

# Phenomenology of Extra Dimensions

Review: J LH, Spiropulu

Formal Theorists: have lived in 4+n dimensions since 1920's (Kaluza-Klein)

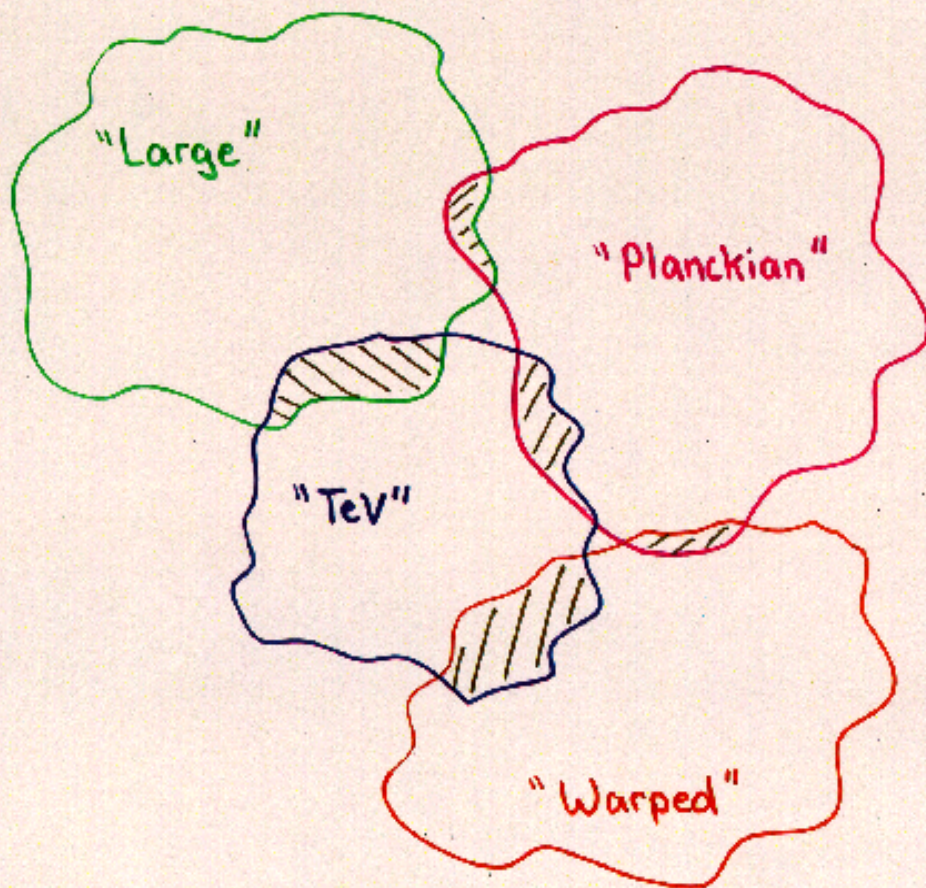
New Theories have testable consequences!

- Several interesting scenarios
- Each has distinctive phenomenology

- General pedagogical introduction
- Examine phenomenology of various scenarios
  - ⇒ Collider signatures



## The Physics of Extra Dimensions



There are many different models with extra dimensions!

Resulting phenomenology is distinct.



## Hierarchy Problem

1) Generation :  $\Lambda_{Pl} \gg \Lambda_{EW}$

2) Stabilization :  $\Lambda_{EW}$  unstable under quantum corrections

Traditional Solution: Add new particles at  $\Lambda_{EW}$   
 $\Rightarrow$  Supersymmetry

New Ideas: Modify Gravity  
Gravity is not well-tested  
@ short distances

Prediction: New Physics at  $\Lambda_{EW}$  !

