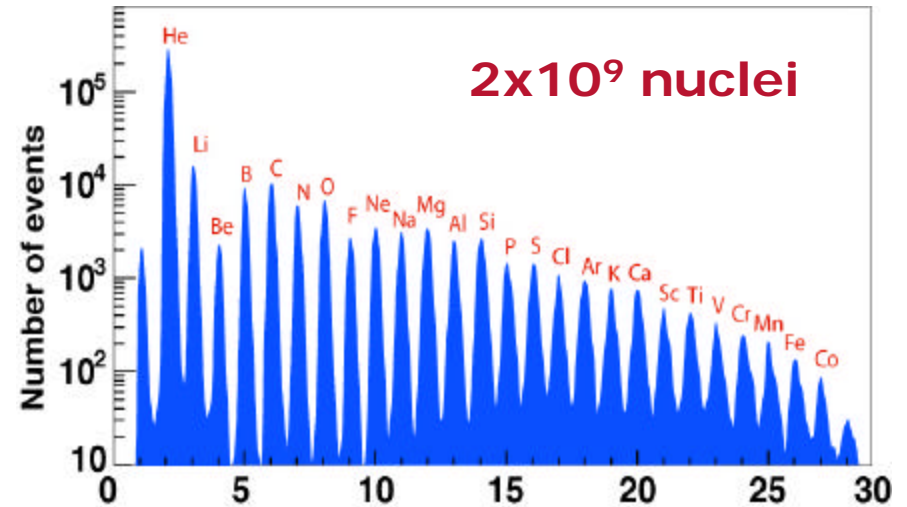
Has been observed in K_L and B only. Both results are in agreement with the Standard Model.

Need new type of CP-violation for Baryogenesis.

Antimatter



Matter



If no antimatter is found \Rightarrow there is no antimatter to the edge (1000 Mpc) of universe.

The physics of antimatter in the universe is based on:

the existence of a strong Time Reversal Violation

the existence of Baryon Number Violation (proton decay)

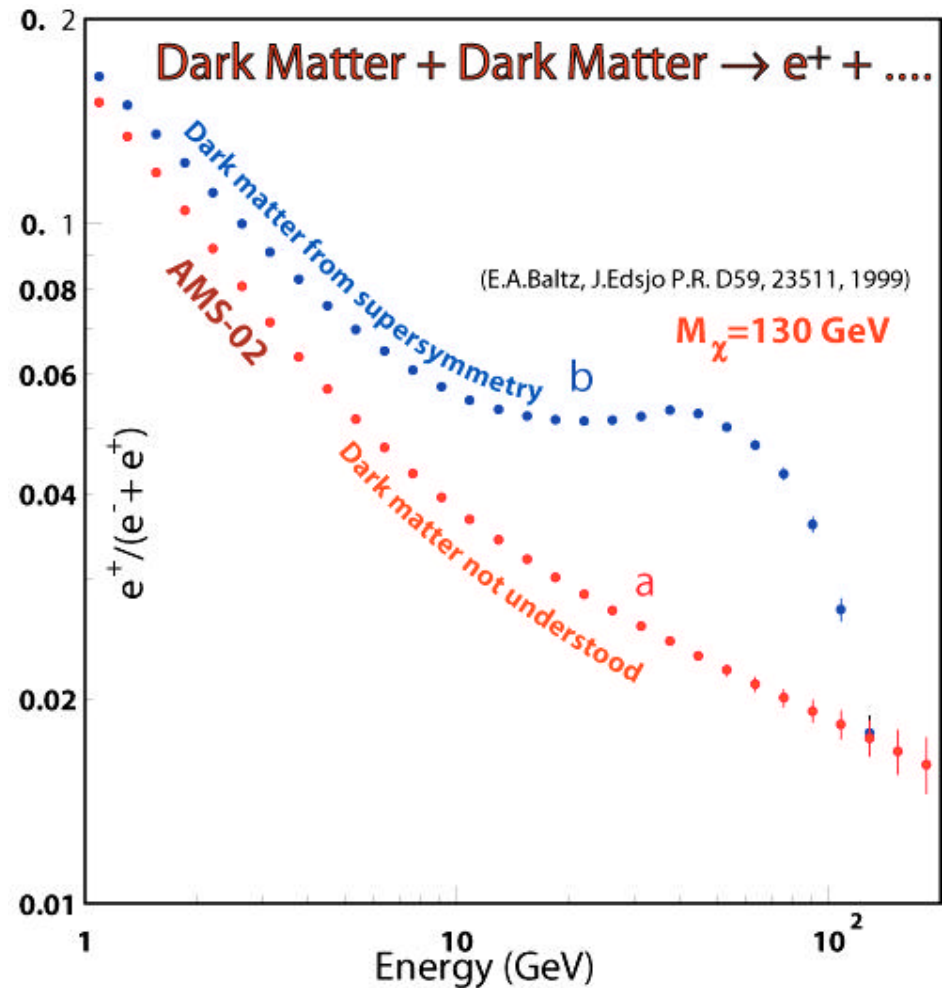
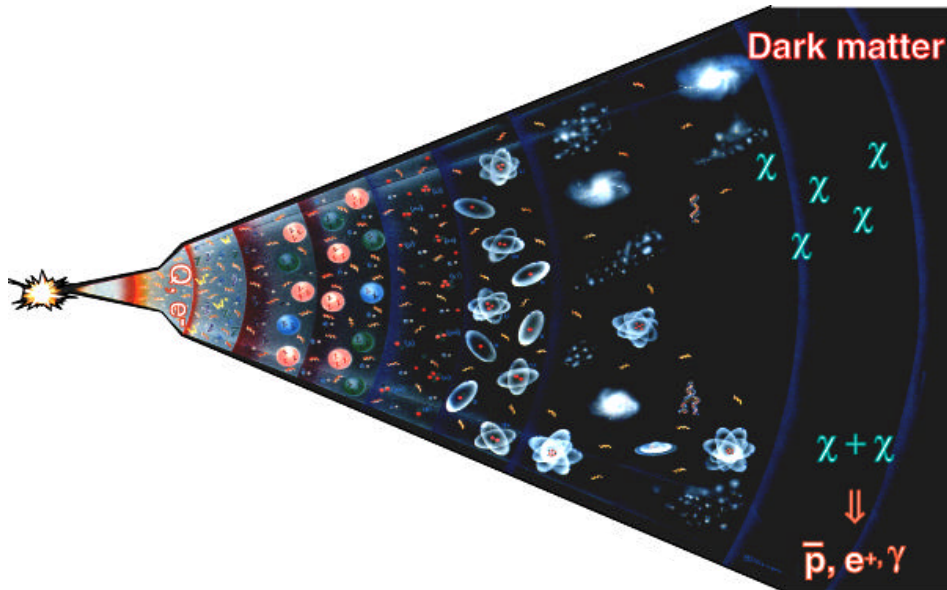
Grand Unified Theory

