

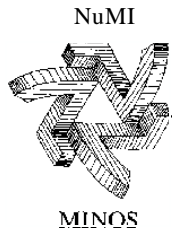
# MINOS Update

Robert C. Webb  
La Thuile 2002  
March 4, 2002

## NuMI/MINOS Update

Robert C. Webb  
Physics Department  
Texas A&M University  
March 4, 2002

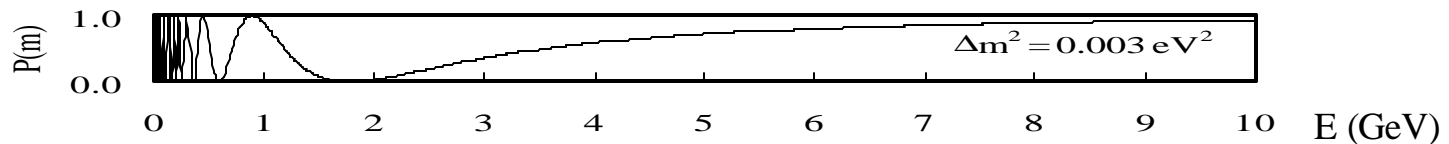
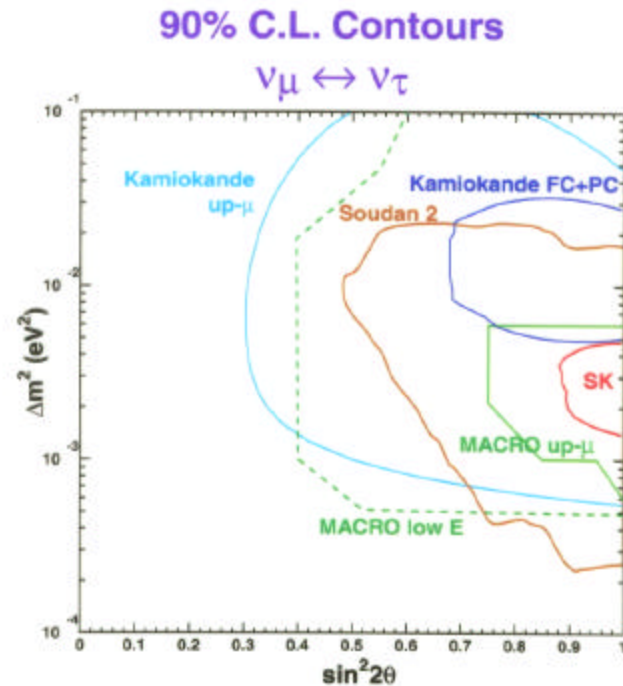
- Introduction
- Neutrino Oscillation Evidence
- Overview of NuMI / MINOS
- Physics Reach of MINOS
- Status of
  - NuMI
  - MINOS
    - Far Detector
    - Cal Det
    - Near Detector
- Summary

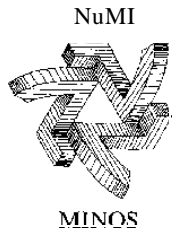


# MINOS Physics Goals

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- A next generation long baseline neutrino accelerator experiment
  - « Explore the region of oscillation parameter space indicated by the atmospheric neutrino results
- “A hotter beam over a longer baseline”
  - « Match flux and efficiency to the region of the first oscillation maximum
  - « Check the oscillation interpretation through spectral measurements
  - « Obtain improved parameter measurements
  - « Search for sub-dominant modes



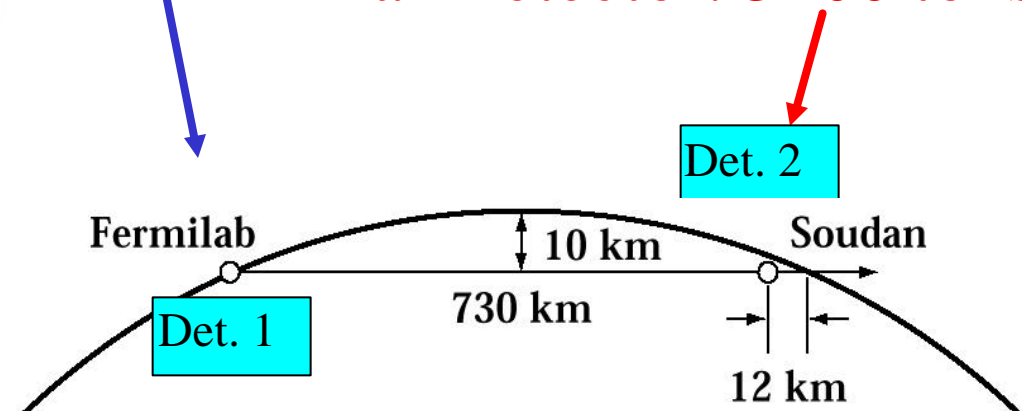


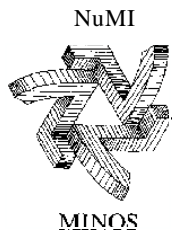
# MINOS Experiment

Two Detector Neutrino  
Oscillation Experiment  
(Start 2004)

Near Detector: 980 tons

Far Detector: 5400 tons





# Producing the NuMI Beam at Fermilab's Main Injector

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$2.7 \times 10^{20}$  120 GeV Protons per year hit the target\*

$\pi^+$  are produced at wide range of angles

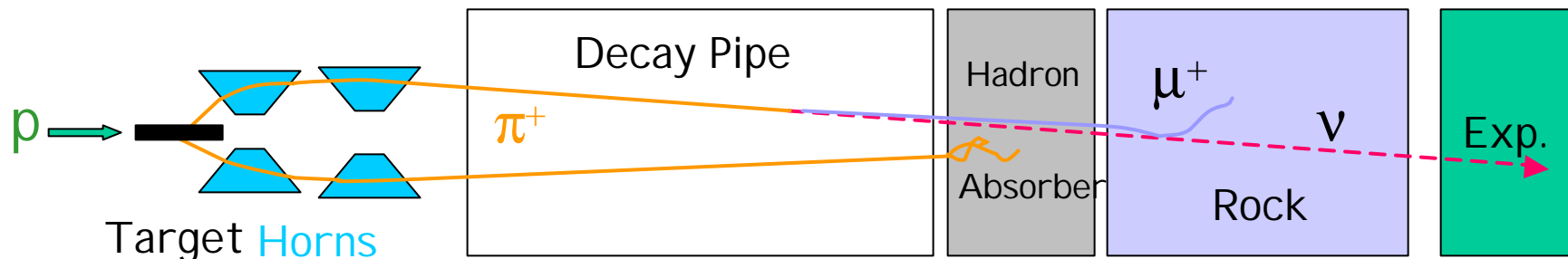
Magnetic horns focus  $\pi^+$

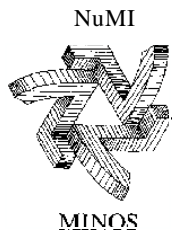
$\pi^+$  decay to  $\mu^+ \nu_\mu$  in a long evacuated pipe