$$\Lambda_{NP} \sim \Lambda_{EW} \sim \text{TeV}$$

+

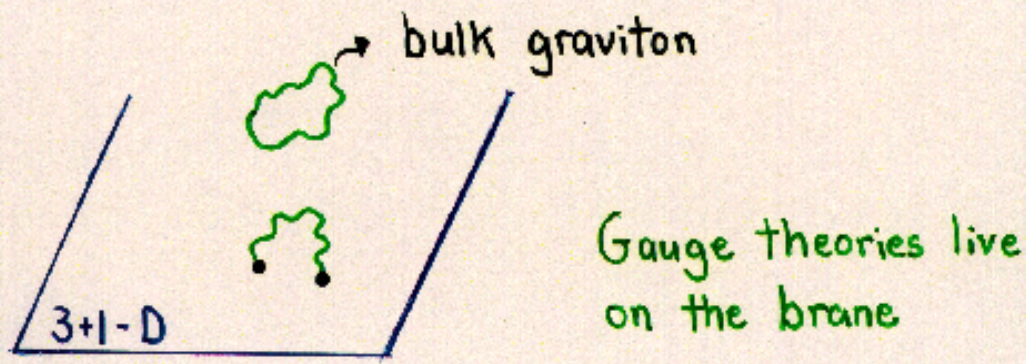
New Colliders @ TeV-scale

= Can decisively prove or disprove these theories !

Exciting times ahead of us!



## Physics of 3-branes: string theory

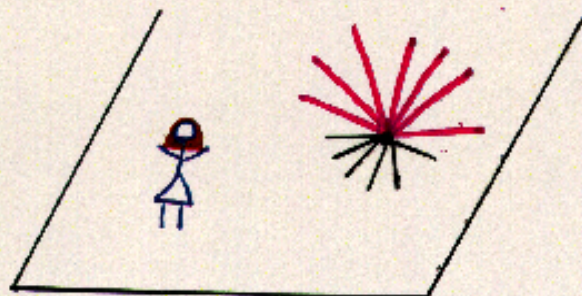


- Gauge particles live at end of strings
- Strings can pop off of brane
  - ⇒ are neutral with no gauge charges
  - = bulk gravitons

N.B.: This is a very simplified picture ....



## A 3-brane Universe



Standard Model  
forces stuck on 3-brane

Gravitational fields  
spread out over  
all spacetime

Are gravitational fields diluting too quickly?

→ Extra dimensions must be compactified

recover  $F_{Gr} \sim 1/r^2$  on the brane



## Compactification

Bulk fields expand into Kaluza-Klein towers



5-d kinetic motion  
appears as mass  
in 4-d

⇒ Get Kaluza-Klein tower state for each  
5-d revolution

$$\Phi(x^M, x^a) = \sum_{n=0}^{\infty} \phi^{(n)}(x^M) e^{inx^a/R_c} \cdot \frac{1}{\sqrt{V_n}}$$

$\delta^2 \Phi = 0$  gives massless mode in 4+n D

$$(\delta_M^2 + \delta_a^2) \Phi = \sum_n \left[ \delta_M^2 - \left(\frac{n}{R_c}\right)^2 \right] \phi^{(n)}(x^M) e^{inx^a/R_c}$$

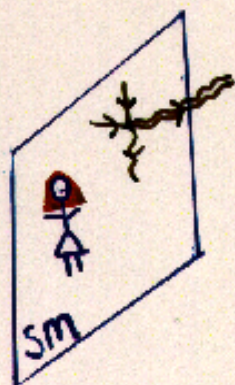
↑ mass term,  $m = \frac{n}{R_c}$



# Large Extra Dimensions

Arkani-Hamed,  
Dimopoulos, Dvali  
SLAC-PUB-7801

Motivation: Solve the hierarchy problem by removing it!



SM fields are constrained to the 3-brane (the wall)

Gravity is allowed to propagate in the  $n$  extra dimensions (the bulk)

Gauss' Law:  $m_{pl}^2 = V_\delta m_*^{2+\delta}$  ;  $V_\delta = A_\delta R^\delta$  \*\*