Electroweak Theory

This is the main research topic for the current and next generation of accelerators world wide

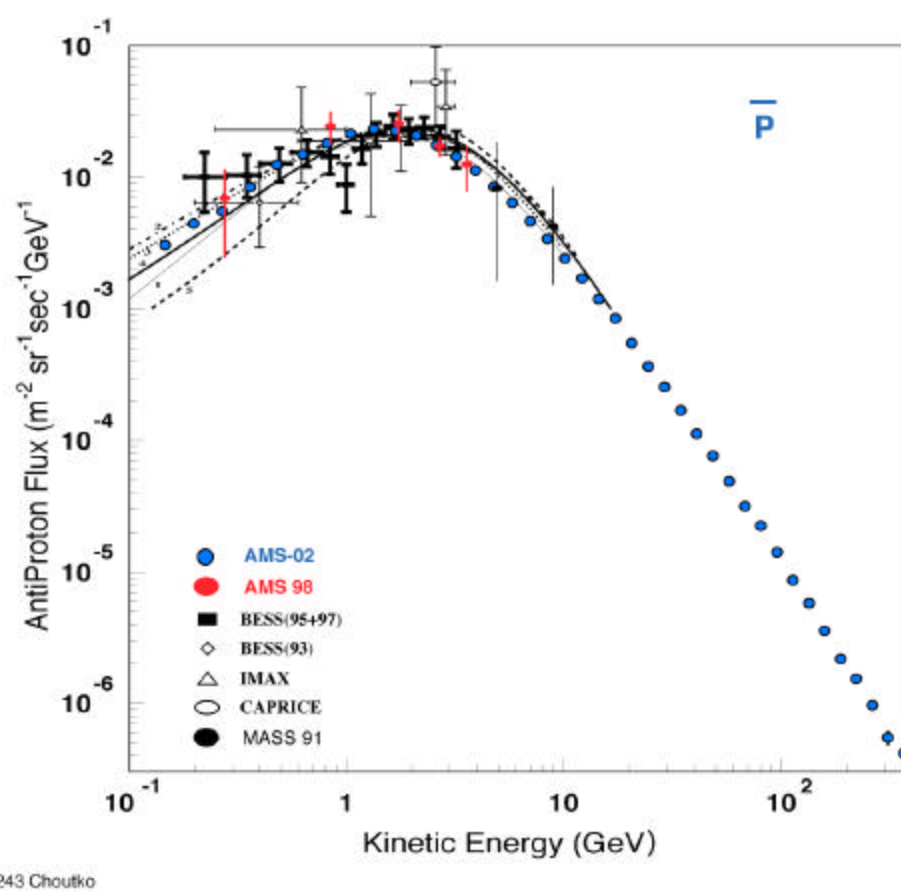
2. Dark matter

There are many theoretical suggestions that **SUSY particles** (c) are at least part of the Dark matter.

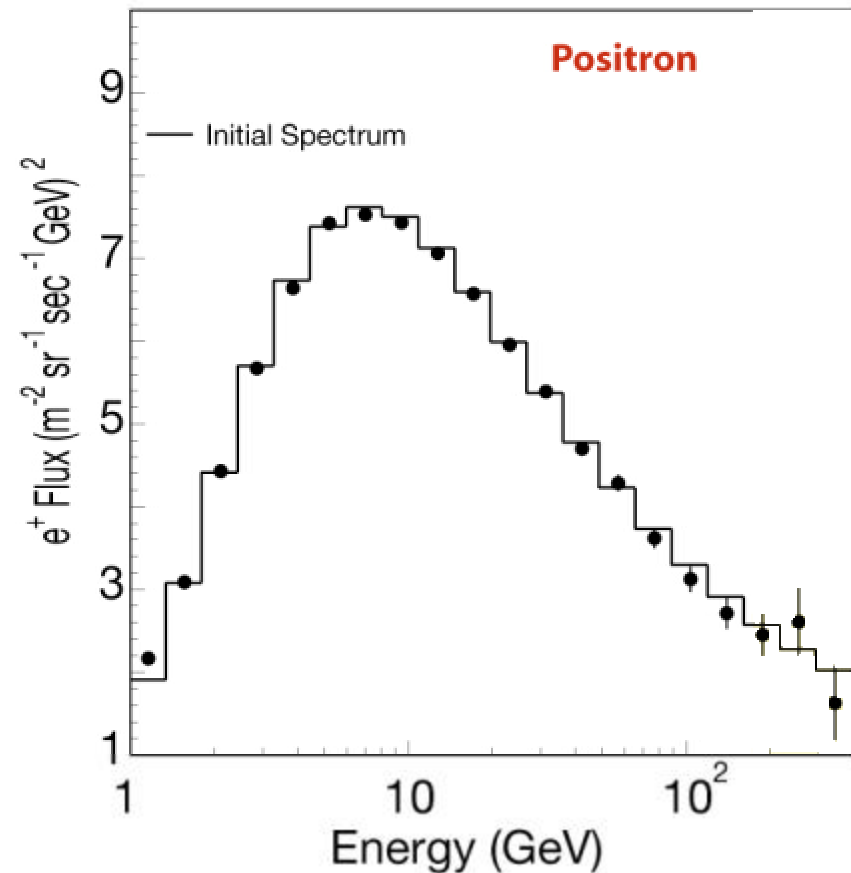


AMS and Dark Matter

The fluxes of antiprotons and positrons may be sensitive to dark matter



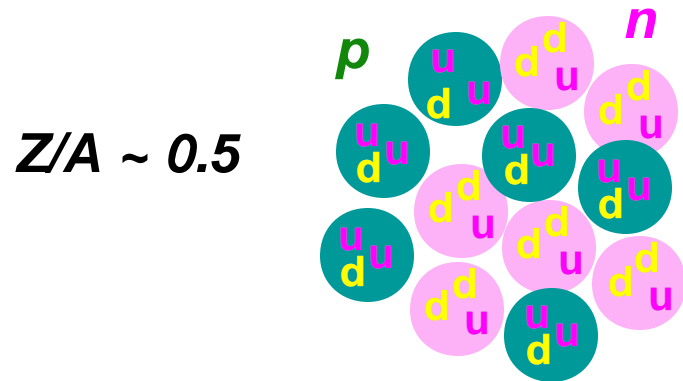
y02K243 Choutko



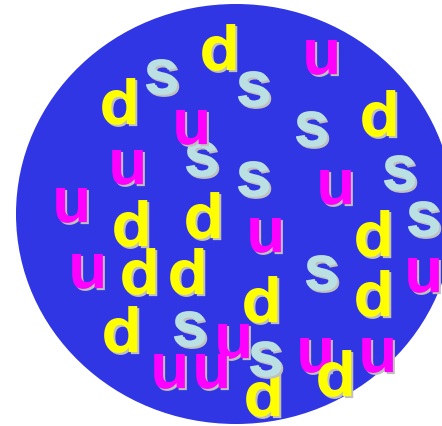
3. Strange Quark Matter – “Strangelets”

Known quarks: u, d, s, c, b, t

Carbon Nucleus



Strangelet



$Z/A < 0.12$

Strangelets: a single “super nucleon” with many u, d & s

- Stable for masses $A > \sim 10$, with no upper limit
- “Neutron” stars may be composed of one big strangelet

Searches

with terrestrial samples – low sensitivity.

with lunar samples – limited sensitivity.

in accelerators – cannot be produced at an observable rate.