Left-over  $\pi^+$  shower in an absorber

Rock shield ranges out  $\mu^+$

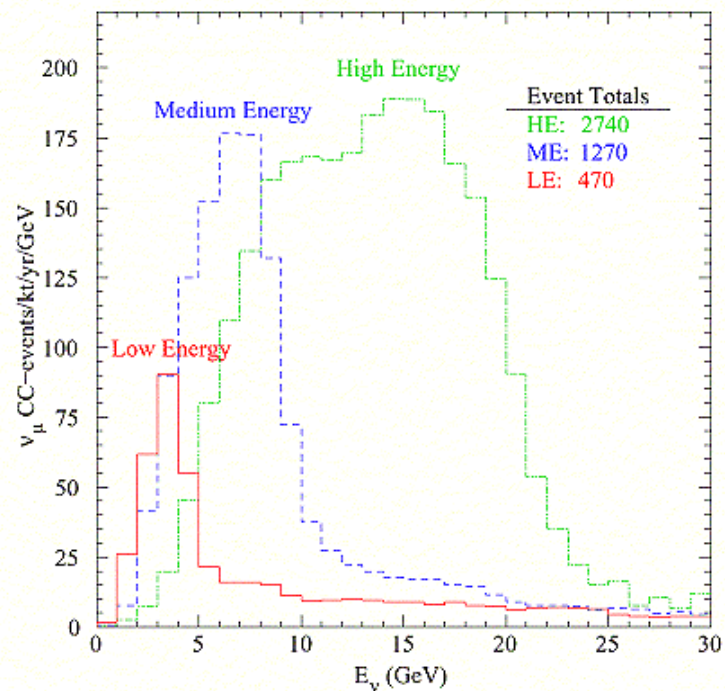
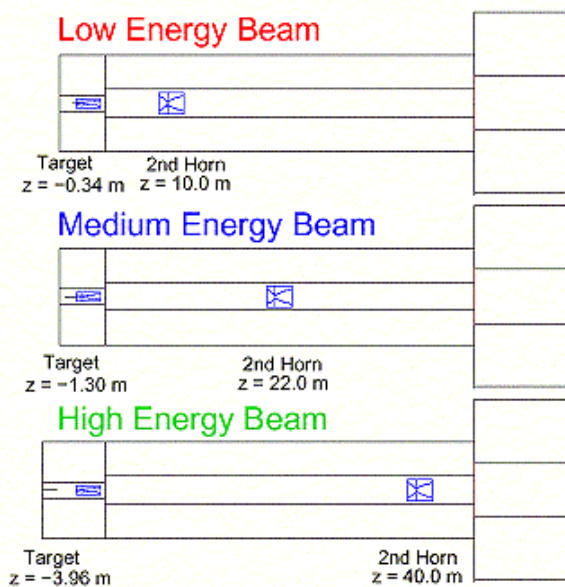
$\nu$  beam travels through earth  
to the experiment



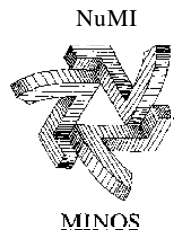


# NuMI Horn Configurations and Beam Spectra

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Initial running in the “Low Energy” configuration  
Expect  $\sim 2500$  Events per “Canonical” NuMI year



# Oscillation effects observable at MINOS

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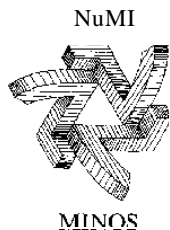
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Title:  
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Creator:  
(ImageMagick)  
Preview:  
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with a preview included in it.  
Comment:  
This EPS picture will print to a  
PostScript printer, but not to  
other types of printers.

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(tau\_rates.eps)  
Creator:  
(ImageMagick)  
Preview:  
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with a preview included in it.  
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This EPS picture will print to a  
PostScript printer, but not to  
other types of printers.

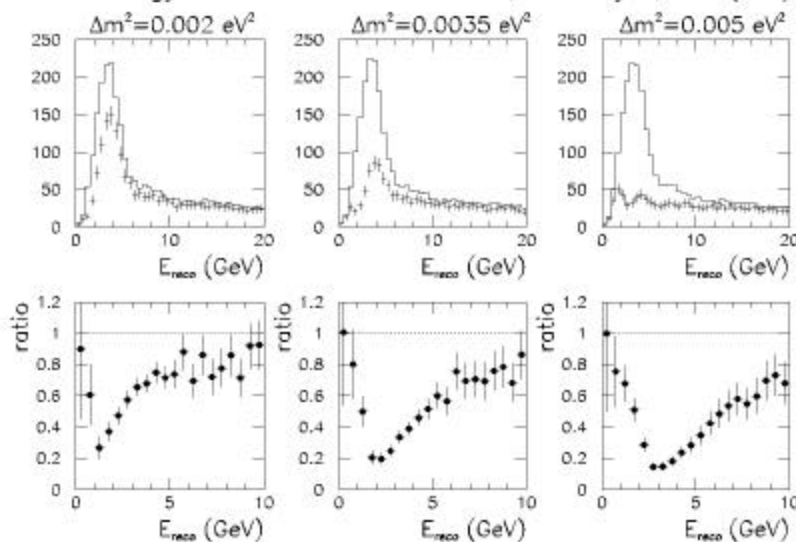


# $\nu_\mu$ CC Event Selection Efficiency and Background Rejection

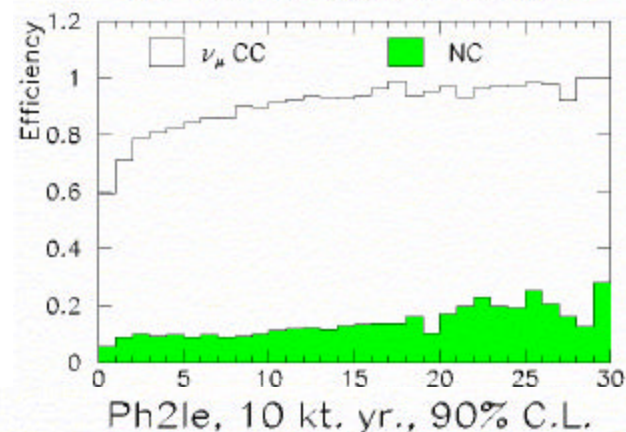
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March 4, 2002  
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With high statistics and good event efficiencies in the energy region of interest MINOS will give substantially improved oscillation parameter measurements in a 2-year run