$m_*$  = Fundamental Planck scale in the bulk  
 $\approx \text{TeV}$

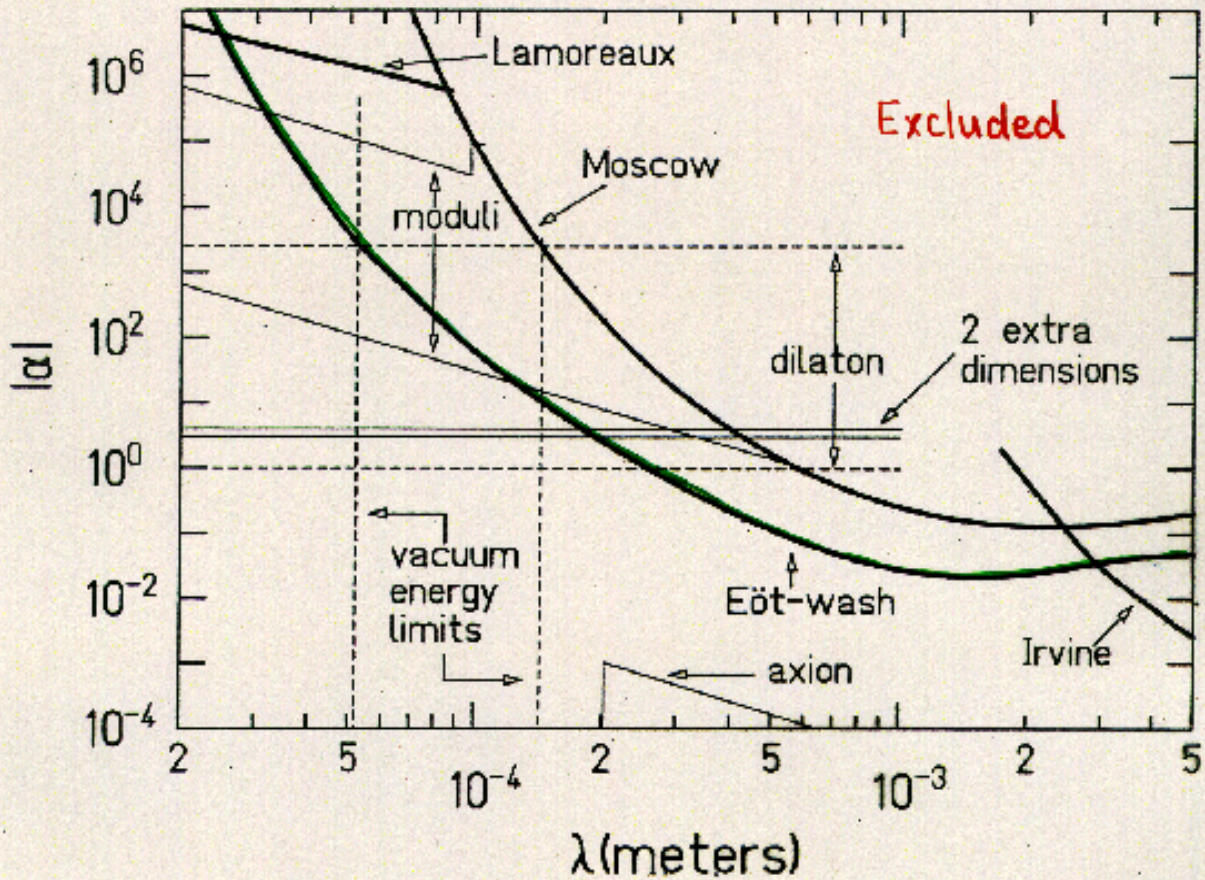
$\delta = 1$	$R = 10^{11} \text{ m}$	Excluded!
2	0.4 mm	$m_c = 1/R = 5 \times 10^{-4} \text{ eV}$
4	$10^{-5} \text{ mm}$	20 KeV
6	30 fm	7 MeV

\*\* Assumed all  $\delta$  new dim's have same size

$A_\delta = \frac{2\pi^{\delta/2}}{\Gamma(\delta/2)}$  (sphere) or  $(2\pi)^\delta$  (torus) or ... 7



# Constraints from Cavendish-type experiments



$\lambda \leq 220 \mu$  Hoyle et al  
 $m_* \geq 1.6 \text{ TeV}$  for  $n=2$

*transmission 2 raised to the power of*

$$V_{\text{gravity}} \sim \frac{m_1 m_2}{m_*^{2+n}} \frac{\lambda_{\text{axio}}}{r^{n+1}}$$

$(r \ll R)$

$$\sim \frac{m_1 m_2}{m_{\text{pl}}^2} \frac{1}{r}$$

$(r \gg R)$



## Constraints from Astro/Cosmology

(i) SN Cooling

Cullen + Perelstein  
Garger et al  
Savage et al

$NN \rightarrow NN + G_n$  can cool supernova too rapidly

(ii)  $\gamma$ -Ray Flux

Hall + Smith

$\nu\bar{\nu} \rightarrow G_n \rightarrow \gamma\gamma$  produces too many soft  $\gamma$ 's

(iii) Matter dominated early universe  
(KK States)