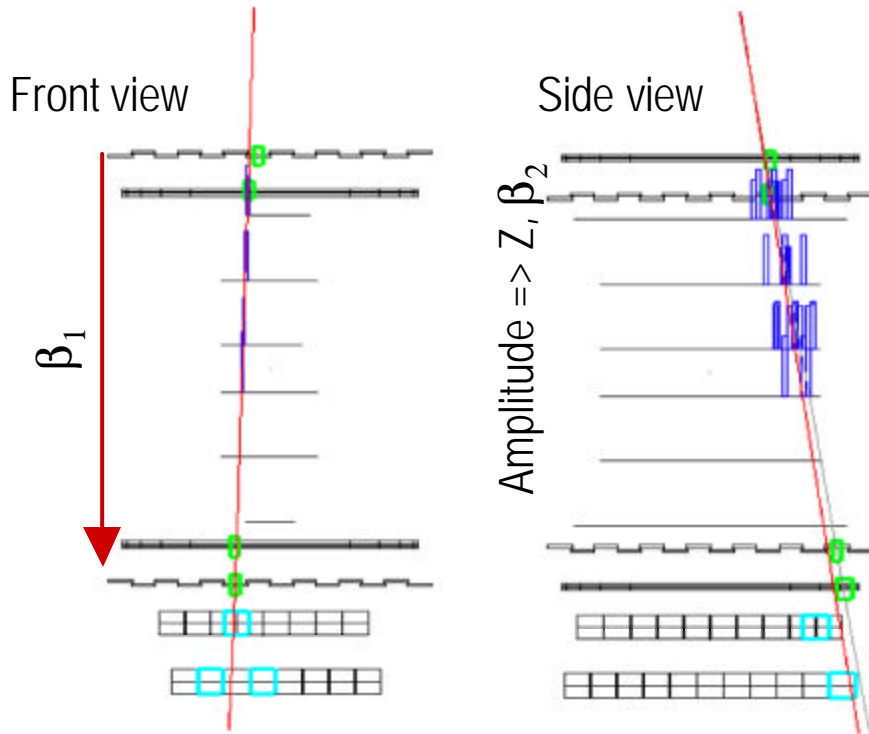
in space – candidates...

Jack Sandweiss, Yale

Strangelet candidate from AMS-01

Observed 5 June 1998 11:13:16 UTC

Lat/Long= -44.38°/+23.70°, Local Cutoff 1.95 ± 0.1 GV, Angle= 77.5° from local zenith