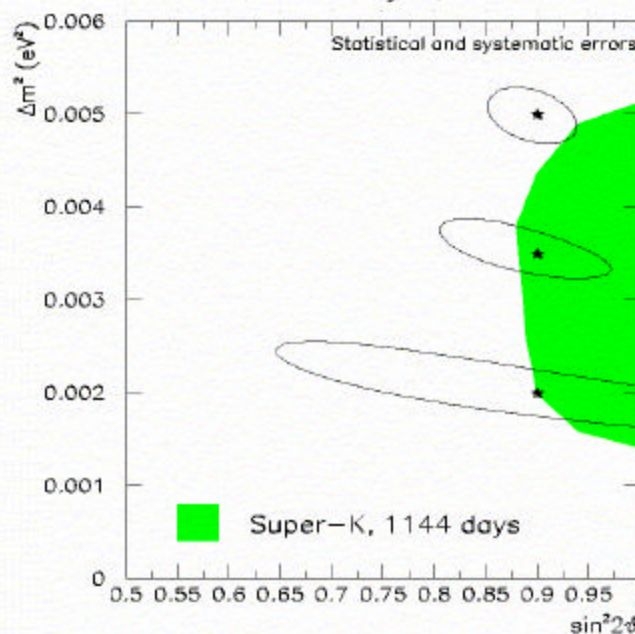
CC energy distributions – Ph2le, 10 kt.yr.,  $\sin^2(2\theta)$



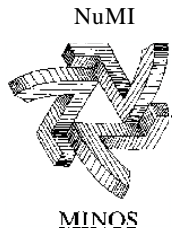
CC-like selection efficiencies



Ph2le, 10 kt. yr., 90% C.L.





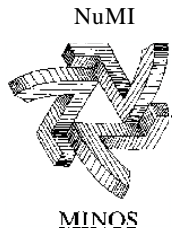


# Control of Systematics

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- Near/Far extrapolation
  - « Primary Control comes from extrapolation of data from the near detector
  - « Efforts aiming at production experiments at CERN & FNAL
- Beam pointing checked by
  - « GPS & laser survey with 1km lever arm on the near site and  $\ll 1\text{m}$  accuracy at the FD site
  - « Checked with ND beam profile, downstream hadron monitors (spot check), and muon monitors in the absorber pile and rock

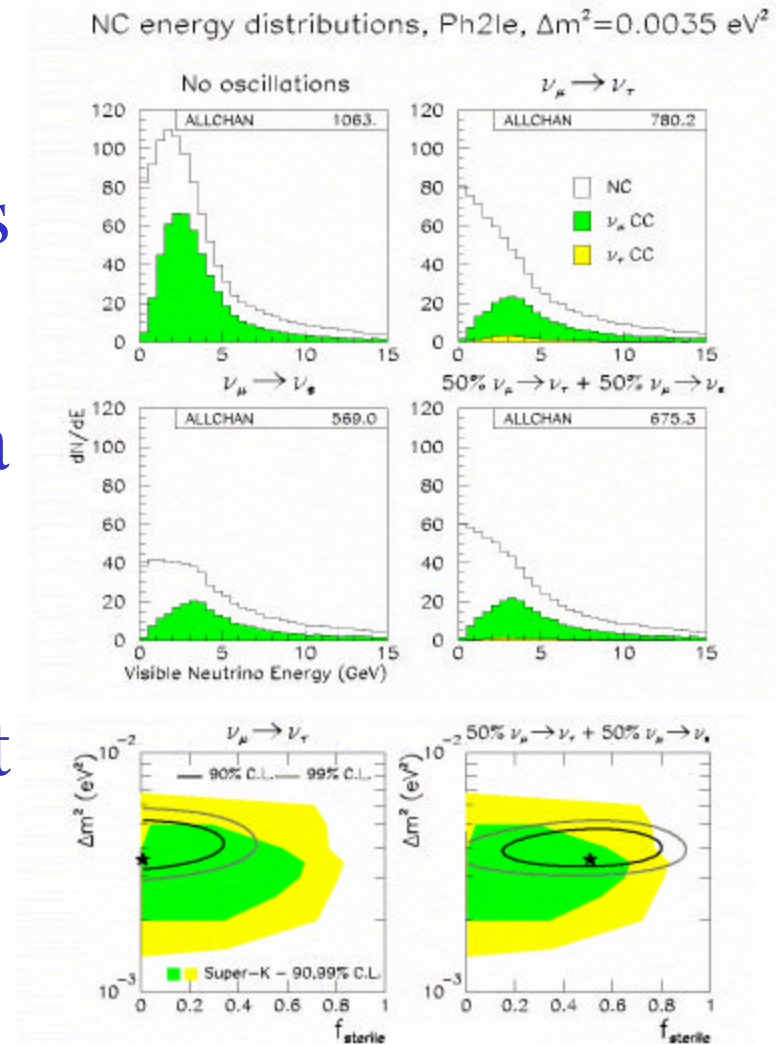


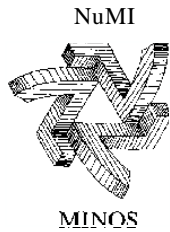


# Improving Limits on Other Oscillation Modes

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- Spectrum of NC-like events in MINOS gives a handle on  $\nu_{\text{sterile}}$
- An analysis of this data and disappearance data give improved constraints over current limits