Fairbairn

(iv) Re-heating of universe

Hannstad

### Constraints on $m_*$ (TeV)

	$\delta = 2$	3	4	5
(i)	31	2.75	—	—
(ii)	110	5	—	—
(iii)	86	7.4	1.5	—
(iv)	167	22	4.75	1.55

$\Rightarrow$  Low  $m_*$  disfavored for  $\delta = 2$

Given theoretical uncertainties,  $\delta \geq 3$  OK



## Bulk Metric: Linearized Quantum Gravity

$$G_{AB} = \eta_{AB} + \frac{h_{AB}(x^m, x^a)}{m_*^{\delta/2+1}}$$

$$\begin{aligned} A &= 0, \dots, 3+\delta \\ m &= 0, 1, 2, 3 \\ a &= 4, \dots, 3+\delta \end{aligned}$$

## Interactions:

$$S_{\text{int}} = \frac{-1}{m_*^{\delta/2+1}} \int d^4x d^{\delta}x^a h_{AB}(x^m, x^a) T_{AB}(x^m, x^a)$$

$$\text{SM on wall: } T_{AB} = \eta^m_A \eta^a_B T_{mv} \delta(x^a)$$

## Graviton KK Reduction

Massless $h_{AB}$ in $\delta$ -dim	{	$h_{\mu\nu}$	1 massless 2-component tensor
		$h_{\mu i}$	$\delta$ massless 2-component vectors
		$h_{ij}$	$\frac{1}{2}\delta(\delta+1)$ real massless scalars

States in KK tower acquire mass by 'eating'  
lower spin fields

## Physical Fields:

1 massive 5-component tensor $h_{\mu\nu}$	{	with massless zero-modes
$\delta-1$ massive gauge fields $A_\mu$		
$\frac{1}{2}\delta(\delta-1)$ scalars		



## Decomposition of Graviton Field-Strength Tensor

$$h_{AB} \sim \begin{pmatrix} h_{\mu\nu} + \eta_{\mu\nu} \phi & A_{\mu a} \\ A_{\mu a} & \phi_{ab} \end{pmatrix}$$

## Graviton-fermion Interactions

$$T_{AB} = \eta^{\mu}_A \eta^{\nu}_B [\bar{\psi} \gamma_{\mu} \partial_{\nu} \psi - \partial_{\mu} \bar{\psi} \gamma_{\nu} \psi]$$

Spin 1 KK states:  $T_{\mu a} = 0$

Spin 0 KK states:  $T_{ab} = 0$

$$\begin{aligned} T_{\mu}{}^{\mu} &= \bar{\psi} \gamma_{\mu} \delta^{\mu} \psi - \partial_{\mu} \bar{\psi} \gamma^{\mu} \psi \\ &= -2m \bar{\psi} \psi \end{aligned}$$

Spin 2 KK states: Given by linearized G.R.

Interactions:

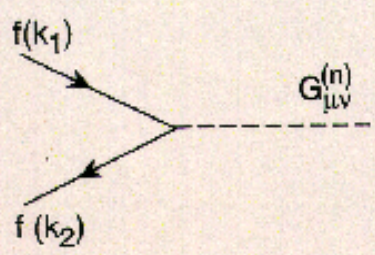
$$S = \frac{-1}{m_*^{\eta/2+1}} \int d^4x d^n x^a h_{AB}(x^{\mu}, x^a) T_{AB}(x^{\mu}, x^a)$$

SM on wall:  $T_{\alpha\beta} = \eta^{\mu}_{\alpha} \eta^{\nu}_{\beta} T_{\mu\nu} \delta(\nu a)$



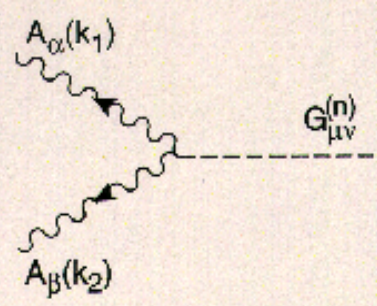
# Feynman Rules

Giudice, Rattazzi, Wells  
Han, Lykken, Zhan



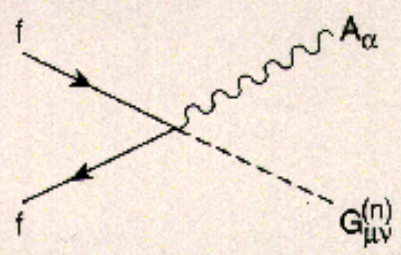
$$-\frac{i}{4\overline{M}_P} [W_{\mu\nu}^{(f)} + W_{\nu\mu}^{(f)}]$$

