

New Results from the Salt Phase at Sudbury Neutrino Observatory

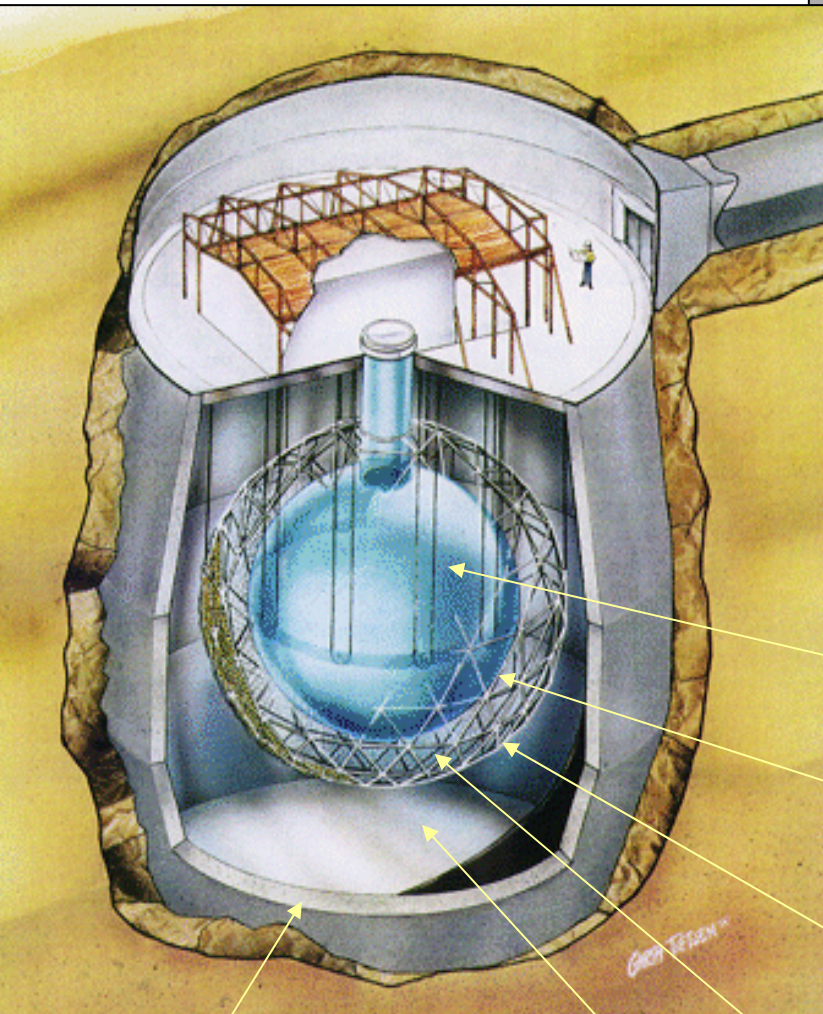
***Jeanne Wilson, University of Sussex,
on behalf of the
Sudbury Neutrino Observatory
(SNO) collaboration***

Whats New?



- First full **Spectral** analysis
 - Full evaluation of energy-dependent systematics
- Salt **Day-Night** asymmetry measurements
- Updated **Flux** measurements for increased statistics in Salt data set.

SNO



Urylon Liner and Radon Seal

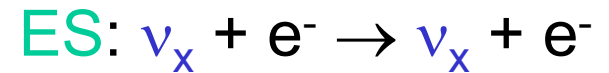
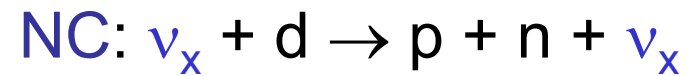
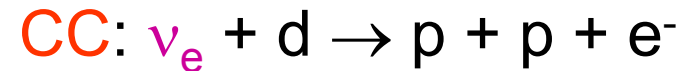
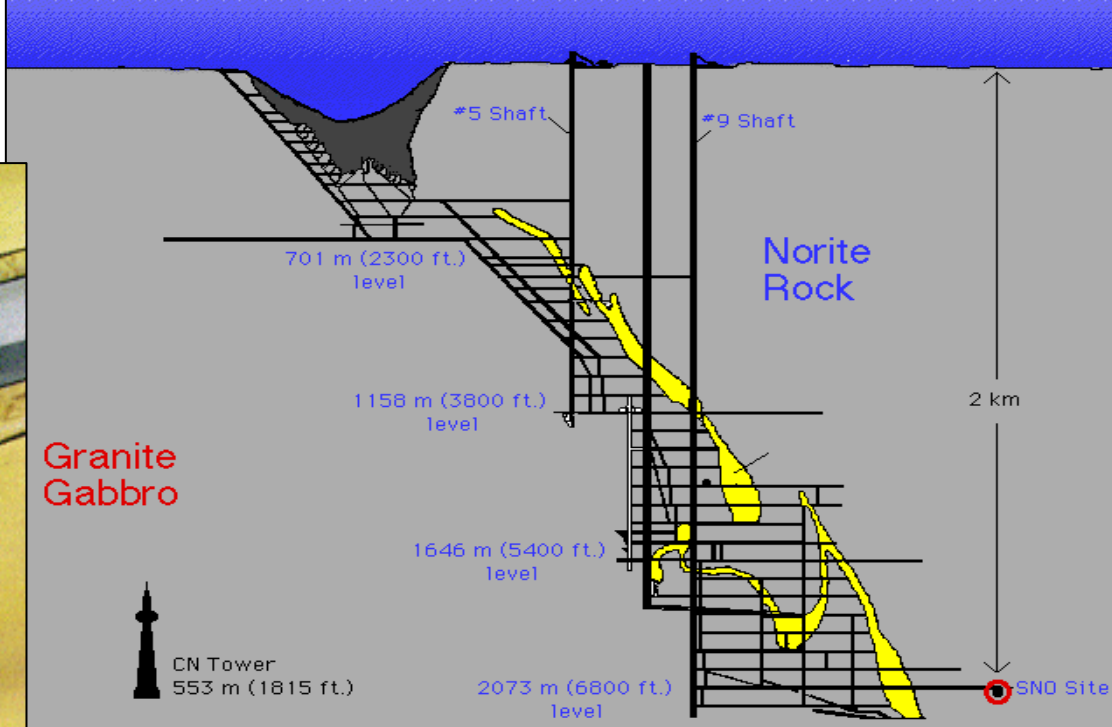
5300 tonnes Outer Shield H₂O