Rigidity = 4.31 ± 0.38 GV

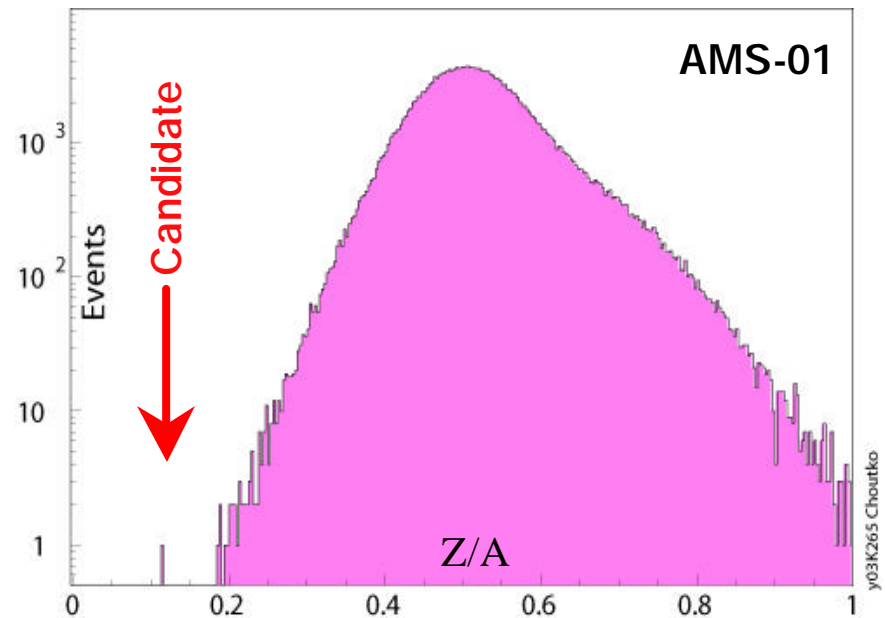
Charge $Z = 2$

$b_1 = b_2 = 0.462 \pm 0.005$

Mass = 16.45 ± 0.15 GeV/ c^2

$Z/A = 0.114 \pm 0.01$

Flux ($1.5 < E_k < 10$ GeV) = 5×10^{-5} (m² sr sec)⁻¹

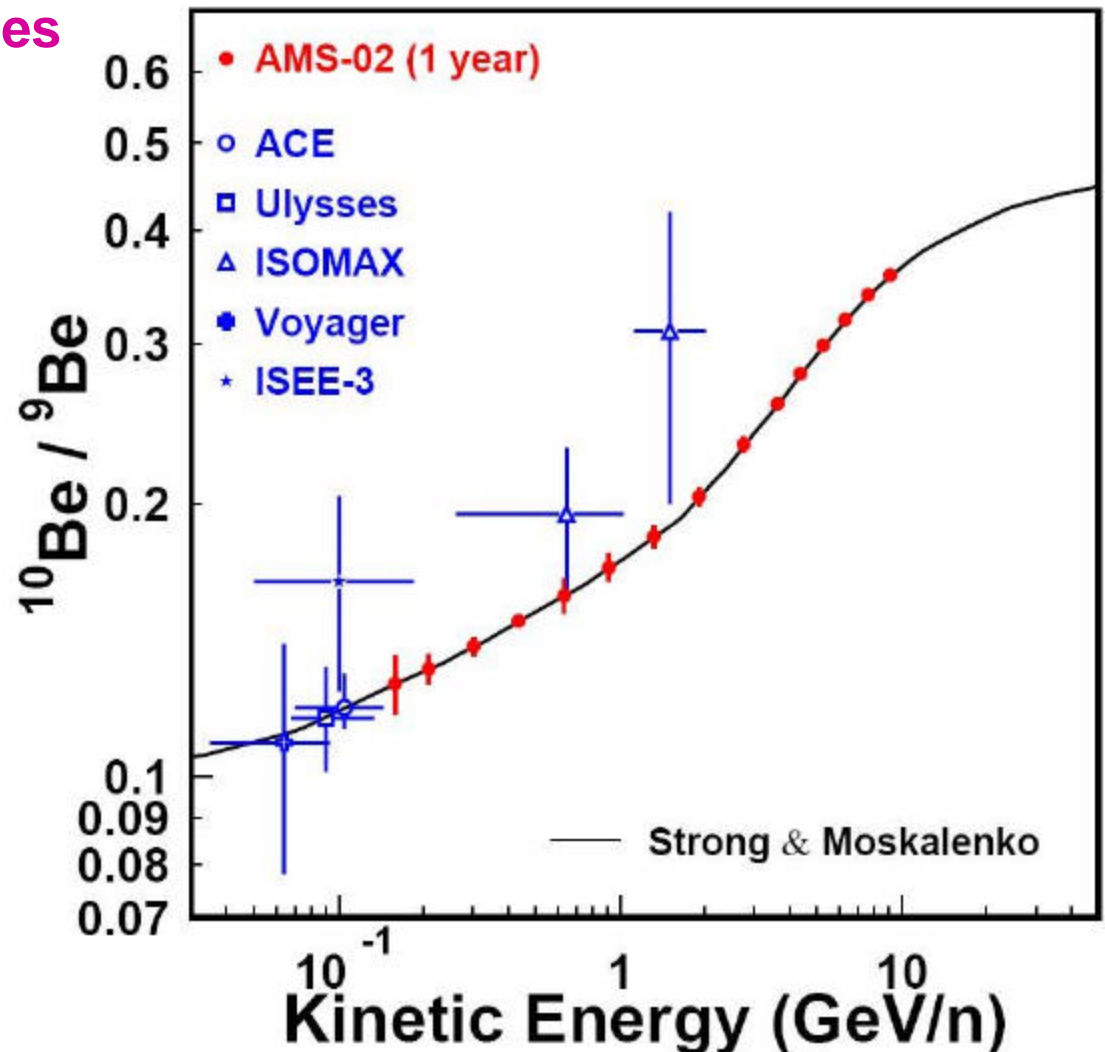


Background probability $< 10^{-3}$

4. The ratio $^{10}\text{Be} / ^9\text{Be}$ determines

i) the cosmic ray confinement time in the galaxy, and

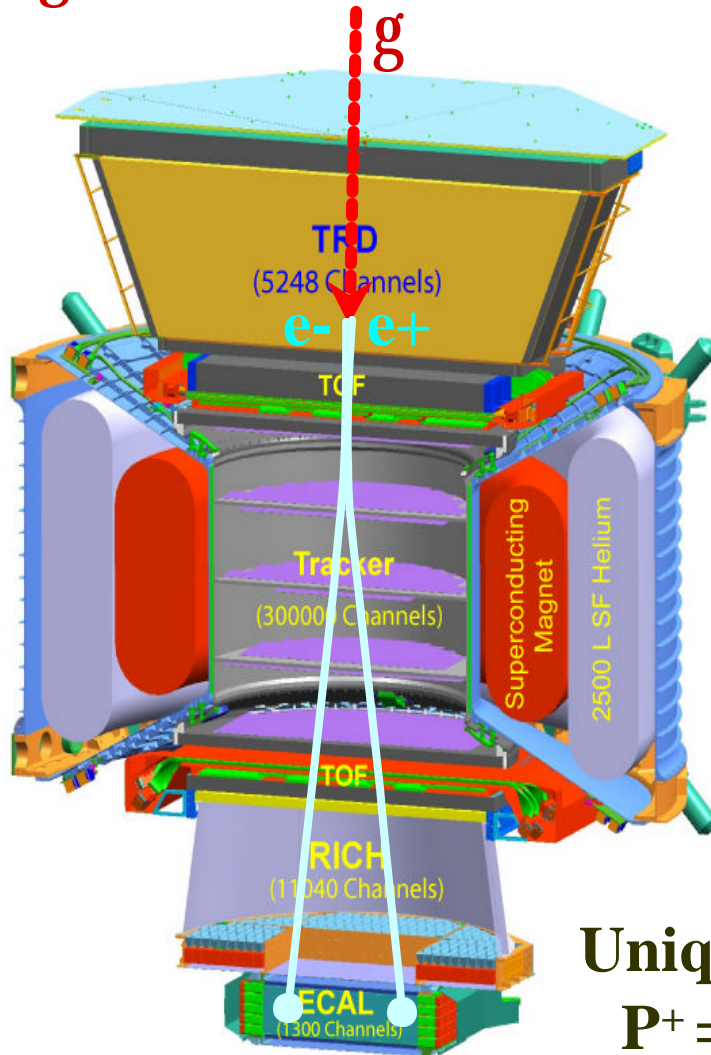
ii) the mean density of interstellar material traversed by cosmic rays.