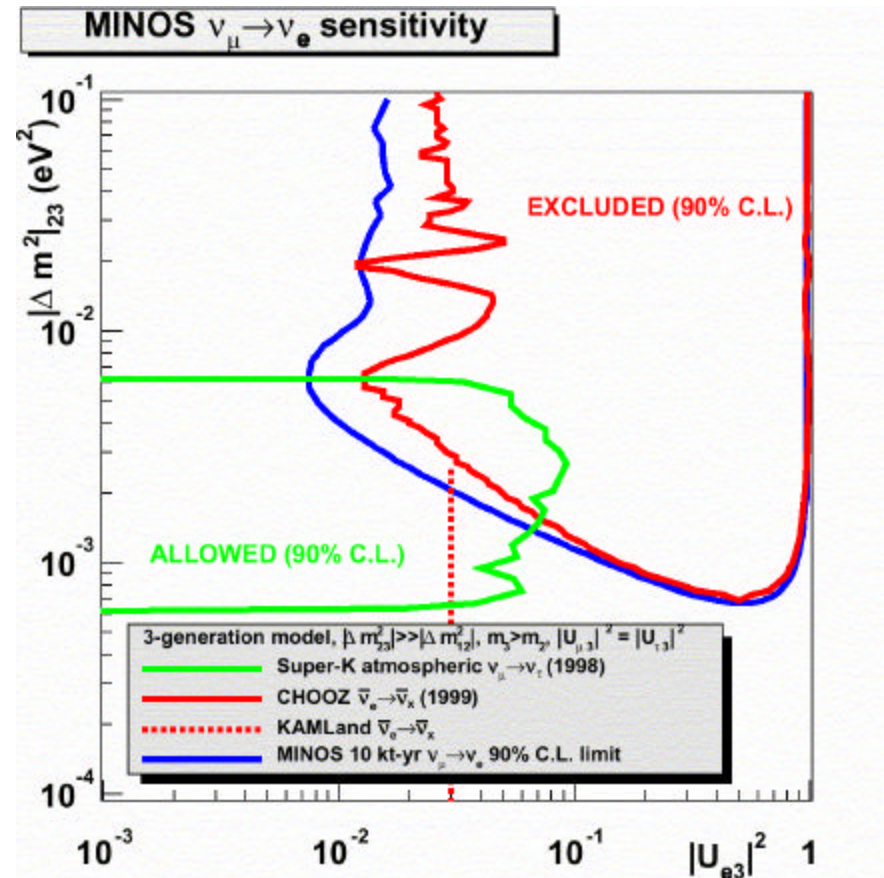


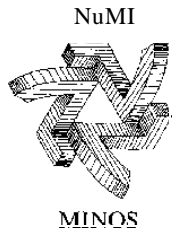


# Improved Limits on Sub-dominate Mode

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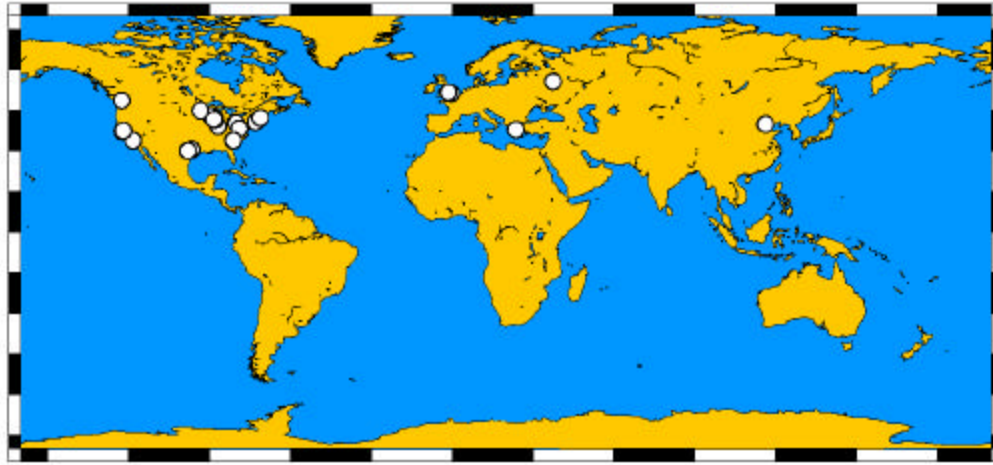
- Analysis uses shower shapes and energies to eliminate backgrounds
- Modest improvement over the current limits
- The limits are shown for a nominal 10kt-year exposure





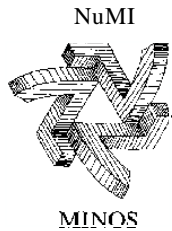
# The MINOS Collaboration

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Over 200 Collaborators  
from 32 Institutions

Argonne • Athens • Brookhaven • Caltech • Cambridge • Chicago • Dubna • Elmhurst  
Fermilab • Harvard • IHEP-Beijing • Illinois Inst. of Technology • Indiana • ITEP-Moscow  
Lebedev • Livermore • Macalester • Minnesota • Minnesota-Duluth • Northwestern  
Oxford • Pittsburgh • Protvino • Rutherford • South Carolina • Stanford • Sussex • Texas A&M  
Texas-Austin • Tufts • University College London • Western Washington • Wisconsin



# MINOS Detectors

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## Basic Ingredients of MINOS detectors

- Fine grained, Iron/scintillator sampling calorimeters
- 2.54cm thick Steel absorbers with toroidal B-field of  $\sim 1.3$  T
- 1 cm thick, 4 cm wide Scintillator strips
- 1.2 mm diameter wavelength shifter and clear fibers
- Multi-Anode PMT
- Front end Electronics & DAQ