1700 tonnes Inner Shielding H₂O

1000 tonnes D₂O

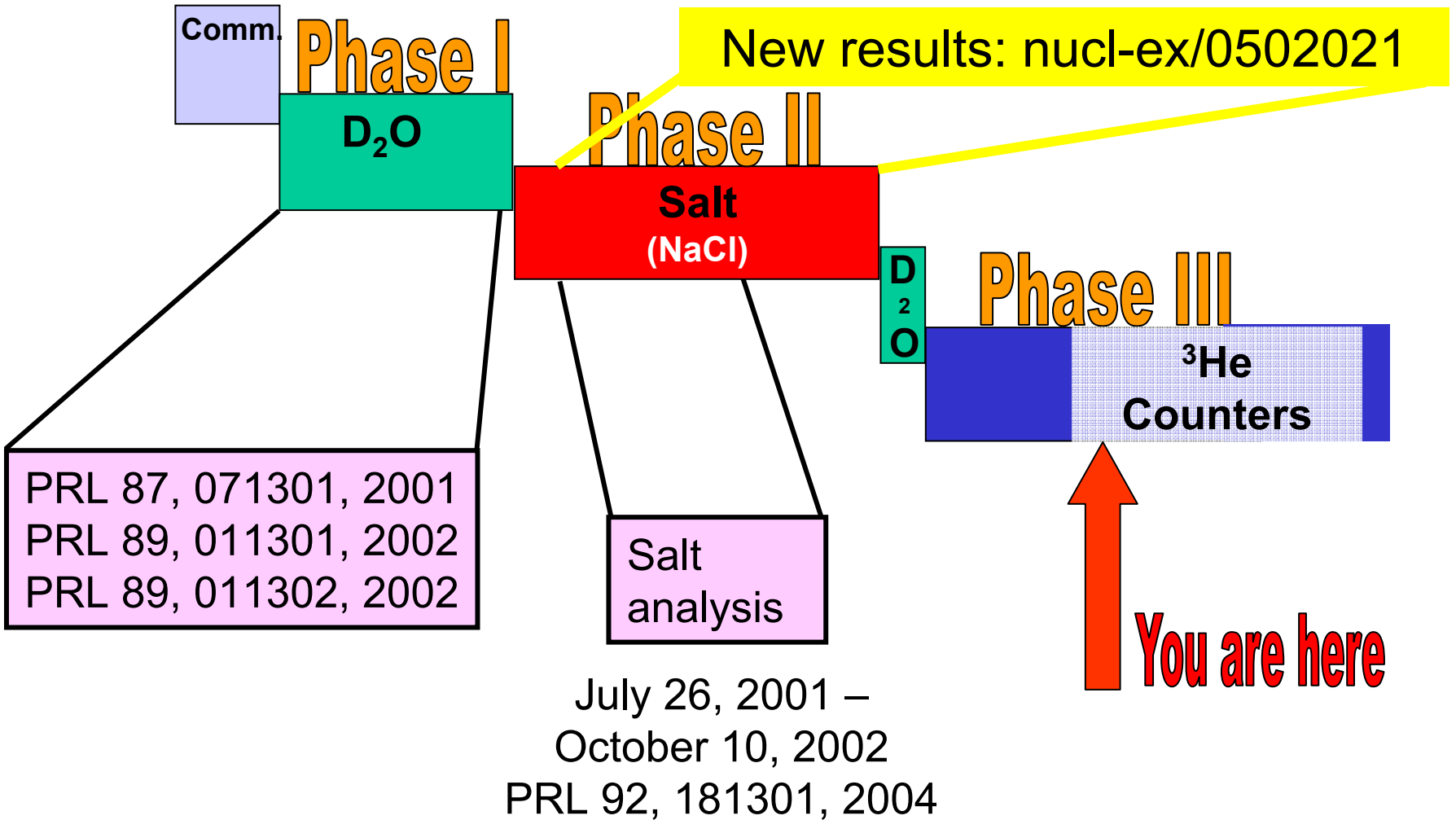
12m Diameter Acrylic Vessel

Support Structure for 9500 PMTs, 60% coverage



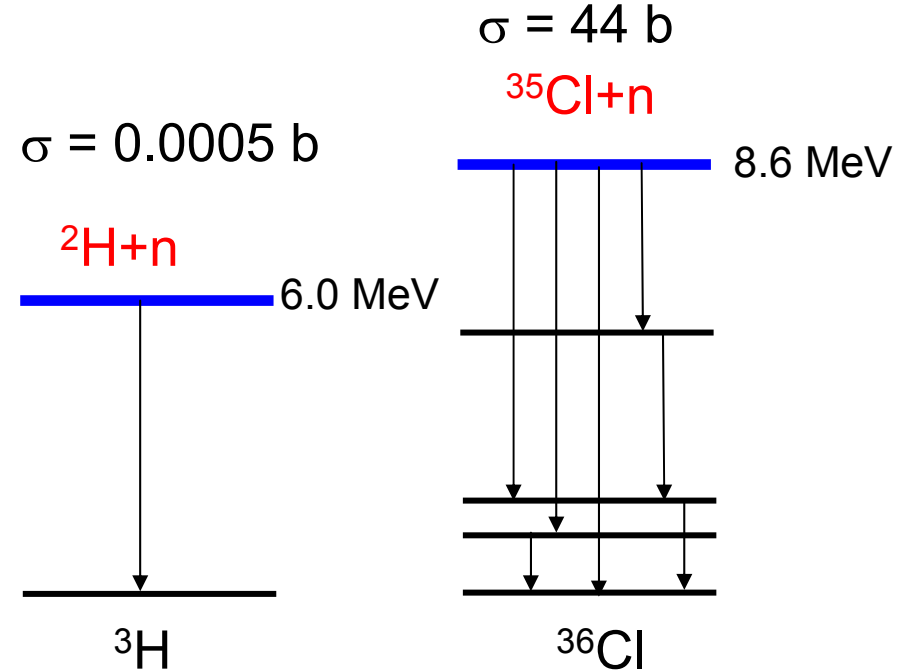
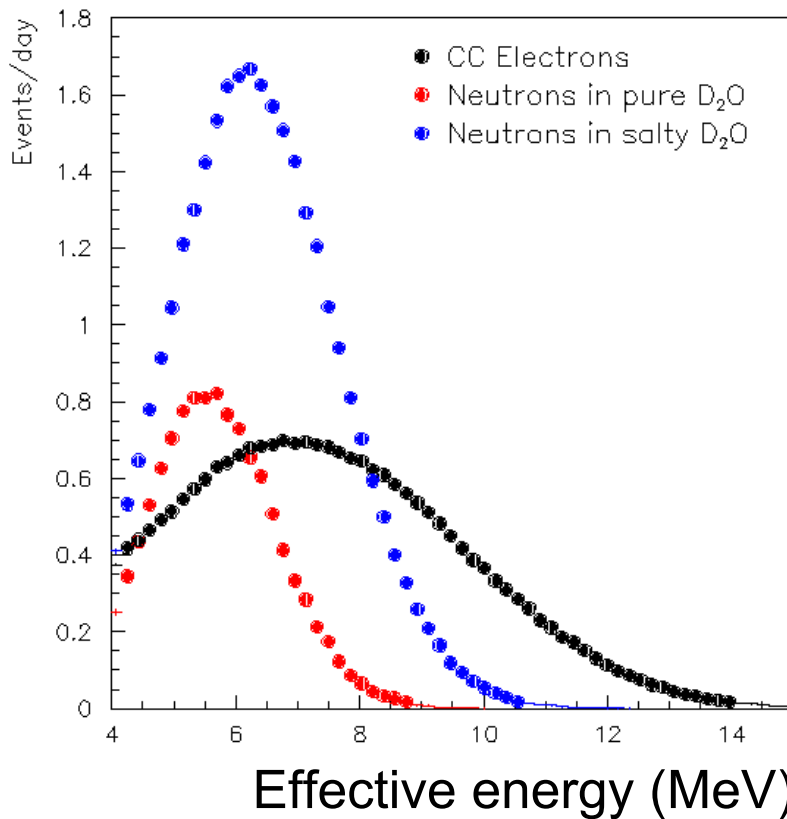
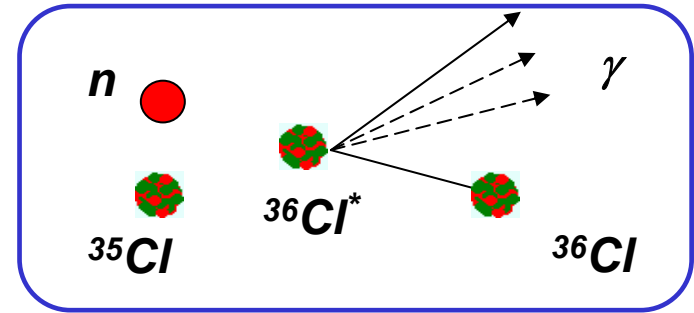
SNO Program

1998 1999 2000 2001 2002 2003 2004 2005 2006



Advantages of NaCl for Neutron Detection

- Higher capture cross section
- Higher energy release
- Many gammas



Event Isotropy

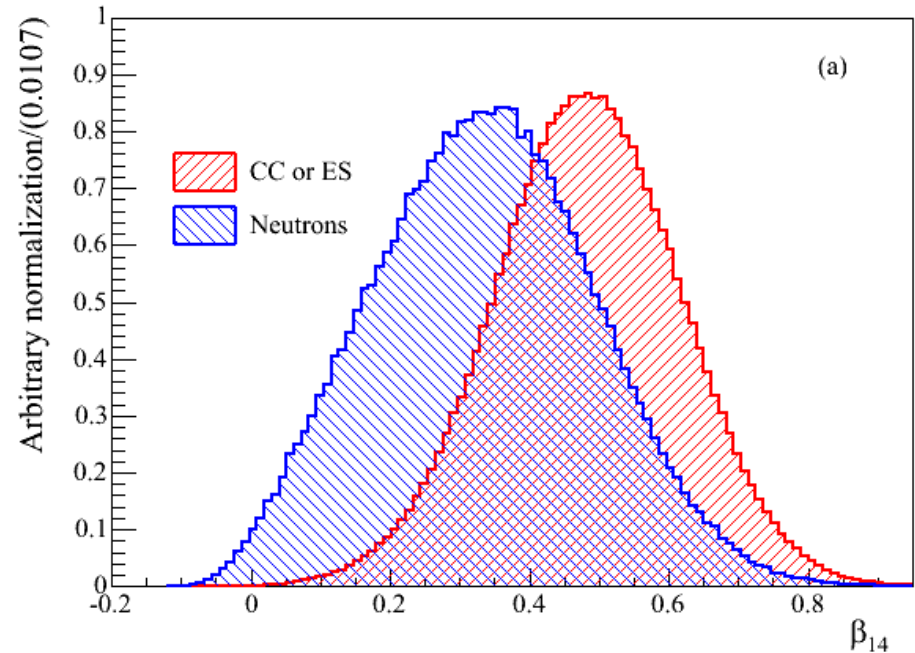
Define the parameters β_l as

$$\beta_l \approx \left\langle P_l(\cos \theta_{ij}) \right\rangle_{i \neq j}$$

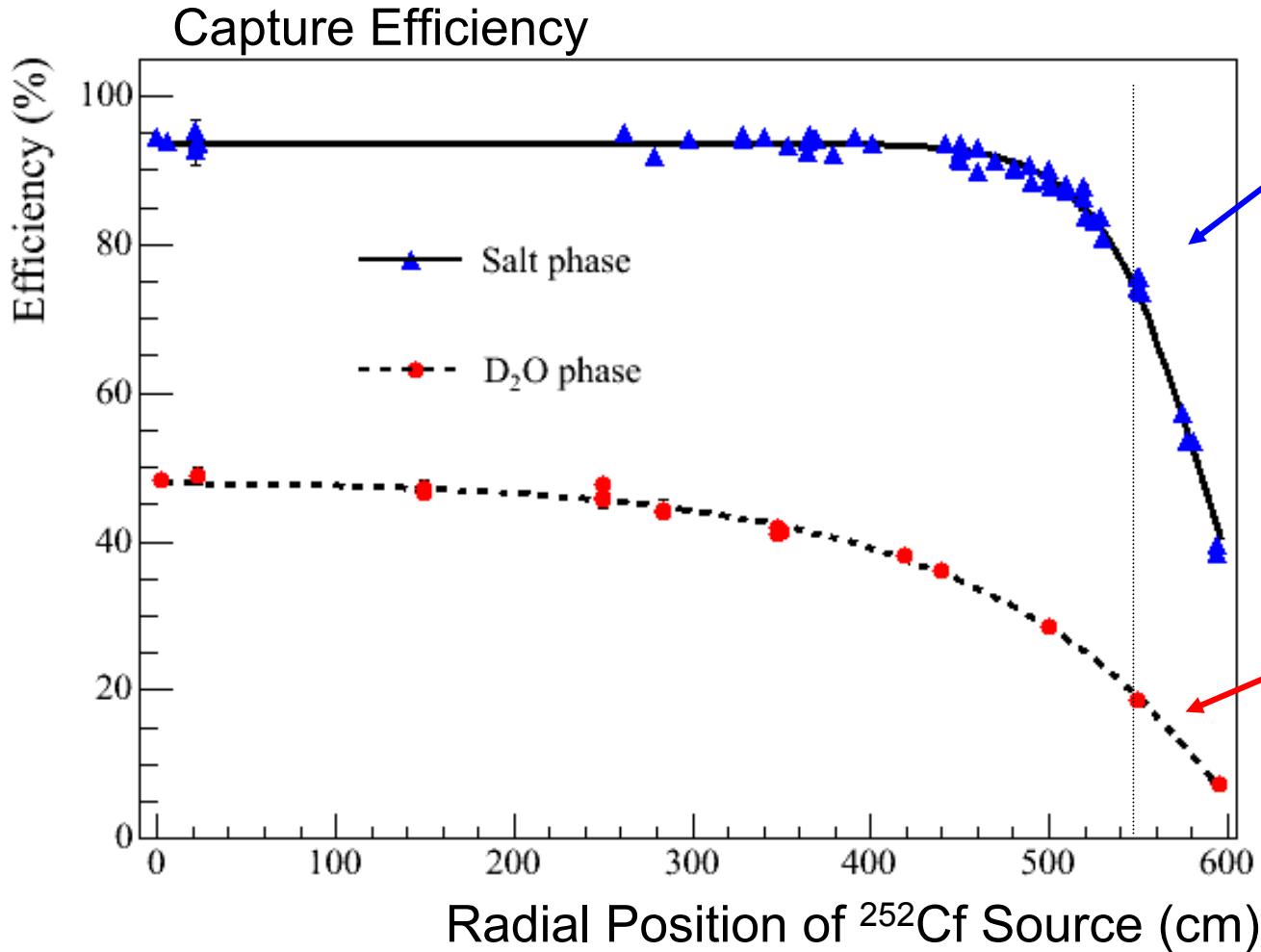
where P_l is the l^{th} order Legendre Polynomial.