One of the most important measurements in cosmic ray physics

Identify γ Sources with AMS

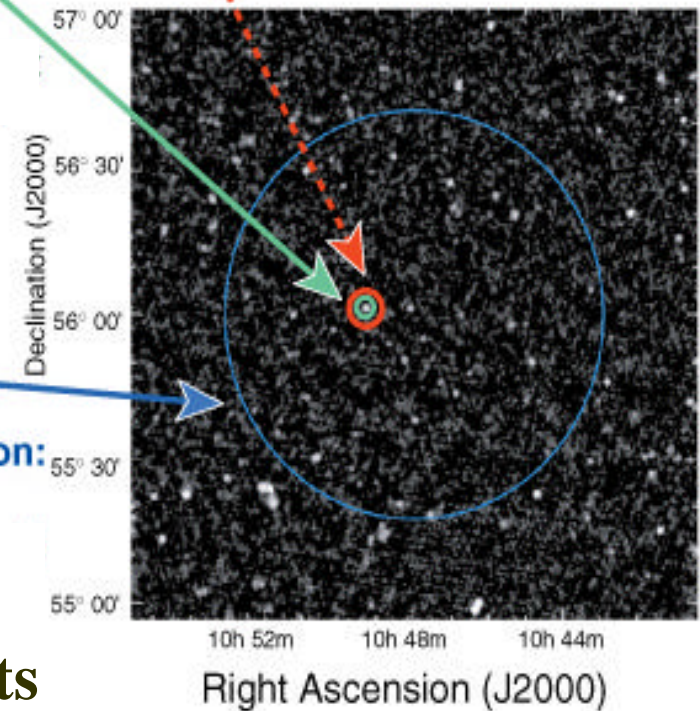
$\gamma \rightarrow e^+ e^-$



AMS
Source localization:
($E > 10$ GeV) $< 2'$

GLAST
Source localization:
 $< 5'$

EGRET
• Source localization:
 $< 30'$



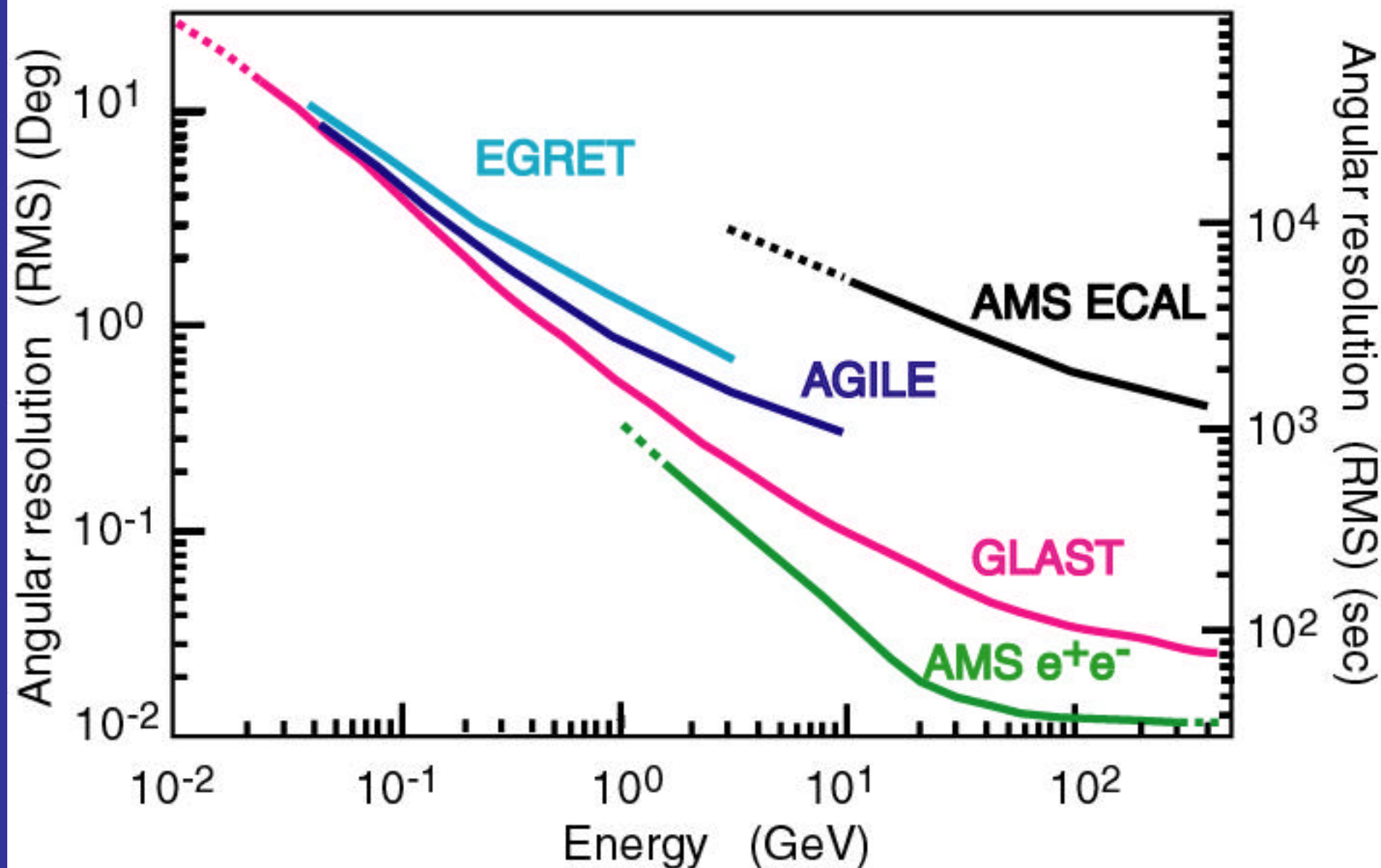