## Far Detector at Soudan

- 8m wide octagonal shape
- 486 steel plates
- $\sim 5,400$  tons

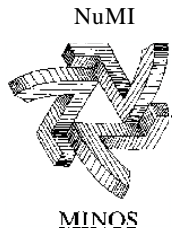
## Near Detector at Fermilab

- 3.8x4.8m squashed octagon
- 282 steel plates
- $\sim 980$  tons

## Both Detectors have

- 2-views for tracking

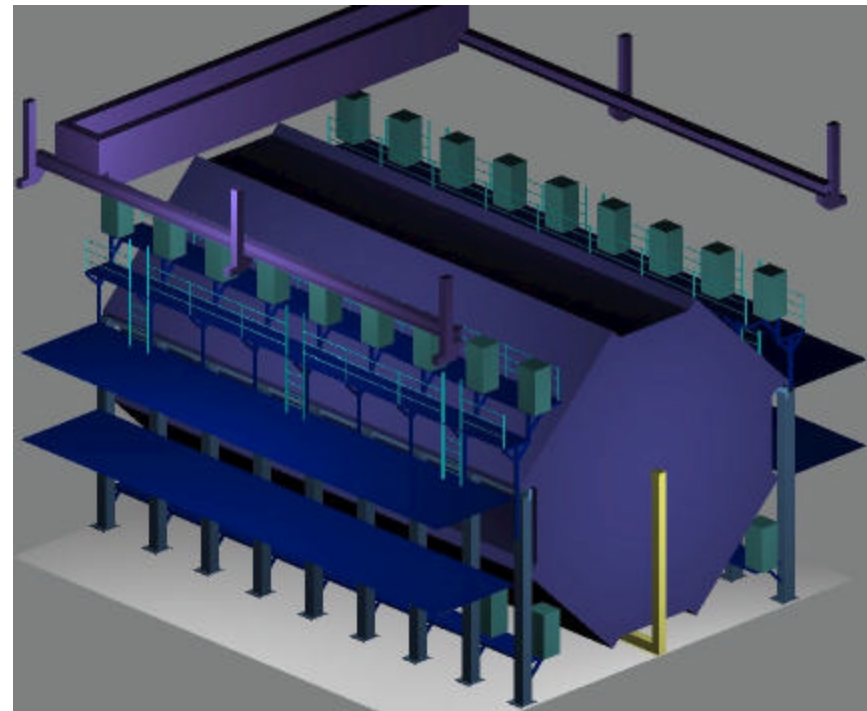




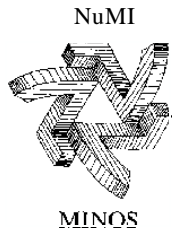
# MINOS Far Detector

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March 4, 2002  
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- 8m Octagonal Tracking Calorimeter
- 486 layers of 2.54cm Fe
- 2 sections, each 15m long
- 4cm wide solid scintillator strips with WLS fiber readout
- 25,800 m<sup>2</sup> active detector planes
- Magnet coil provides  $\langle B \rangle \approx 1.3\text{T}$
- 5.4kt total mass



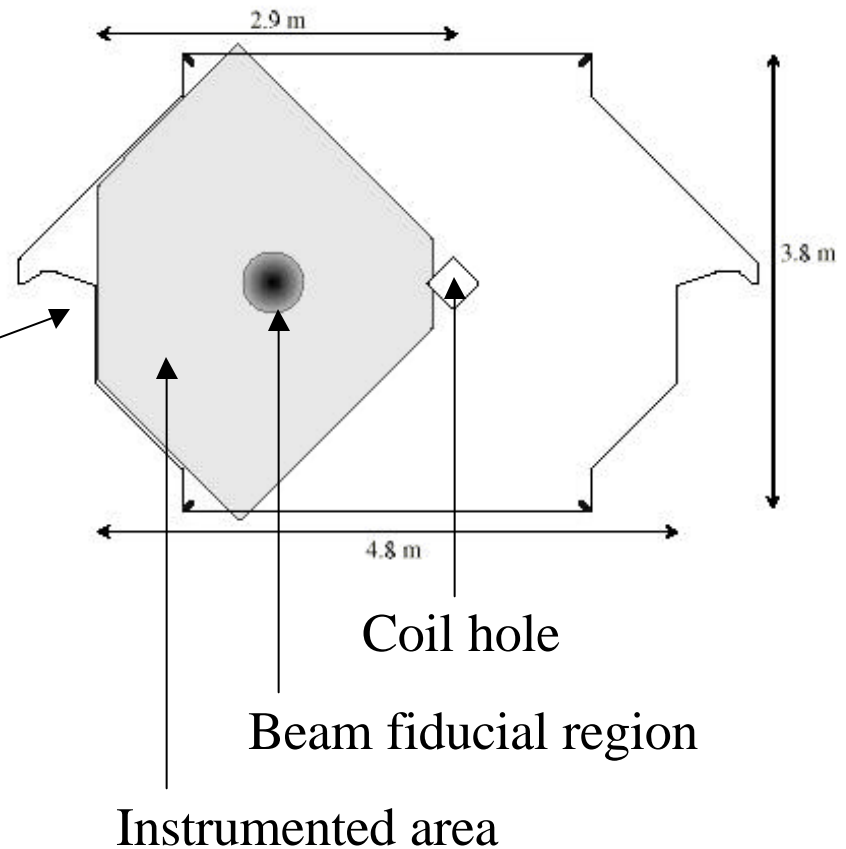
Half of the MINOS Detector

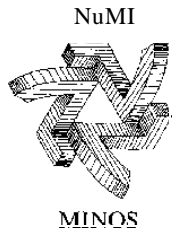


# MINOS Near Detector

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- 16.6 m long, 980 tons
- 282 “squashed octagon” planes
- **Forward section:** 120 planes
  - 4/5 partially instrumented
  - 1/5 planes: full area coverage
- **Spectrometer section:** 162 planes
  - 4/5 planes not instrumented
  - 1/5 planes: full area coverage

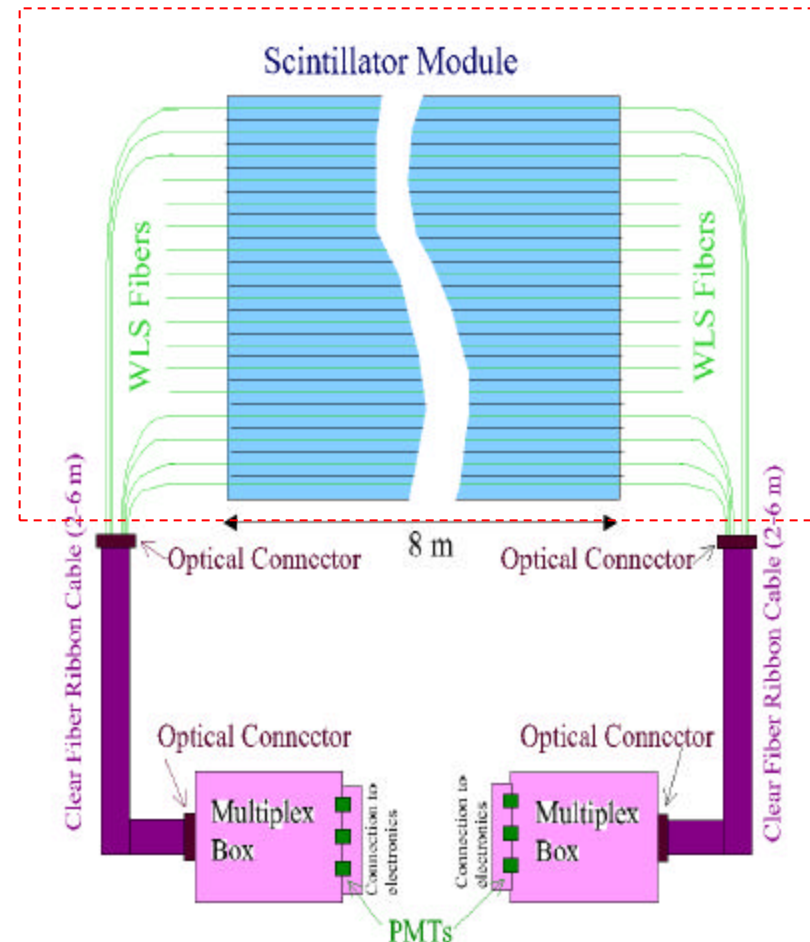




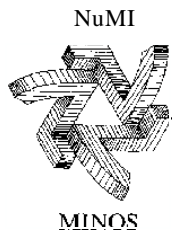
# Scintillator Readout Schematic (Far Detector)