θ_{ij} = angle between pair of hits subtended at event vertex.

$\beta_{14} = \beta_1 + 4\beta_4$
is found to give good
CC/NC separation



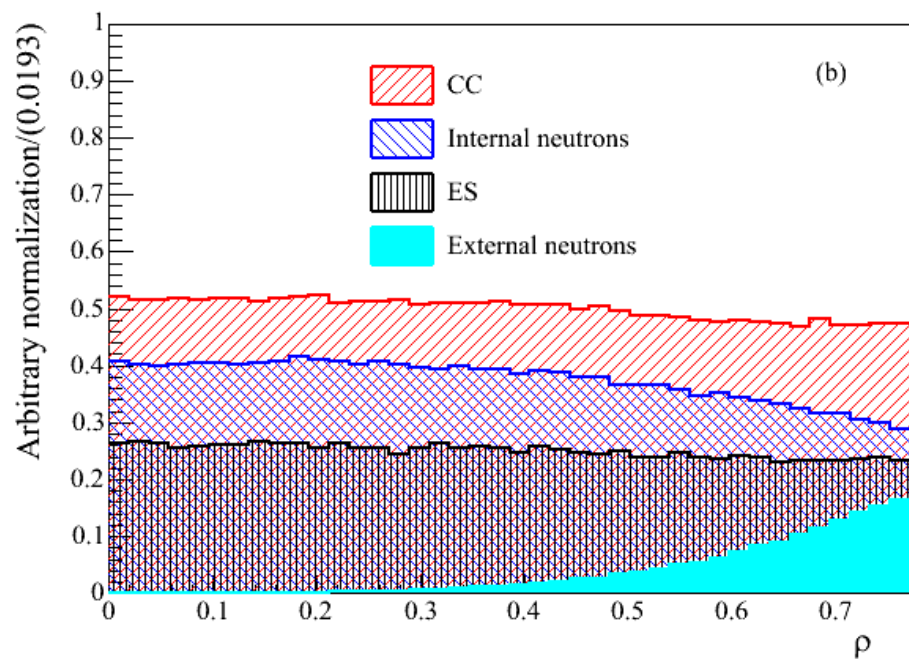
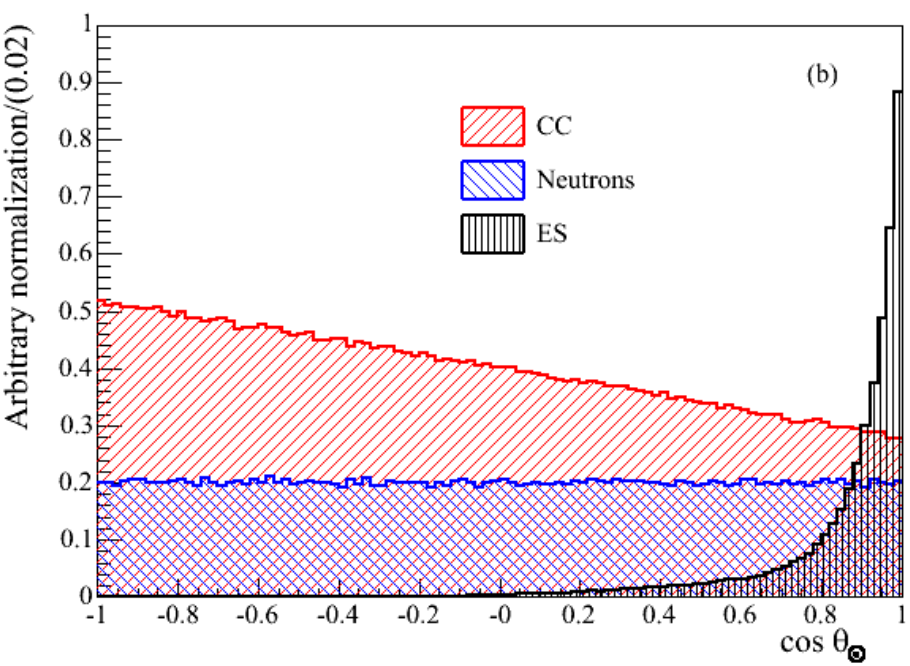
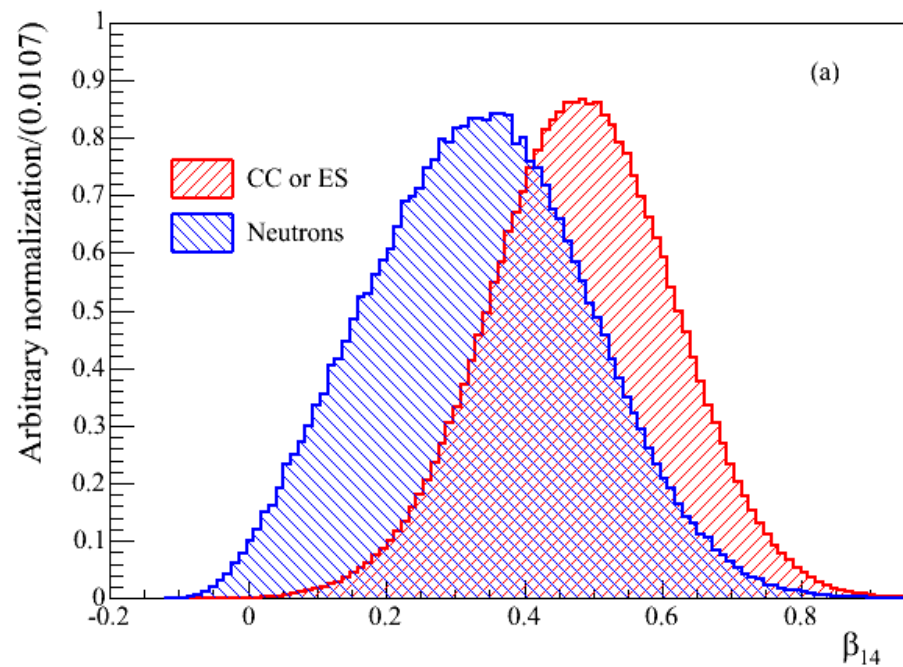
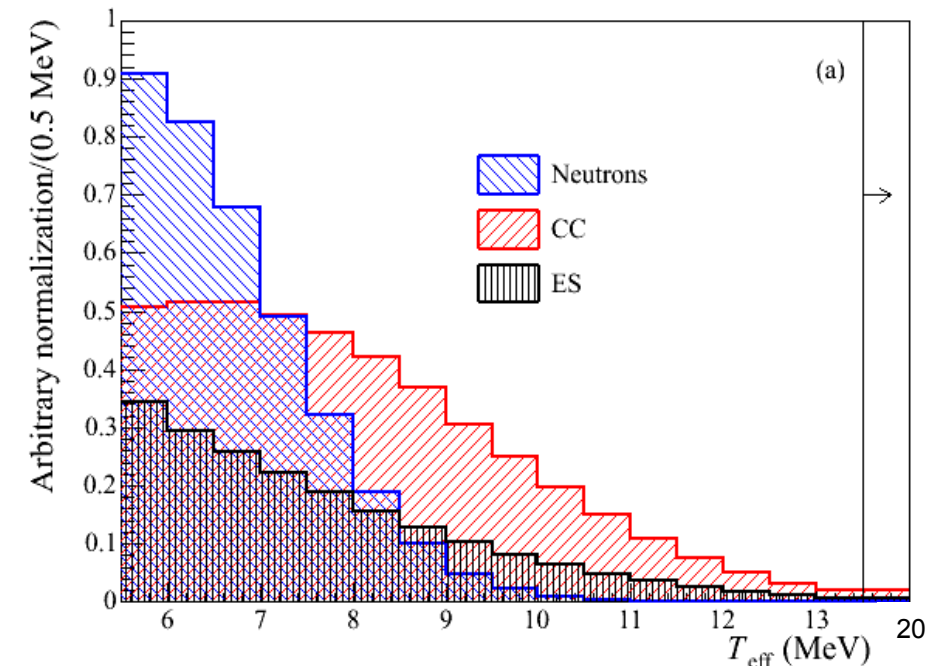
Neutron Capture Efficiency in SNO



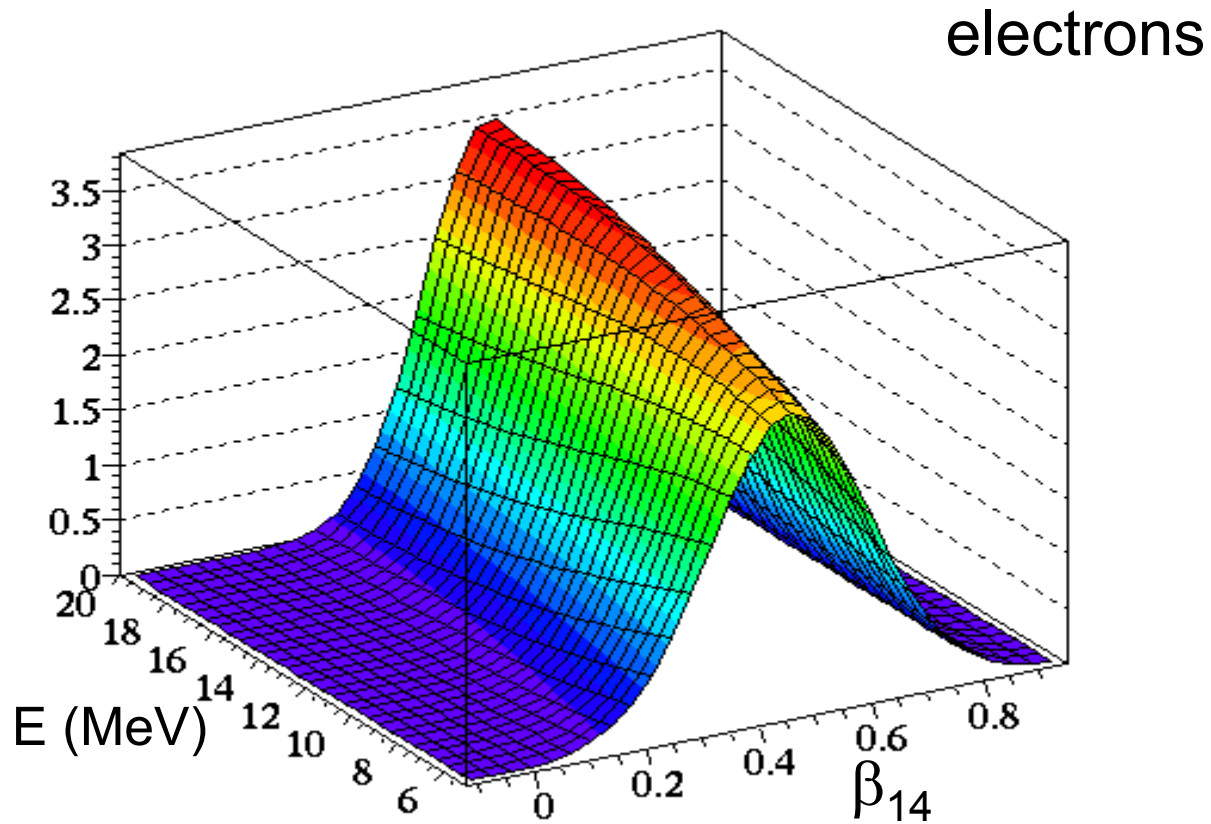
Detection Efficiency

$^{35}\text{Cl}(n,\gamma)^{36}\text{Cl}$
Average Eff. =
0.407
 $T_{\text{eff}} \geq 5.5 \text{ MeV}$ and
 $R \leq 550 \text{ cm}$

$^2\text{H}(n,\gamma)^3\text{H}$
Average Eff. =
0.144
 $T_{\text{eff}} \geq 5.0 \text{ MeV}$ and
 $R \leq 550 \text{ cm}$



Correlated Parameters



Correlated Parameters

- Ideal scenario – 4 dimensional PDF
 - Statistical limitations
- Signal extraction procedure tested on 100 simulated data sets.