Unique constraints

$$P^+ = E^+ = P^- = E^-$$

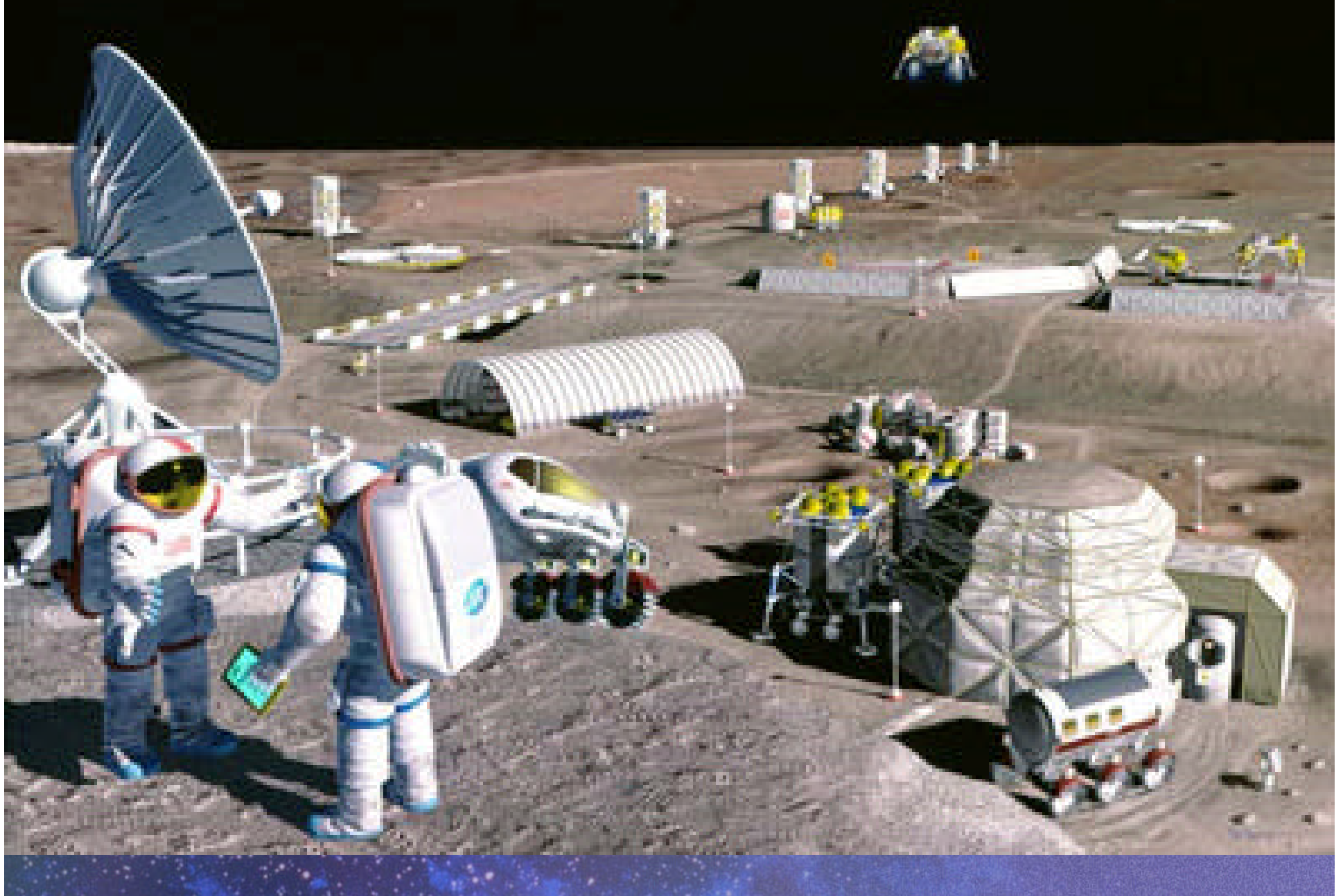
$$?^+ \sim ?^-$$

y01K062_2a

Angular resolution of space born γ -ray Detectors

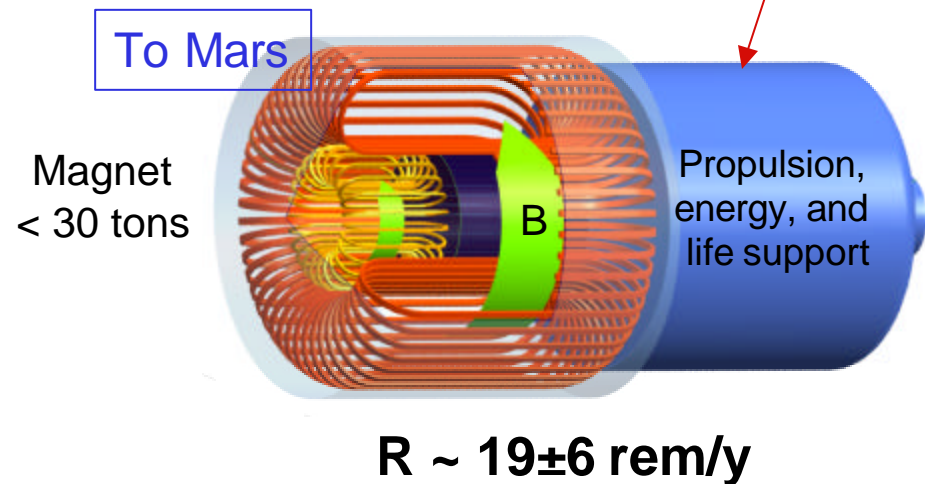
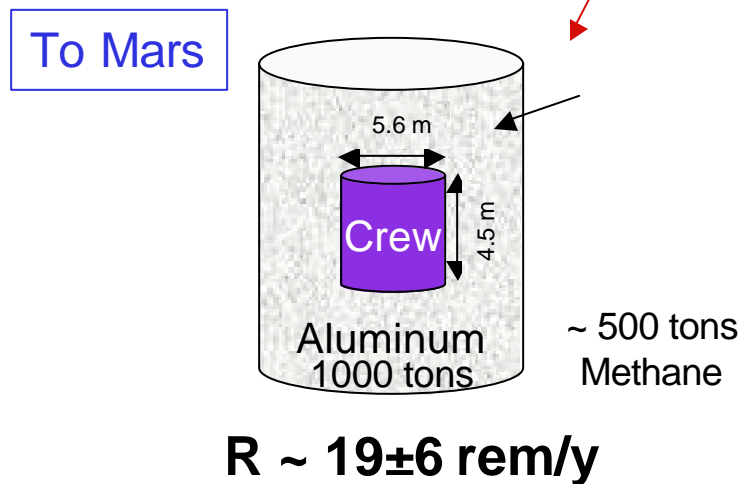
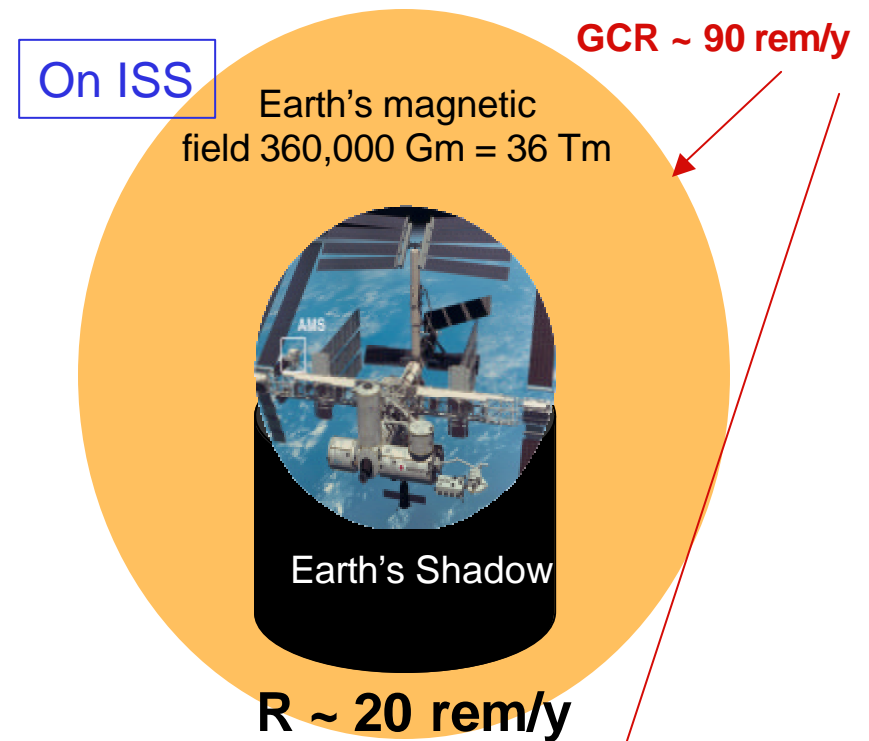
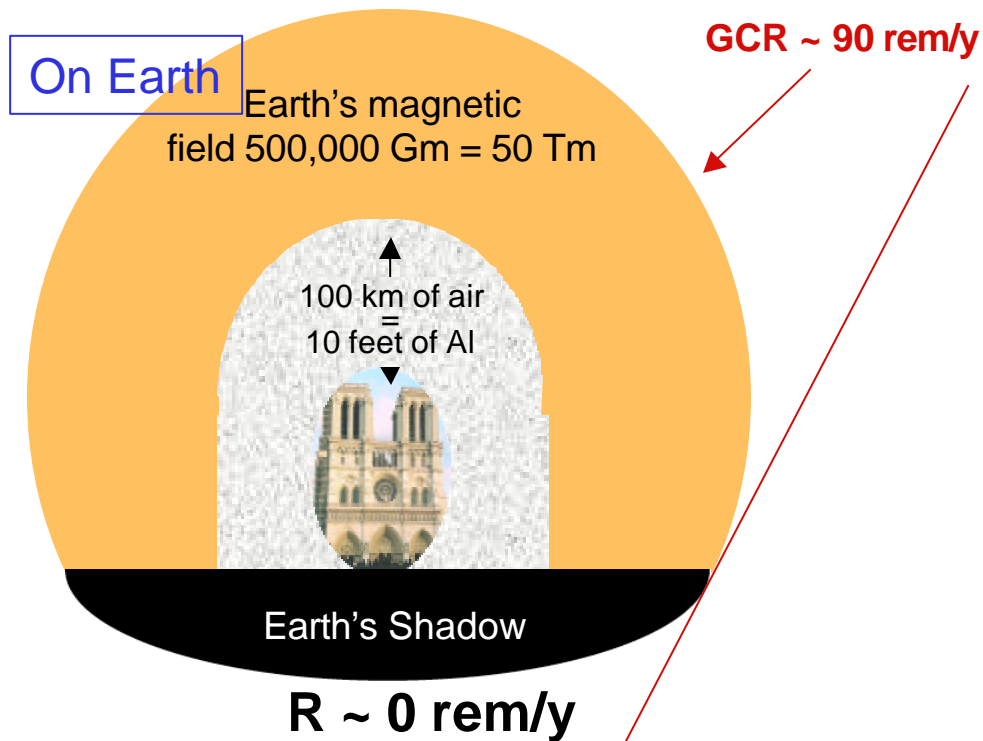