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- Strips assembled into 20- or 28-strip “modules”
- Fire resistant aluminum light cases
- 2-ended WLS fiber readout
- WLS to clear fiber cables at module connectors
- MUX boxes route 8 fibers to one PMT pixel



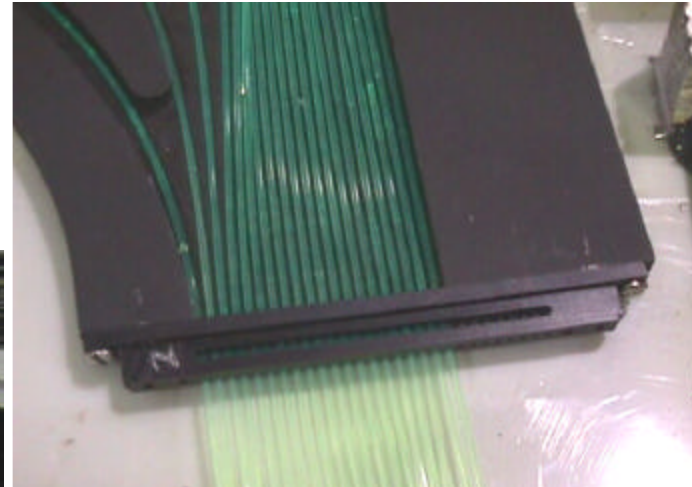


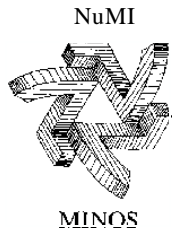


# MINOS Scintillator System

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- An 8 m Scintillator Module
- Optical fiber readout
- A 16-channel M16 PMT





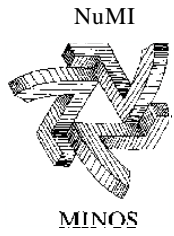
# NuMI Excavation

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- Drill and blast for the shafts, target hall, support labyrinth, absorber hall, and MINOS near detector hall
- Tunnel boring machine for the decay tunnel and access to the ND hall and beam monitoring areas



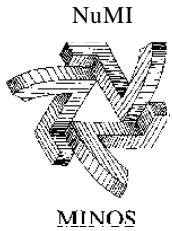




# NuMI Excavation Summer 2001

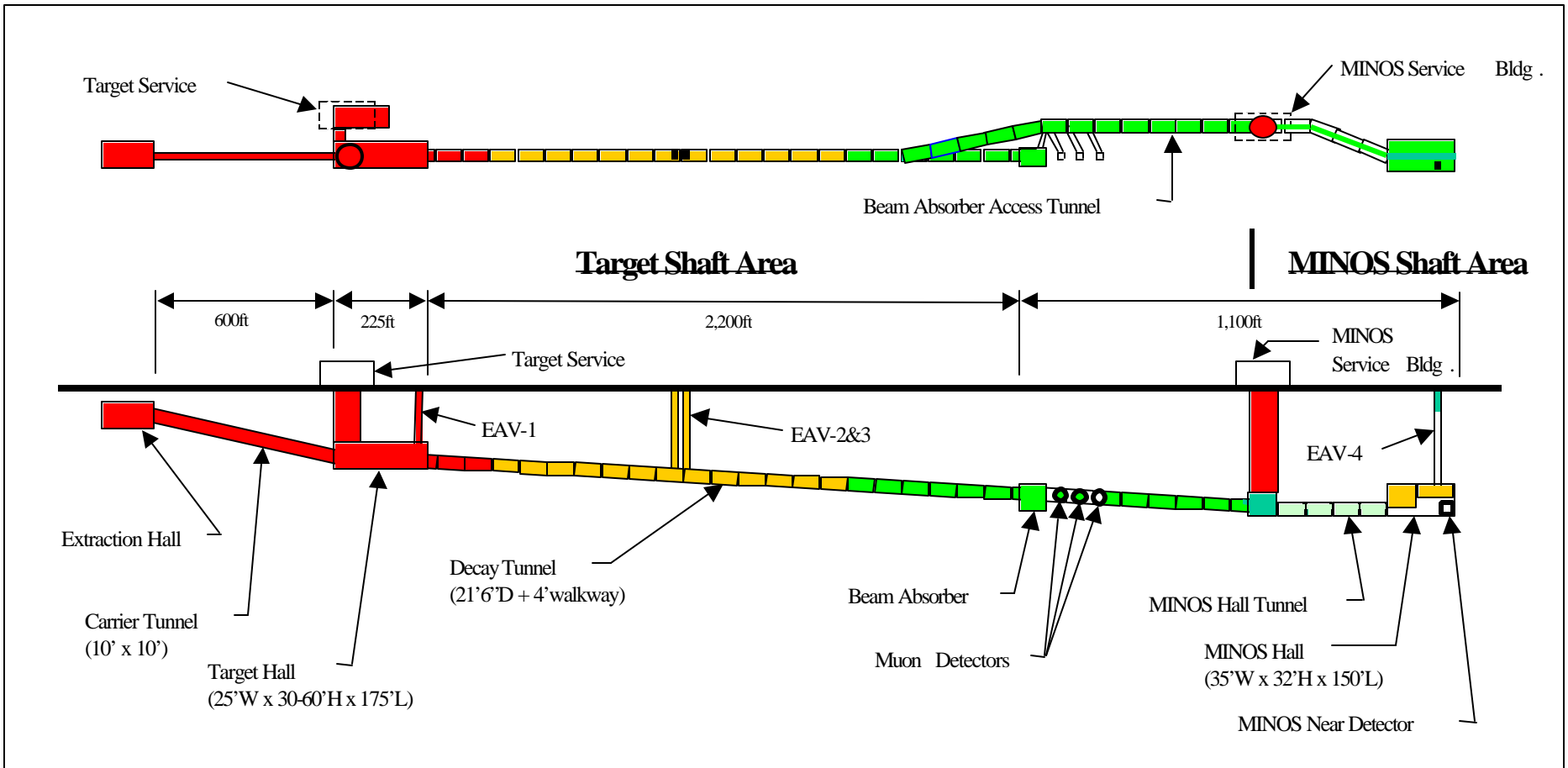
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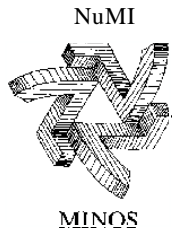


# NuMI Tunnels and Halls Construction Progress

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- As of June 2001
- As of Sep 2001
- As of Jan 2002