- Small bias ($\sim 0.8\%$ on CC) seen using:

$$P(T_{\text{eff}}, \beta_{14}, R^3, \cos\theta) = P(T_{\text{eff}}, \beta_{14}) \cdot P(\cos\theta) \cdot P(R^3)$$

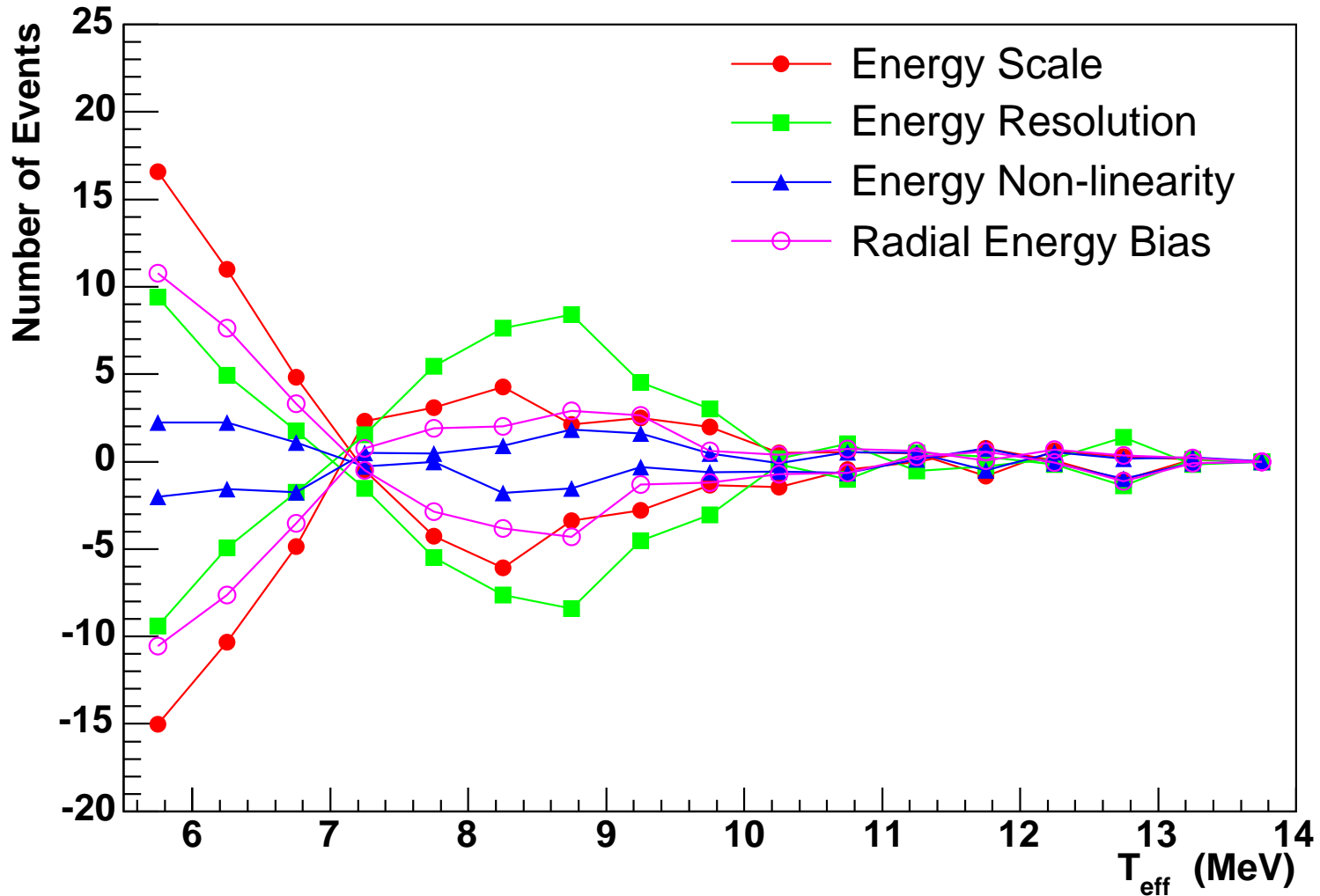
- Bias becomes insignificant using:

$$P(T_{\text{eff}}, \beta_{14}, R^3, \cos\theta) = P(T_{\text{eff}}, \beta_{14}, R^3) \cdot P(\cos\theta / T_{\text{eff}}, R^3)$$

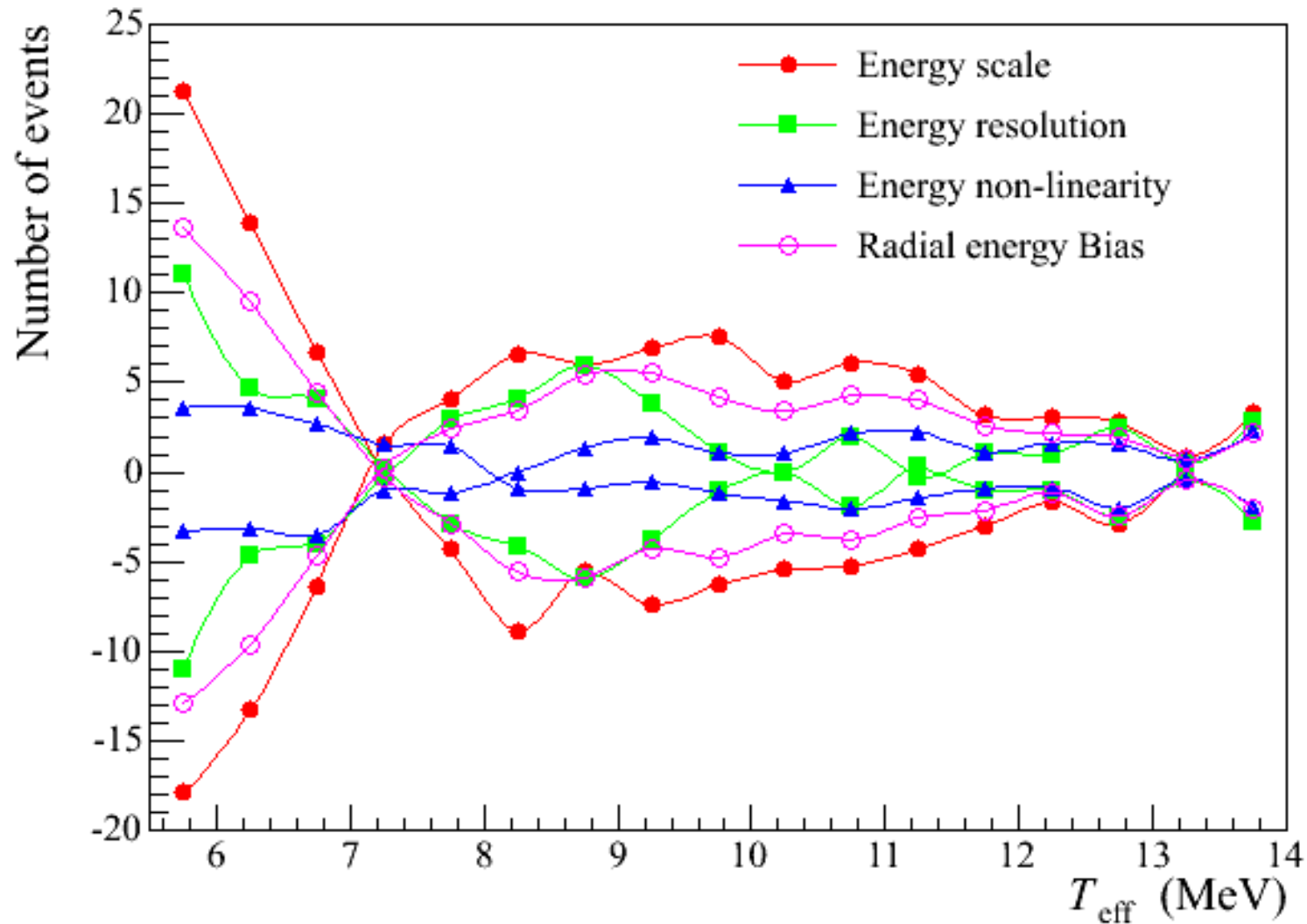
Major Systematic Uncertainties


- A number of systematics re-evaluated for possible differential effects in energy.
 - Energy systematics
 - β_{14} systematics
 - Fiducial volume
 - Selection efficiency
- High(er) energy calibration data used to parameterise non-linear contributions to systematic uncertainty.
 - ^8Li , ^{252}Cf in addition to ^{16}N