$$W_{\mu\nu}^{(f)} = (k_1 + k_2)_\mu \gamma_\nu \tag{53}$$



$$-\frac{i}{\overline{M}_P} [W_{\mu\nu\alpha\beta}^{(\gamma)} + W_{\nu\mu\alpha\beta}^{(\gamma)}]$$

$$W_{\mu\nu\alpha\beta}^{(\gamma)} = \frac{1}{2} \eta_{\mu\nu} (k_{1\beta} k_{2\alpha} - k_1 \cdot k_2 \eta_{\alpha\beta}) + \eta_{\alpha\beta} k_{1\mu} k_{2\nu} + \eta_{\mu\alpha} (k_1 \cdot k_2 \eta_{\nu\beta} - k_{1\beta} k_{2\nu}) - \eta_{\mu\beta} k_{1\nu} k_{2\alpha} \tag{54}$$



$$-\frac{i}{2\overline{M}_P} eQ (X_{\mu\nu\alpha} + X_{\nu\mu\alpha})$$

$$X_{\mu\nu\alpha} = \gamma_\mu \eta_{\nu\alpha} \tag{55}$$

Massless 0-mode ⊕ all tower gravitons have same couplings to matter

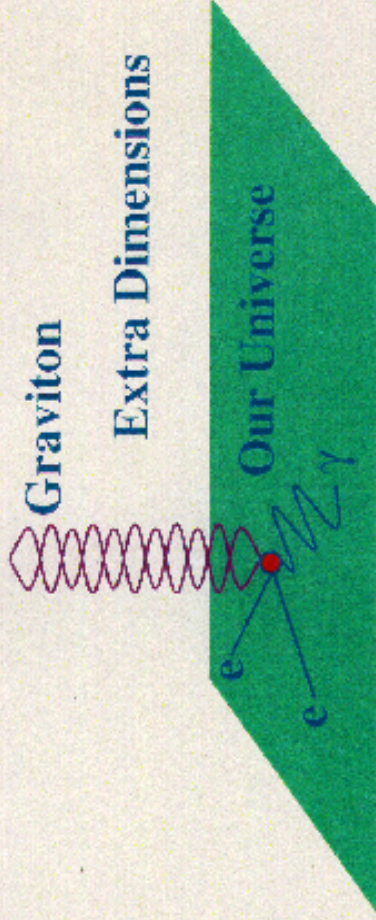




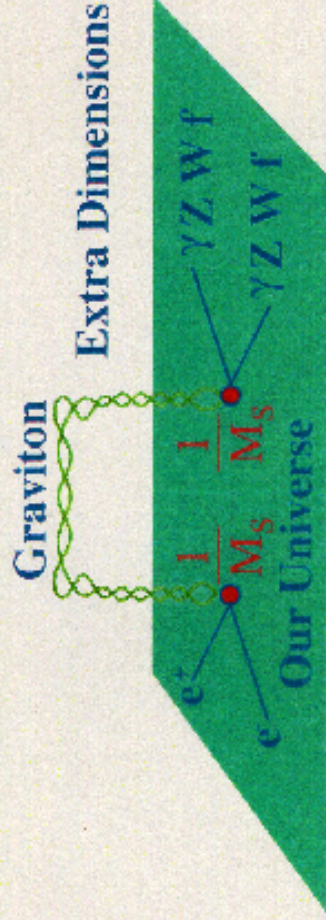
### Search Strategy



Direct Search: 1 photon or 1 Z boson + missing energy.



Indirect Search: Look for deviations from  $(d\sigma/d\Omega)_{SM}$ .



## 2 Classes of Collider Tests



• Graviton Tower Exchange  $XX \rightarrow G_n \rightarrow YY$

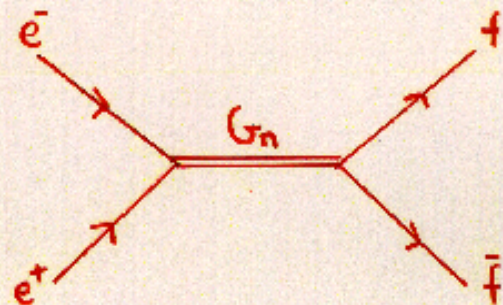
Search for 1) Deviations in Standard processes  
2) New processes!