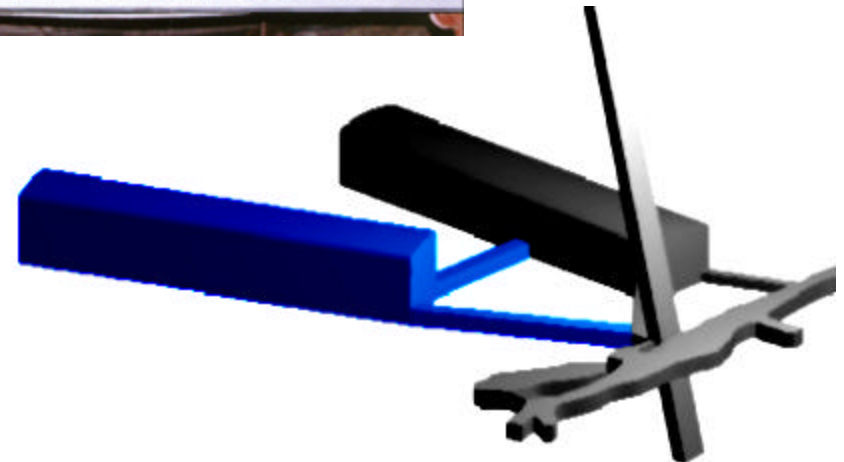


# Soudan Underground Laboratory

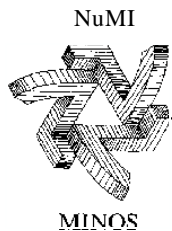
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The Soudan shaft  
limits objects to a  
maximum size of  
1m by 2m by 9m





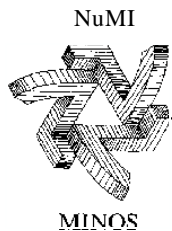


# MINOS Progress - Far Detector Hall

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# MINOS Far Detector Installation - Detector Plane Assembly

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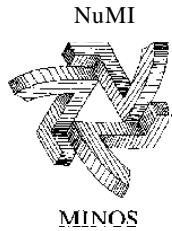


Detector plane strongbacks



Scintillator module mounting





## Current Status of Far Detector

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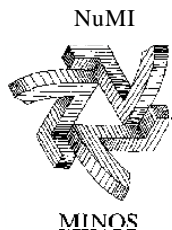
146 Planes mounted as of 3/1/2002

136 Planes being readout through DAQ

1.6kT of fiducial mass

1/3 pe threshold with 4 out of 5 planes required  
for trigger

~2% of the detector is being mounted per day at  
the present rate of assembly



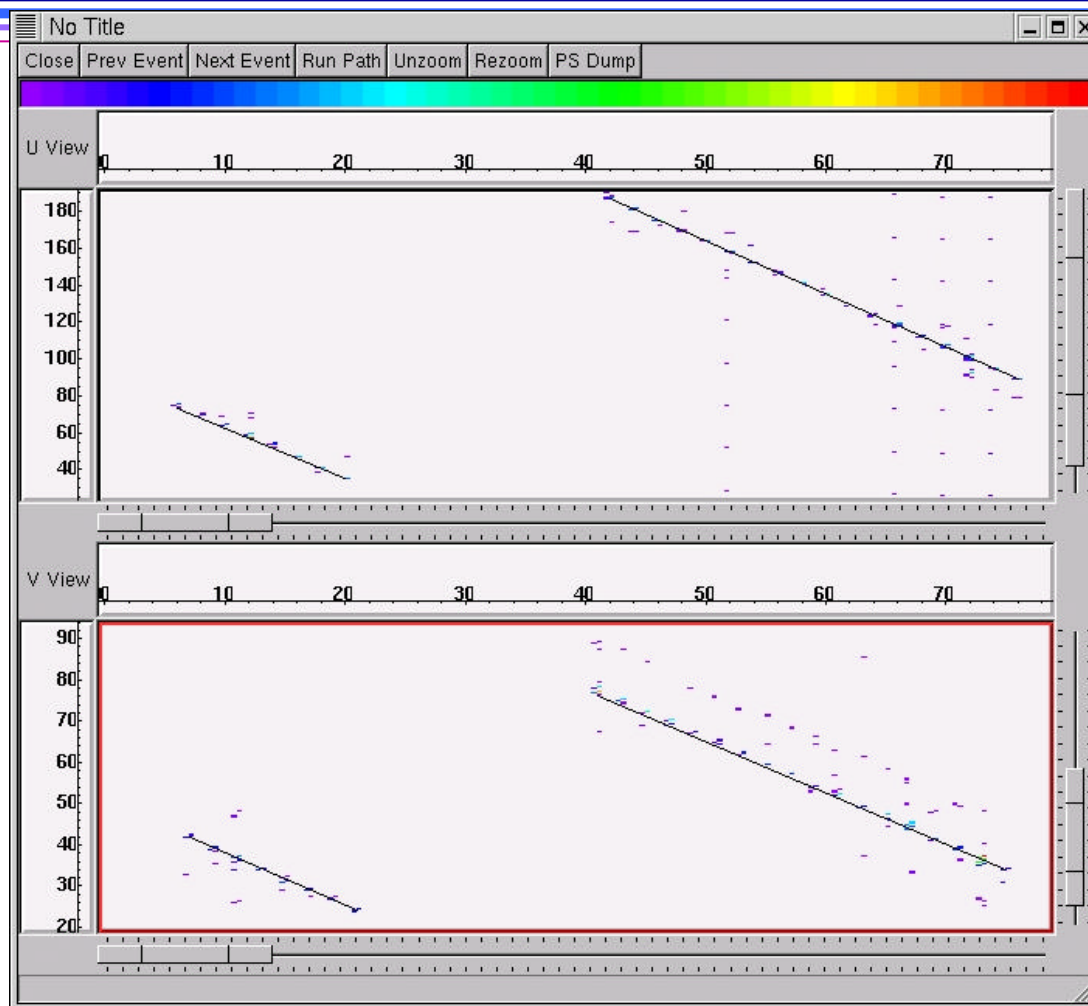
# Double Cosmic-Ray Muon in MINOS Far Detector

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Scintillator strip number →

‘u’ view

‘v’ view

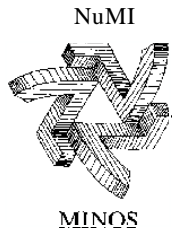


Color indicates pulse size

Extra hits from cross talk and deMUX failures

Lines are from fits by track reconstruction software

Detector plane number (1-76) →



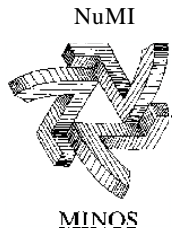
# Near Detector Construction

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Construction beginning in early 2002 **(NOW)**