Differential Systematic Effects



Differential Systematic Effects

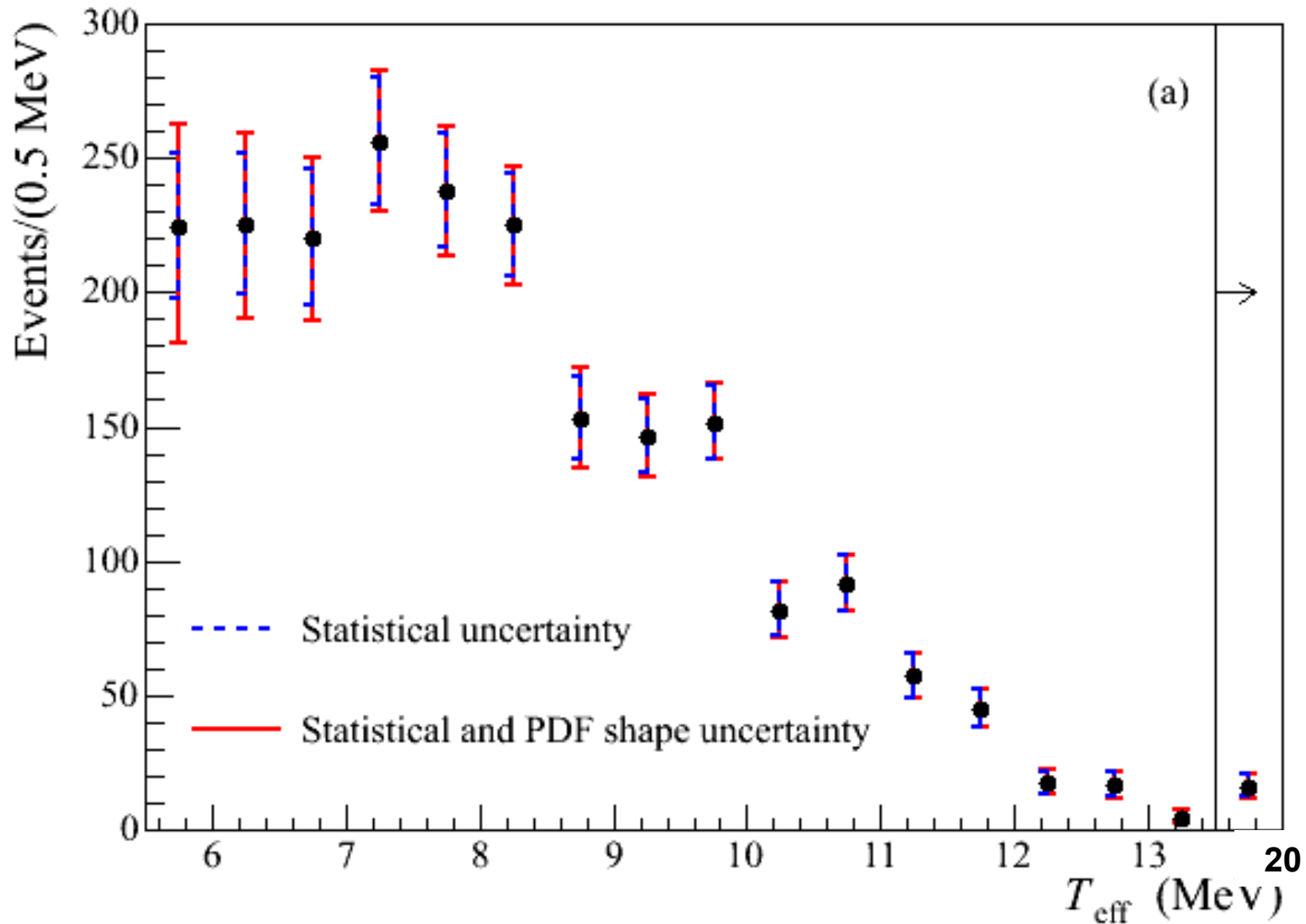


A photograph of the Sudbury Neutrino Observatory building, a large industrial-style structure with light blue corrugated metal siding and a grey stone base. The building features a long row of windows on the left, a central entrance, and a large white garage door on the right. The name "Sudbury Neutrino Observatory" is mounted on the upper right side of the facade. The scene is captured at dusk or dawn, with a cloudy sky and some interior lights visible through the windows.