NASA Exploration: Establish a base on the Moon



Future NASA Exploration: Manned mission to Mars



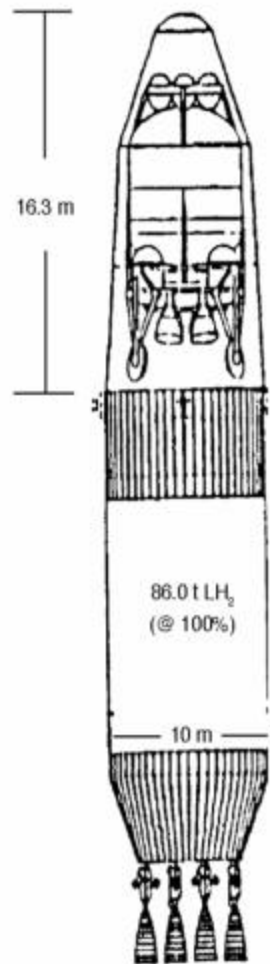


Of the 26892 man-days spent in space only 303 have been in Apollo Mission outside the magnetosphere (1.1%)

Mars reference design: weight estimates

2009 Piloted Mission 1
Surface Hab with Crew Lander

Exploration payload (Earth-Return Habitat Element):



212.1 t

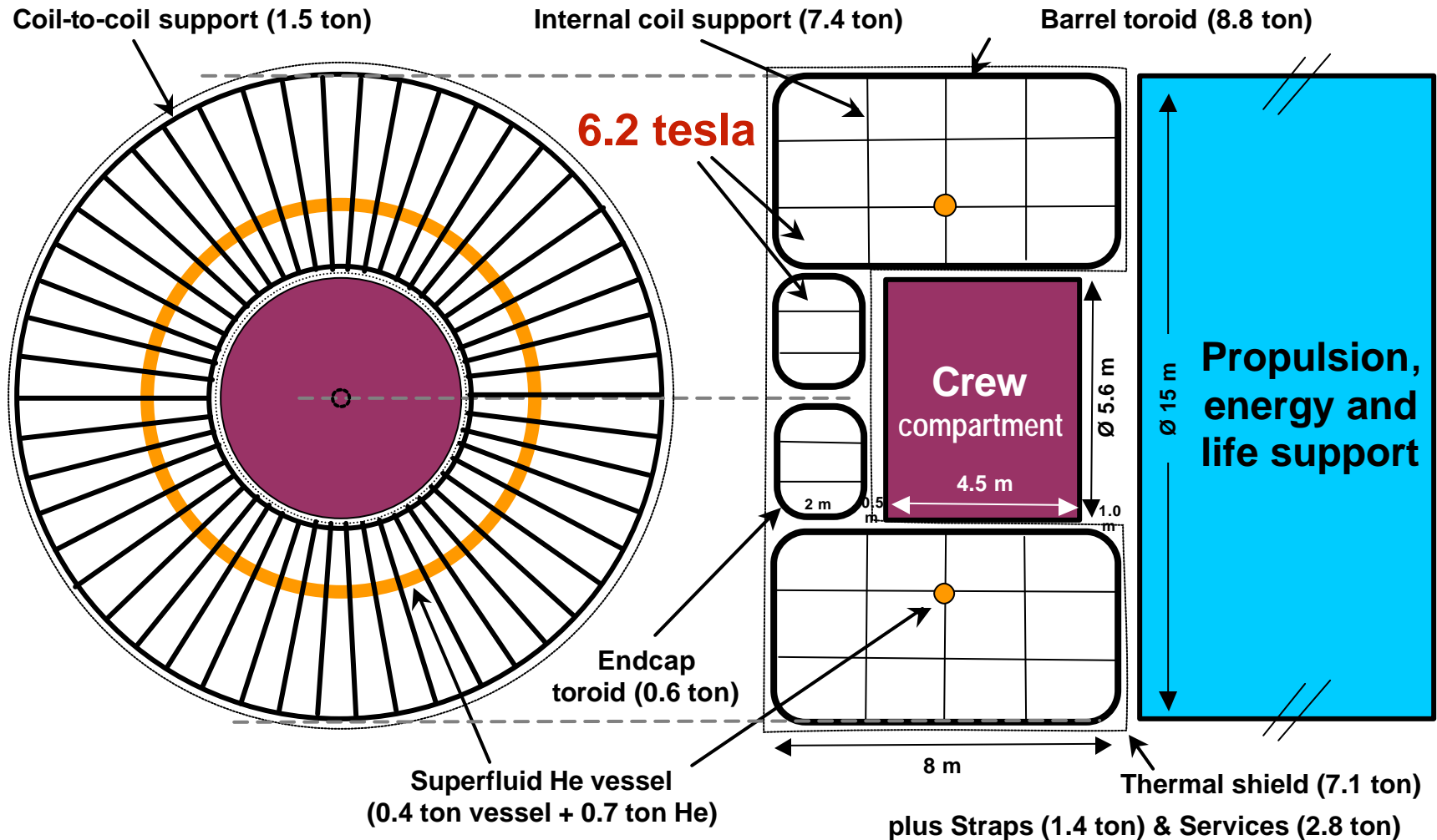
Subsystem	Subsystem Mass (tonnes)	Consumables Subtotal (tonnes)	Dry Mass Subtotal (tonnes)
Physical/chemical life support	6.00	3.00	3.00
Plant growth	0.00	0.00	0.00
Crew accommodations	22.50	17.50	5.00
Health care	2.50	0.50	2.00
Structures	10.00	0.00	10.00
EVA	4.00	3.00	1.00
Electrical power distribution	0.50	0.00	0.50
Communications and information management	1.50	0.00	1.50
Thermal control	2.00	0.00	2.00
Power generation	0.00	0.00	0.00
Attitude control	0.00	0.00	0.00
Spares/growth/margin	3.50	0.00	3.50
Radiation shielding	0.00	0.00	0.00
Science	0.90	0.00	0.90
Crew	0.50	0.50	0.00
Total estimate	53.90	24.50	29.40