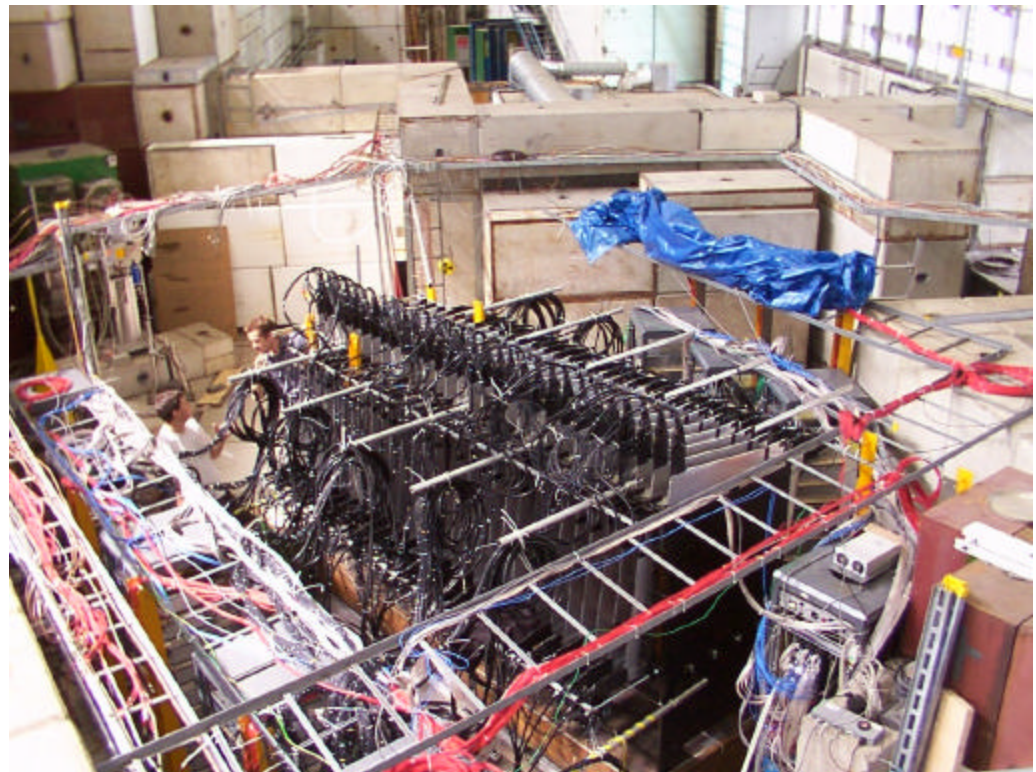


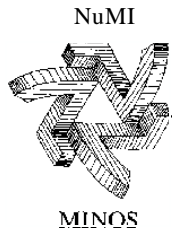


# The MINOS Calibration Detector

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- 60 1m X 1m MINOS modules with Far Detector readout Electronics.
- Assembled in the CERN T11 beam in the Spring of 2001.
- Completed run with energies up to 3 GeV .
- Schedule to run at higher energies in 2002.





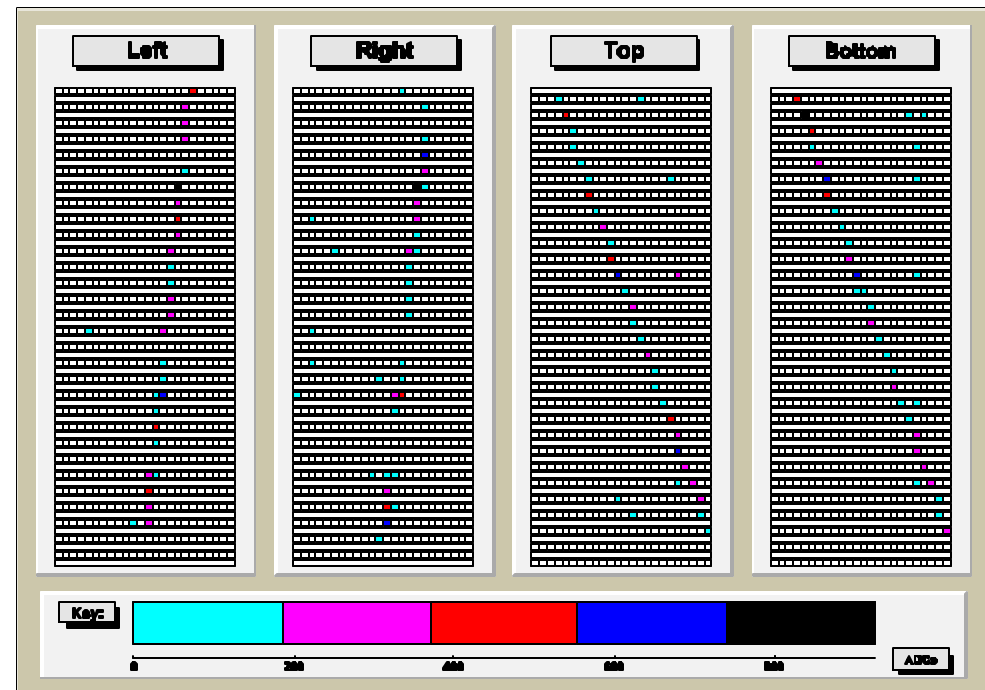
# The MINOS Calibration Detector

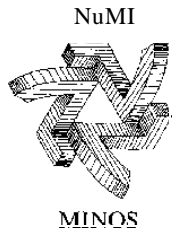
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March 4, 2002  
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Data taken with

- Pions
- Electrons
- Protons(using TOF)
- Muons

Sample Muon event  
displayed to the right

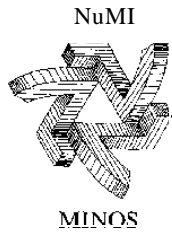




# Schedule

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March 4, 2002  
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- Far Detector installation ('01-'03)
  - « Cosmic rays and atmospheric neutrinos with half the detector (and the magnetic field) in summer 2002
  - « 10kt-years by 2005
  - « CP tests in upward muons and momentum analyzed partially-contained events
- Near Detector assembly and installation ('01-03')
- Beam line commissioning ('04-'05)



# Summary

- MINOS will confront the atmospheric neutrino signal with -
  - High statistics accelerator data
  - Good flavor separation
  - Good control of systematics
- With initial calibration data runs completed at CERN and beginning at Soudan with the *real* detector...
  - We are on the road to exciting neutrino physics in the next few years!