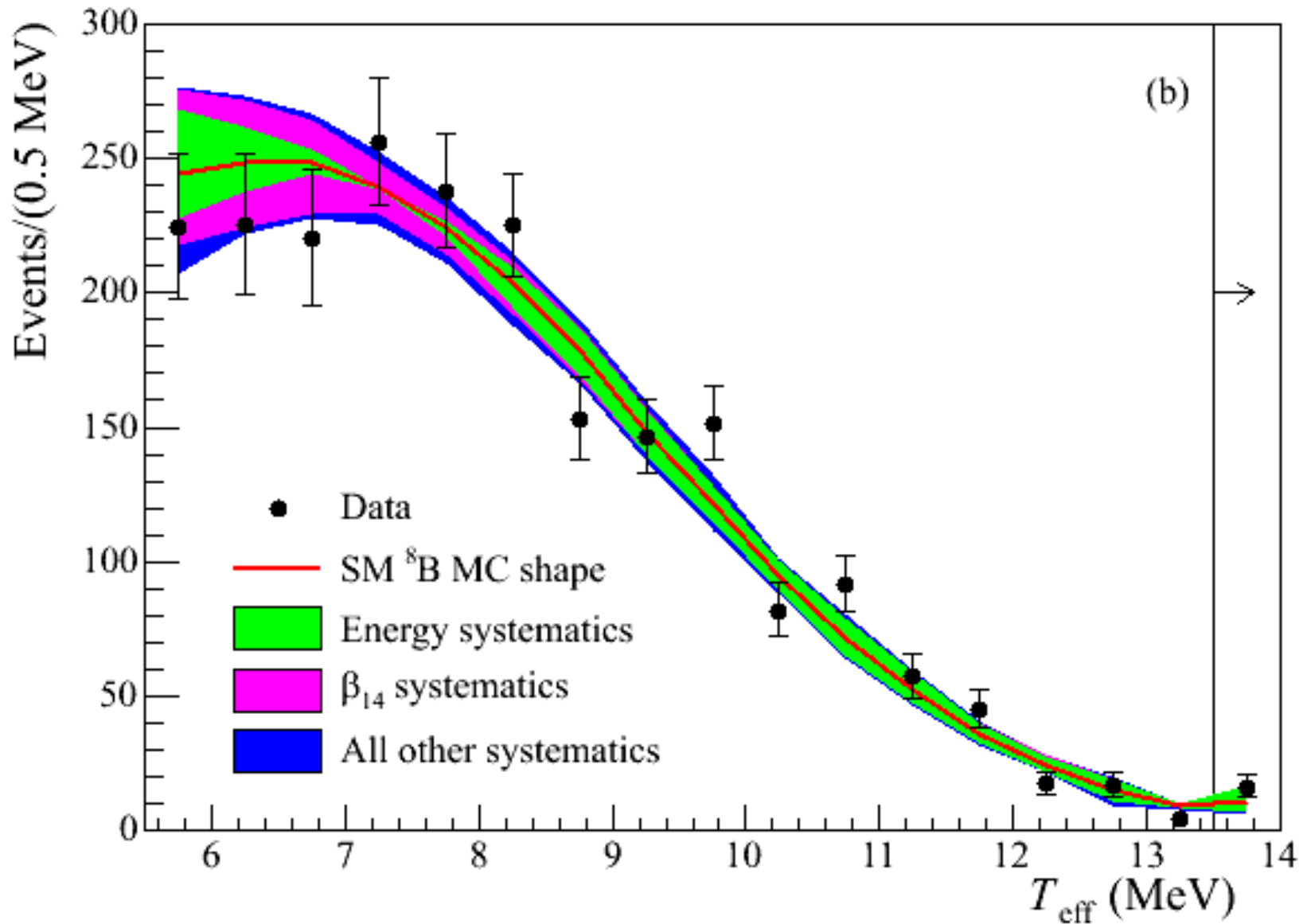
Sudbury
Neutrino
Observatory

Spectrum

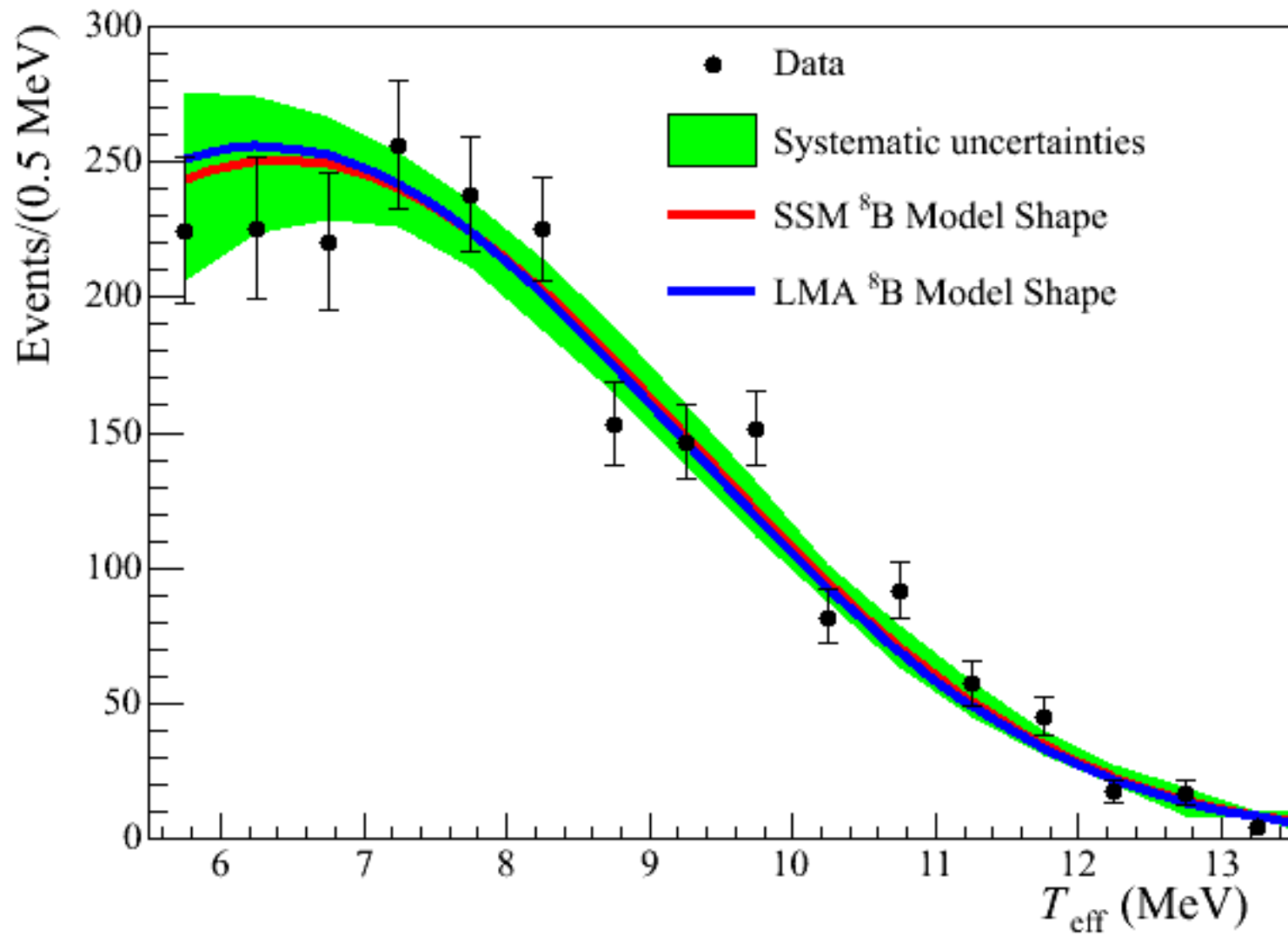
Charged Current Spectrum



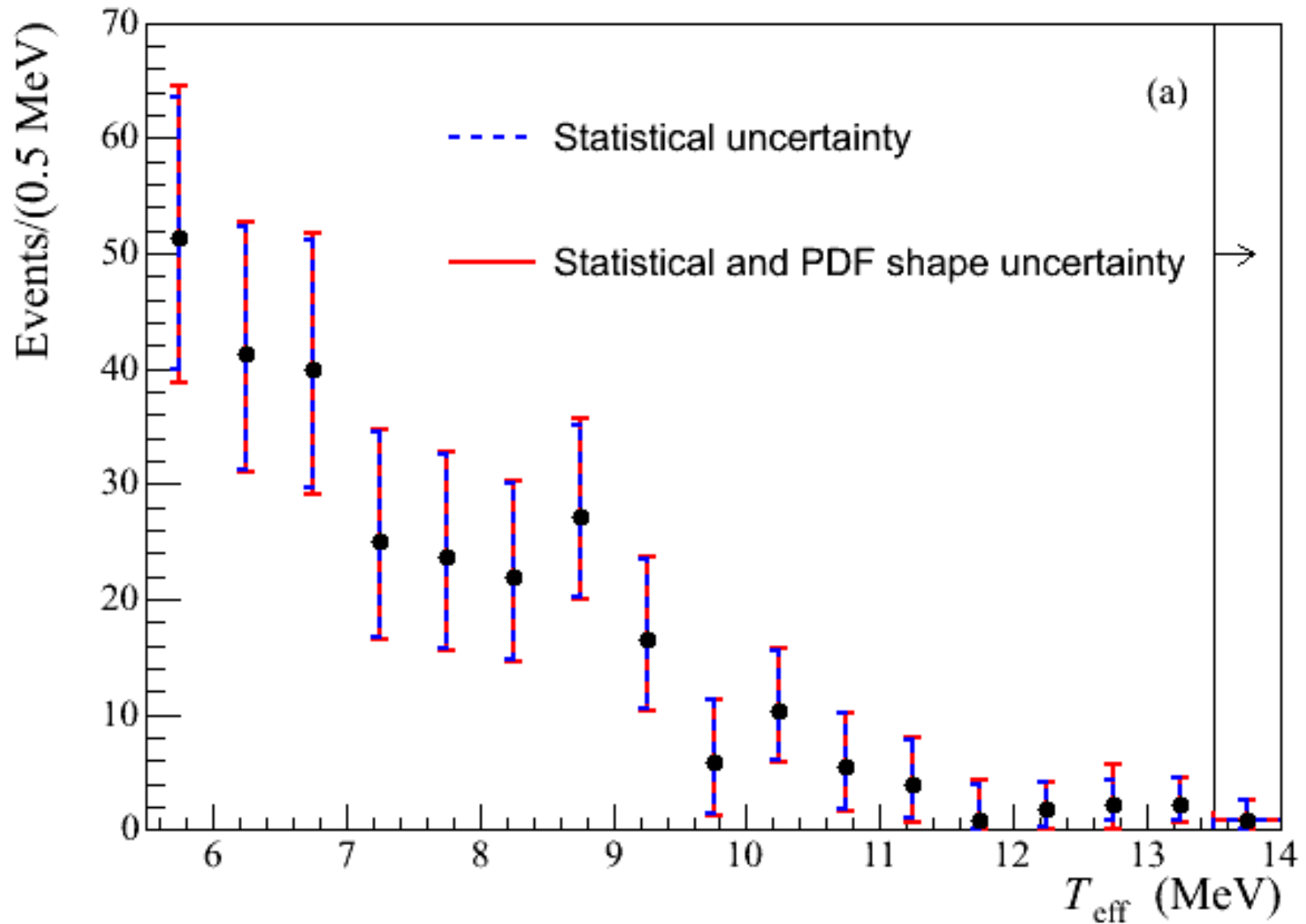
Charged Current Spectrum



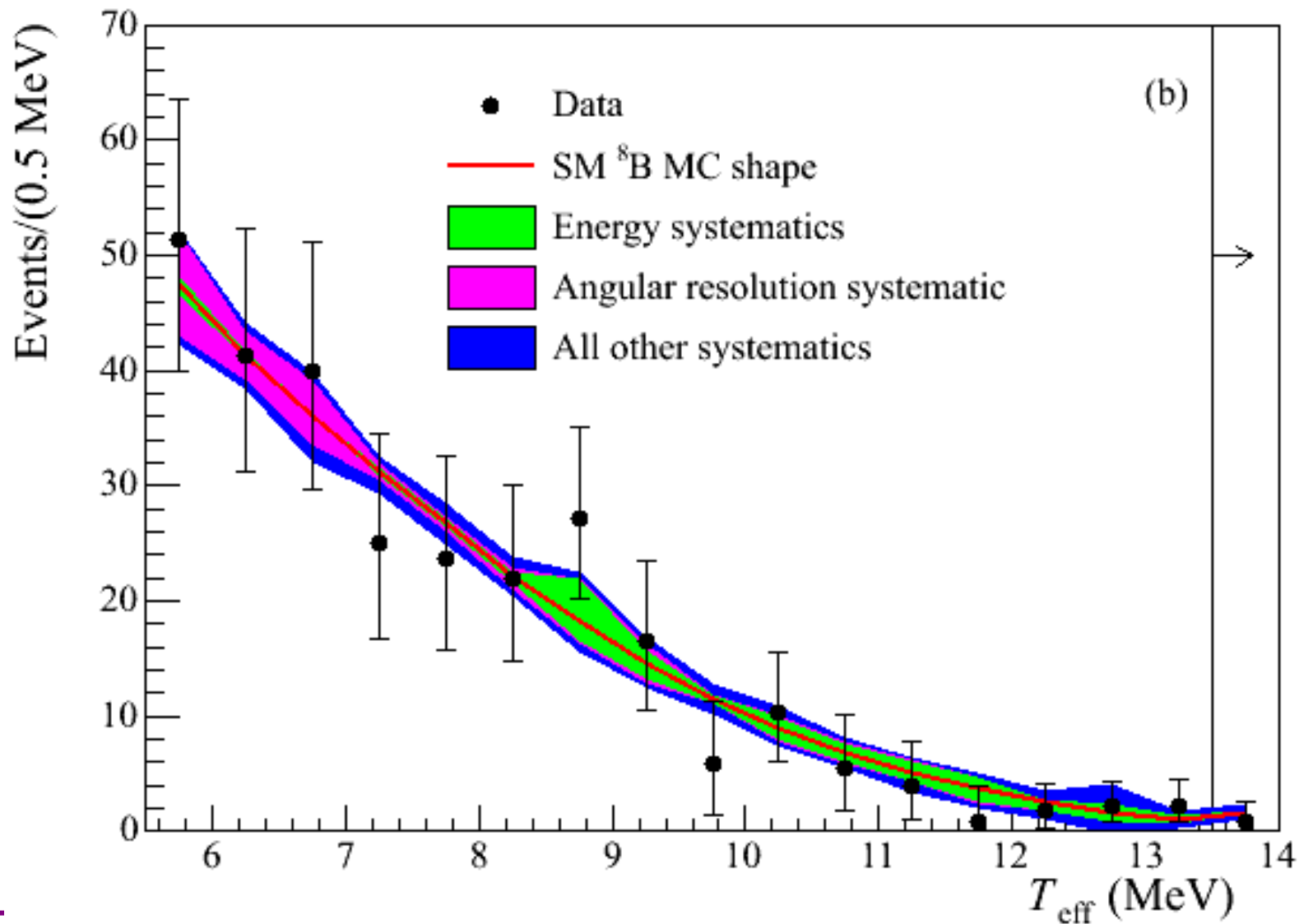
Predicted LMA spectral distortion



Elastic Scattering Spectrum



Elastic Scattering Spectrum





Day Night

Day Night Asymmetry

CC,ES asymmetry

ν_e regeneration inside the Earth

NC asymmetry

Admixture of sterile neutrinos or unexpected matter interactions inside the Earth

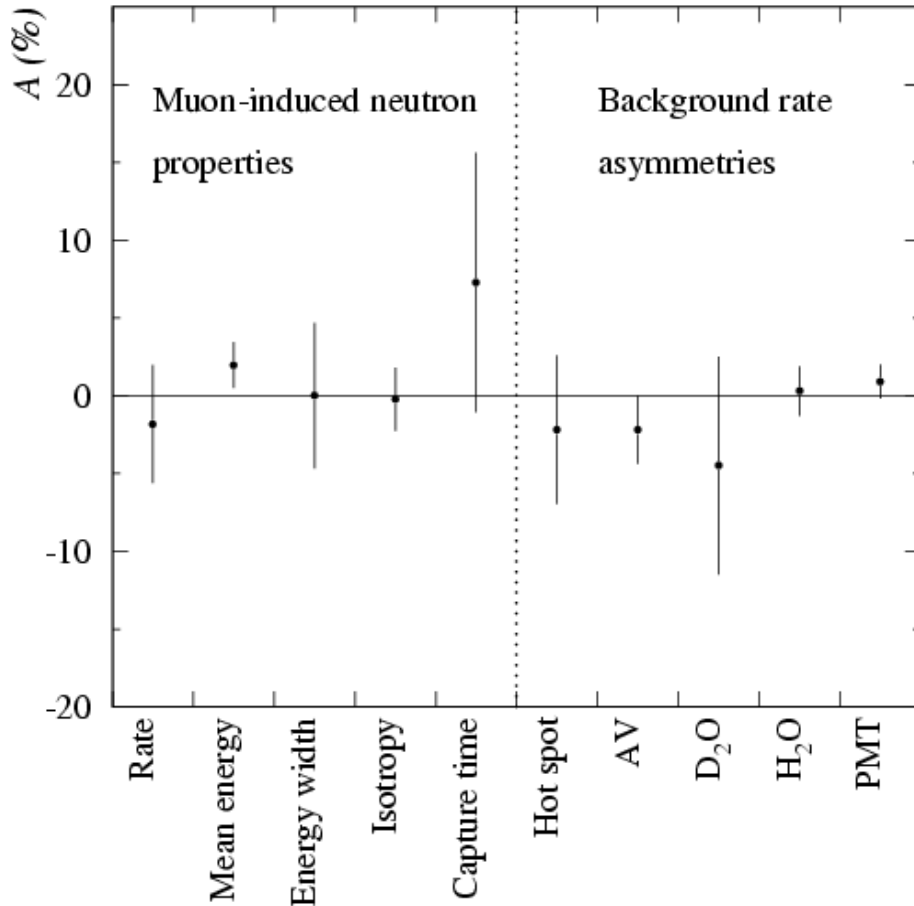
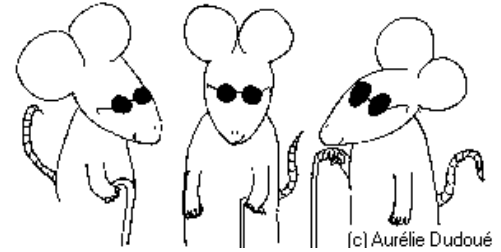
LMA \rightarrow small A_{CC} , A_{ES}
 $A_{NC} = 0$