Magnetic radiation protection system for Mars mission

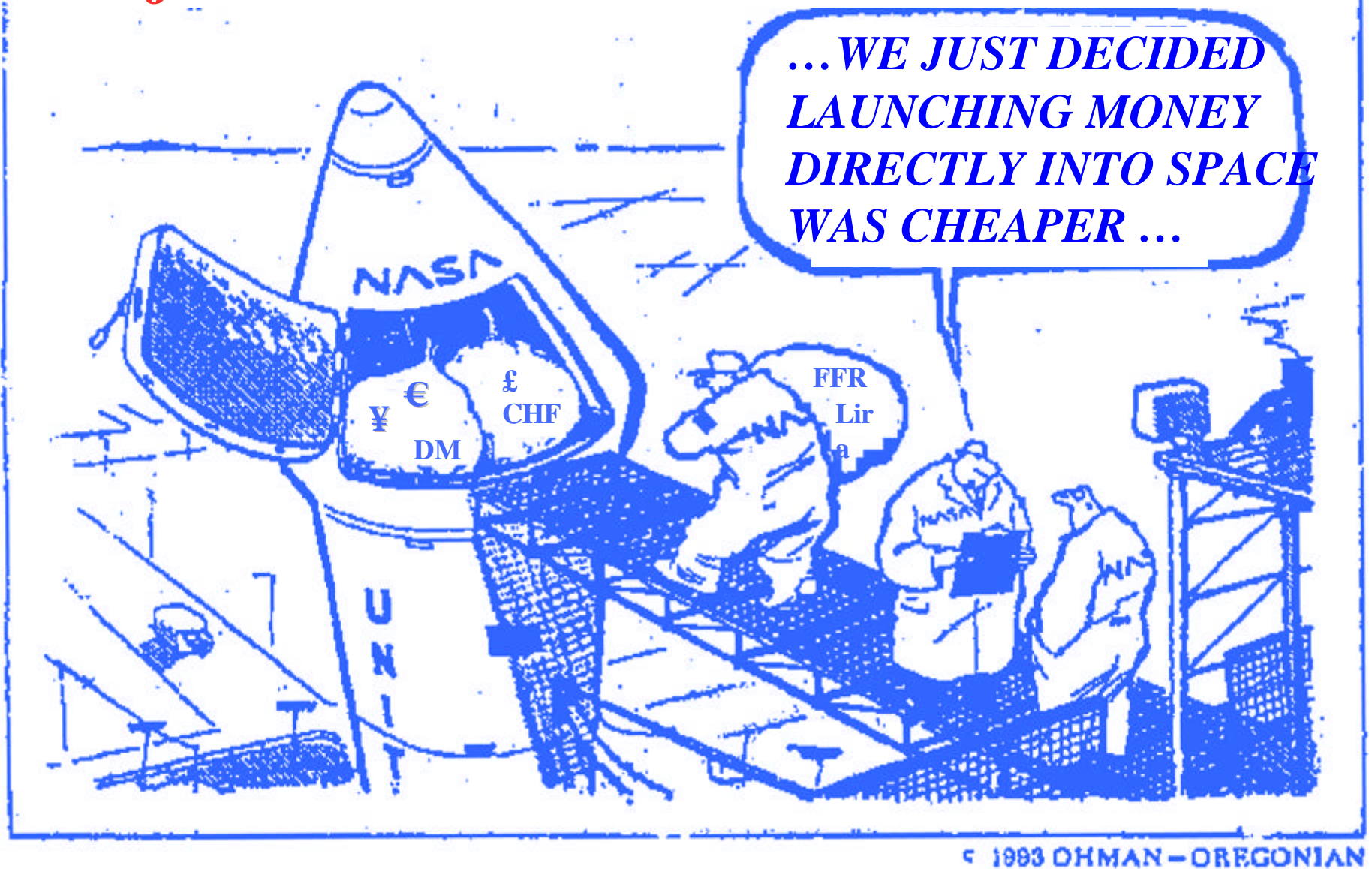
by B. Blau, V.Choutko, S. Harrison, A. Herve, S. Horvarth, H. Hofer,
H.P. Marti, I. Vetlitskiy, et al, (CERN, ETH, MIT, SCL)

with AMS technology



Weight of complete system: 30.7 ton

Major International Commitments to NASA



Discoveries in Physics

Facility	Original purpose, Expert Opinion	Discovery with Precision Instrument
P.S. CERN (1960)	π N interactions	Neutral Currents \rightarrow Z, W
AGS Brookhaven (1960)	π N interactions	2 kinds of neutrinos, Time reversal non-symmetry, New form of matter (4 th Quark)
FNAL Batavia (1970)	Neutrino physics	5 th Quark, 6 th Quark
SLAC Spear (1970)	ep, QED	Partons, 4 th Quark, 3 rd electron
ISR CERN (1980)	PP	Increasing PP Cross section
PETRA Hamburg (1980)	6 th Quark	Gluon
Super Kamiokande (2000)	Proton decay	Neutrinos have mass
Hubble Space Telescope	Galactic survey	Curvature of the universe, dark energy

Exploring a new territory with a precision instrument is the key to discovery.