$$A = \frac{2(\Phi_N - \Phi_D)}{(\Phi_N + \Phi_D)}$$

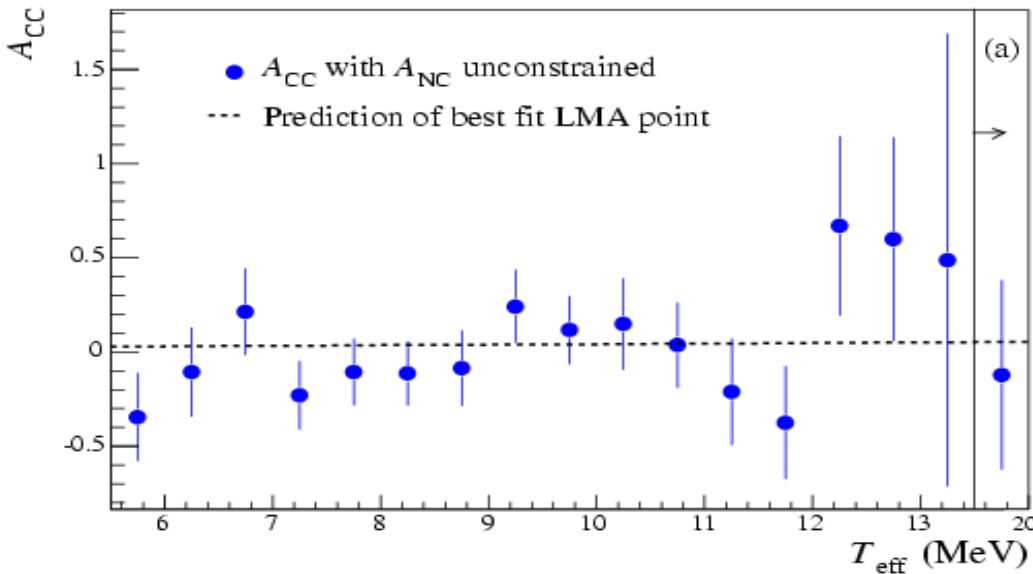
Day Night Asymmetry

- Blind analysis to avoid statistical bias (analysis developed on 20% of data)



Constant background sources used to investigate diurnal systematic effects.

Shape Unconstrained Analysis

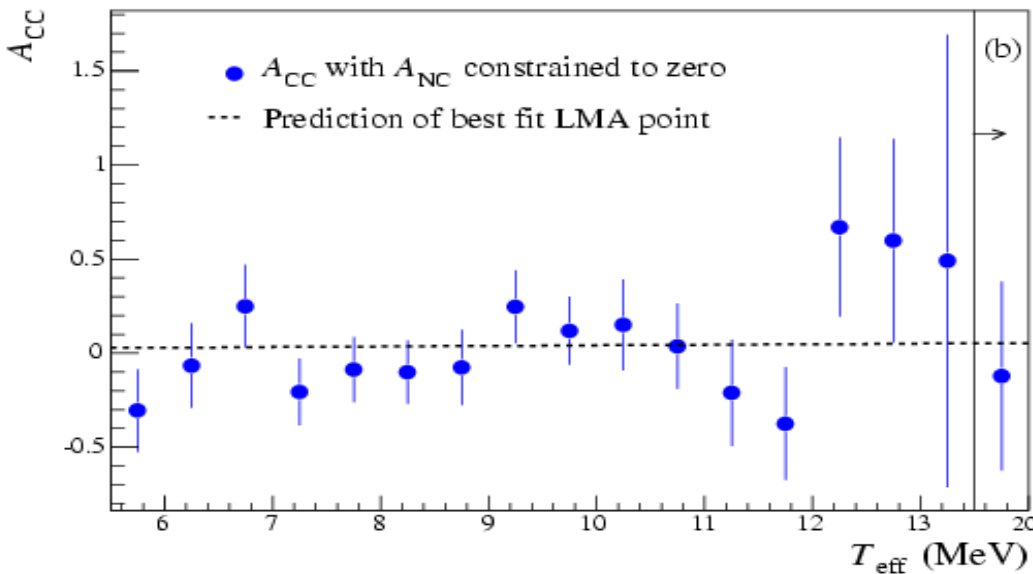


No constraint on A_{NC}

$$A_{CC} = -0.056 \pm 0.074 \pm 0.051$$

$$A_{NC} = 0.042 \pm 0.086 \pm 0.067$$

$$A_{ES} = 0.146 \pm 0.198 \pm 0.032$$



$$A_{CC} = -0.037 \pm 0.063 \pm 0.032$$

$$A_{ES} = 0.153 \pm 0.198 \pm 0.030$$

$$A_{NC} = 0$$



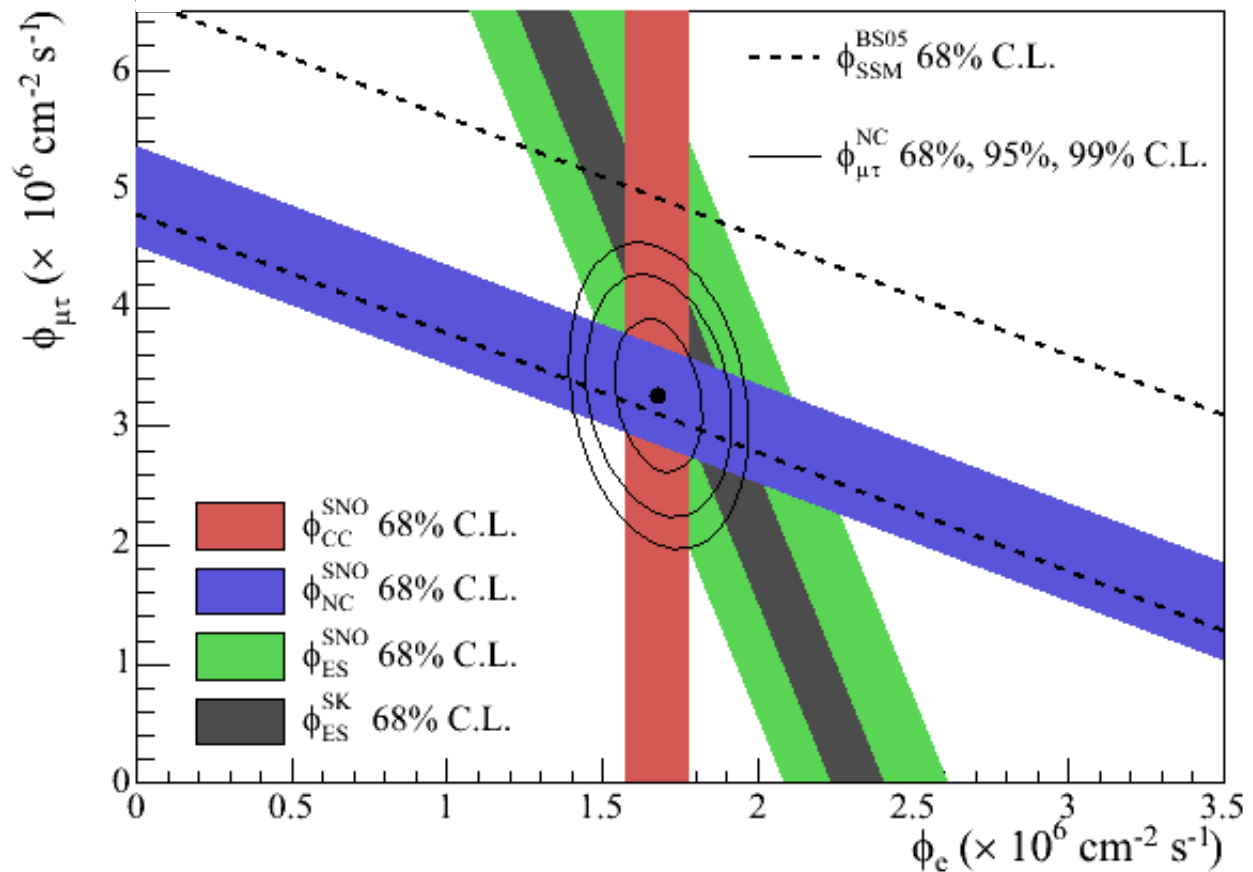
Updated Flux

Energy-Unconstrained Fluxes

$$\Phi_{CC} = 1.68 \pm 0.06 \pm {}^{0.08}_{0.09}$$

$$\Phi_{ES} = 2.35 \pm 0.23 \pm {}^{0.15}_{0.14} \quad \times 10^6 \text{ cm}^{-2} \text{ s}^{-1}$$

$$\Phi_{NC} = 4.94 \pm 0.21 \pm {}^{0.38}_{0.34}$$



SNO flux results - Summary

Phase I (306 days)

constrained fit



Phase I (306 days)

constrained fit



Phase II (391 days)

constrained fit



Phase II (391 days)

constrained fit



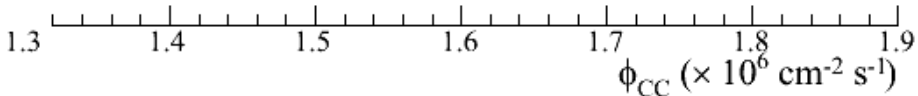
Phase II (391 days)

unconstrained fit

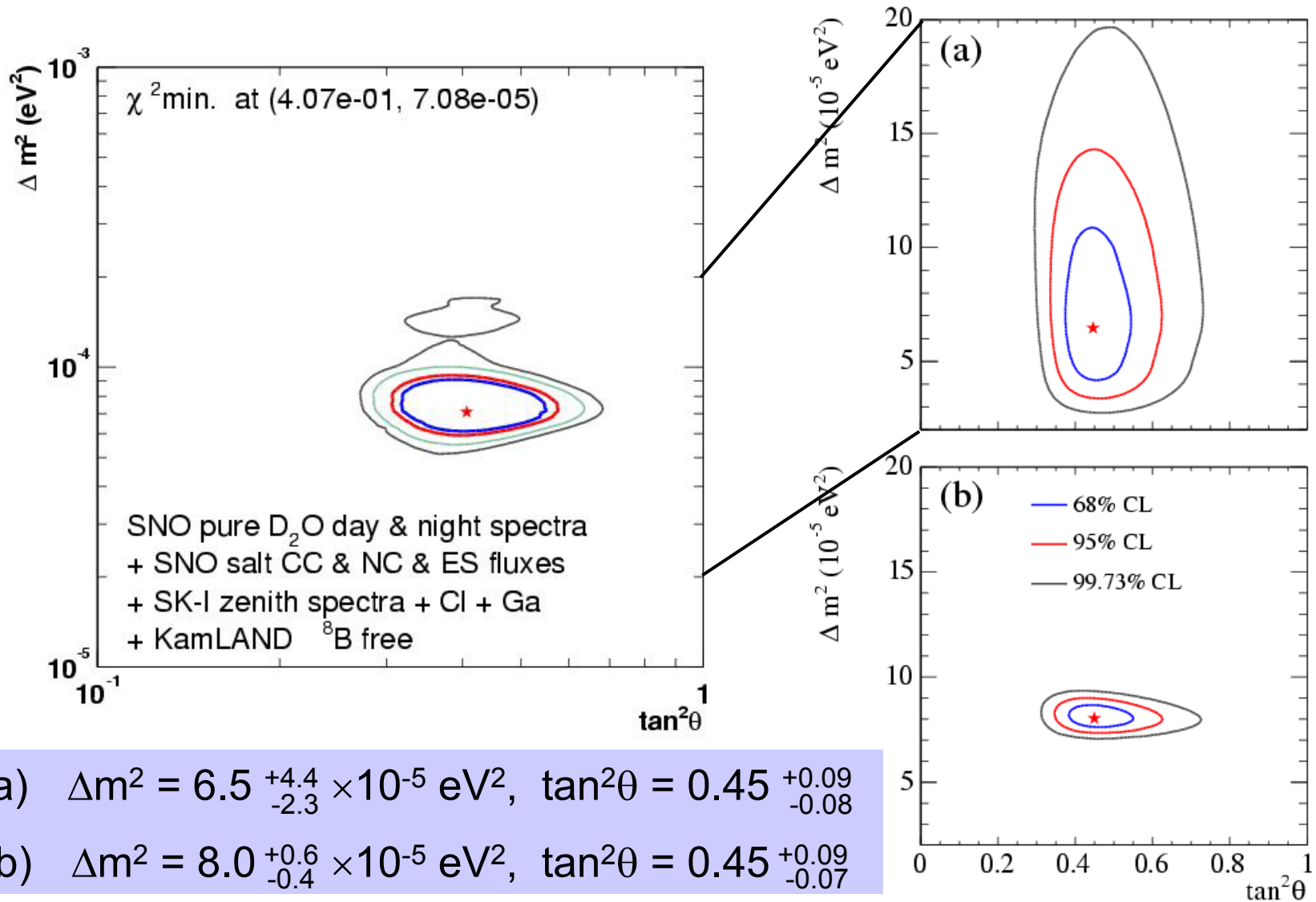


Phase II (391 days)

unconstrained fit



MSW interpretation of Results





Current Status

Where next?

- Lower energy threshold
- Combined analysis of the full salt and D₂O data sets.
- NCD phase

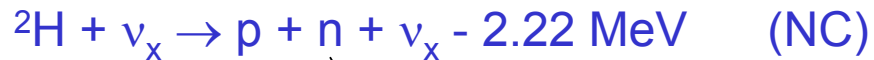
SNO Phase III (NCD Phase)

- ^3He Proportional Counters (“NC Detectors”) now installed
- Production data taking underway

40 Strings on 1-m grid

440 m total active length

Detection Principle

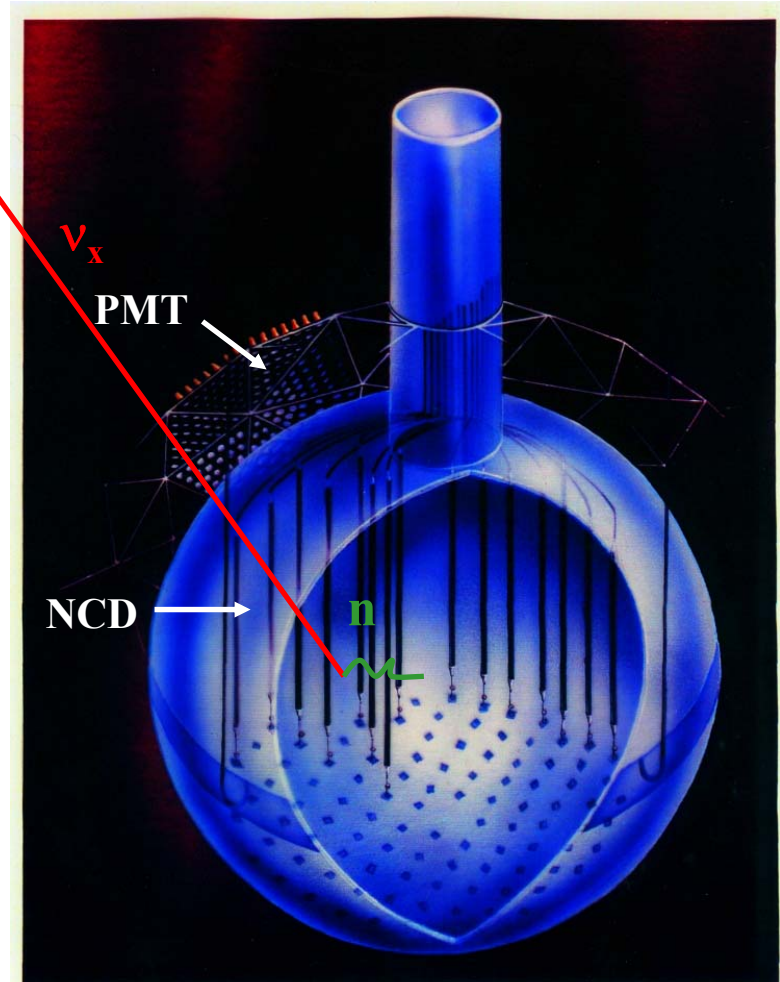


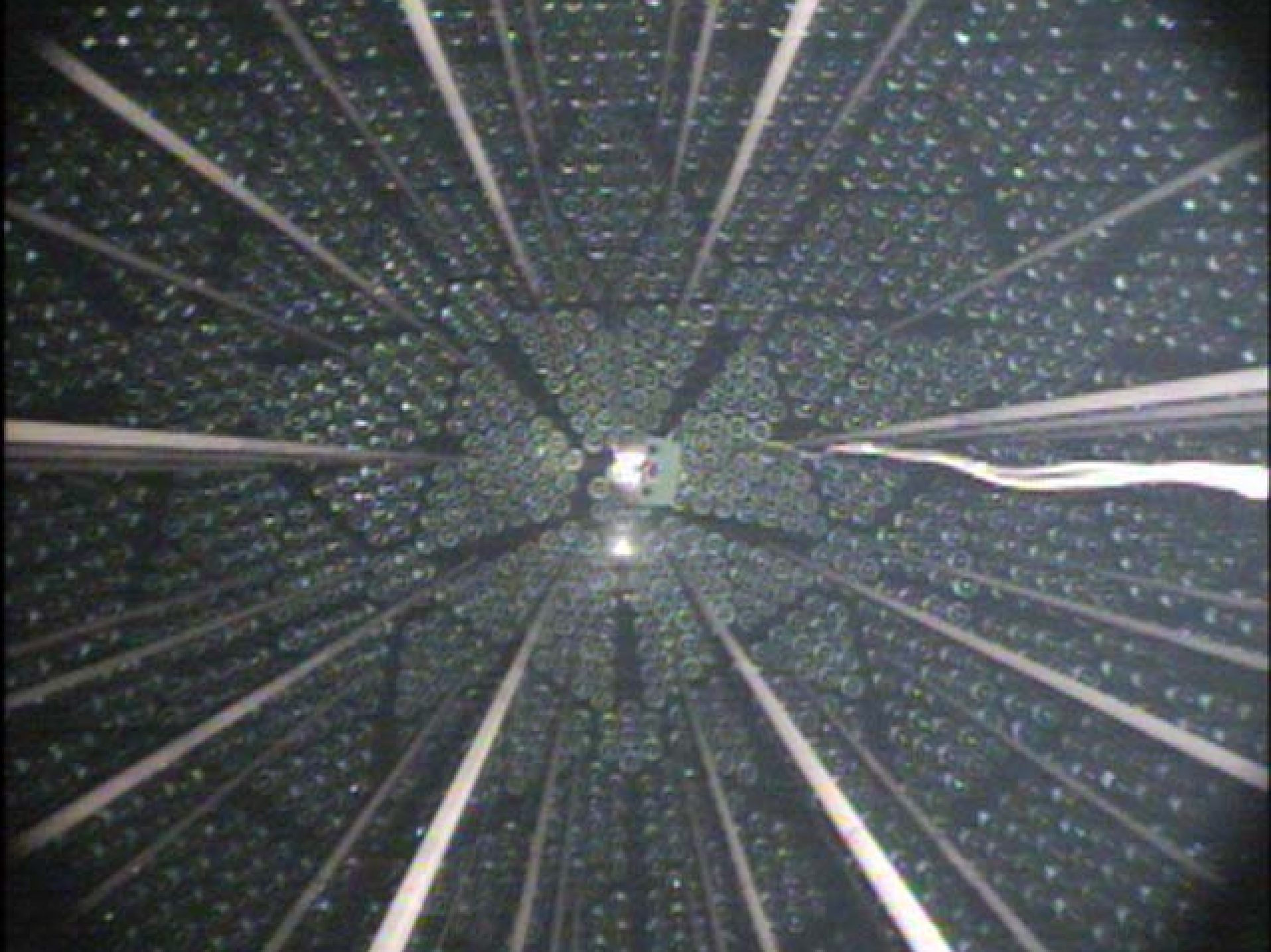
Physics Motivation

Event-by-event separation.

Measure NC and CC in separate data streams.

Different systematic uncertainties than neutron capture on NaCl.







Summary

- 4722 candidate neutrinos in 391 days data
- New salt results confirm and improve previous SNO data
- First full CC spectrum presented
 - consistent with best fit LMA
- First salt day:night measurements
 - A_{cc} consistent with zero and LMA prediction
- Best fit mixing parameters

$$\Delta m^2 = 8.0^{+0.6}_{-0.4} \times 10^{-5} \text{ eV}^2, \quad \tan^2 \theta = 0.45^{+0.09}_{-0.07}$$