Graviton couples 1) universally to everything  
2) via stress-energy tensor

Angular distributions reveal spin-2 exchange

Consider  $e^+e^- \rightarrow f\bar{f}$

$$M = \frac{1}{16 m_{Pl}^2} \sum_n \frac{T_{\mu\nu}^e P^{\mu\nu\lambda\sigma} T_{\lambda\sigma}^f}{s - m_n^2 + i\epsilon}$$



$G_n$  are densely packed!

$$\Rightarrow \sum_n \rightarrow \int dm^2 p(m^2)$$

$$\frac{1}{m_{Pl}^2} \sum_n \frac{1}{s - m_n^2} \rightarrow \frac{1}{m_{*}^4}$$

JLH  
Phys Rev Lett  
Giudice, Rattae  
We



## Poor theoretical control!

∫ KK propagators is divergent

⇒ Sensitivity to unknown ultraviolet physics

### Approaches:

- 'Naive' cut-off JLH; Giudice et al; Han et al
- Brane fluctuations Bando et al
- Weakly Coupled string theory Dudas, Mourad  
Accomando et al  
Cullen et al

### Cut-off Approach

Contact interaction limit for  $G_n$  exchange

Examine leading dimension-8 operators

\* constrain  $m_* |\lambda|^{-1/4}$

$$m = \frac{\lambda}{m_*^4} \left\{ \begin{array}{l} \bar{f} \gamma^\mu f \bar{\ell} \gamma_\mu \ell (p_f - p_{\bar{f}}) \cdot (p_{\bar{\ell}} - p_\ell) \\ + \bar{f} \gamma^\mu f \bar{\ell} \gamma_\nu \ell (p_f - p_{\bar{f}})_\nu (p_{\bar{\ell}} - p_\ell)_\mu \end{array} \right\}$$





# Hewett, GRW, and HLZ Formalisms

- ↓ Hewett: neither sign of the interference nor the dependence on the number of extra dimensions is known; therefore the **interference term** is  $\sim \lambda/M_S^4(\text{Hewett})$ , where  $\lambda$  is of order 1; numerically uses  $\lambda = \pm 1$
- ↓ GRW: sign of the interference is fixed, but the dependence on the number of extra dimensions is unknown; therefore the **interference term** is  $\sim 1/\Lambda_T^4$  (where  $\Lambda_T$  is their notation for  $M_S$ )
- ↓ HLZ: not only the sign of interference is fixed, but the n-dependence can be calculated in the effective theory; thus the **interference term** is  $\sim \mathcal{F}M_S^4(\text{HLZ})$ , where  $\mathcal{F}$  reflects the dependence on the number of extra dimensions:

$$\mathcal{F} = \begin{cases} \log\left(\frac{M_S^2}{s}\right), & n = 2 \\ \frac{2}{n-2}, & n > 2 \end{cases}$$

- ↓ **Correspondence** between the three formalisms:

$$M_S(\text{Hewett})_{\lambda=\pm 1} = \sqrt{\frac{2}{\pi}} \Lambda_T(\text{GRW})$$

$$\frac{\lambda}{M_S^4(\text{Hewett})} = \frac{\pi}{2} \frac{\mathcal{F}}{M_S^4(\text{HLZ})}$$

$$\frac{1}{\Lambda_T^4(\text{GRW})} = \frac{\mathcal{F}}{M_S^4(\text{HLZ})}$$

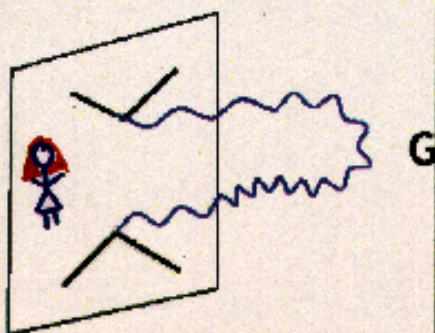
- ↓ Rule of thumb:

$$M_S(\text{Hewett})_{\lambda=\pm 1} \approx M_S(\text{HLZ})_{n=5}$$

$$\Lambda_T(\text{GRW}) = M_S(\text{HLZ})_{n=4}$$

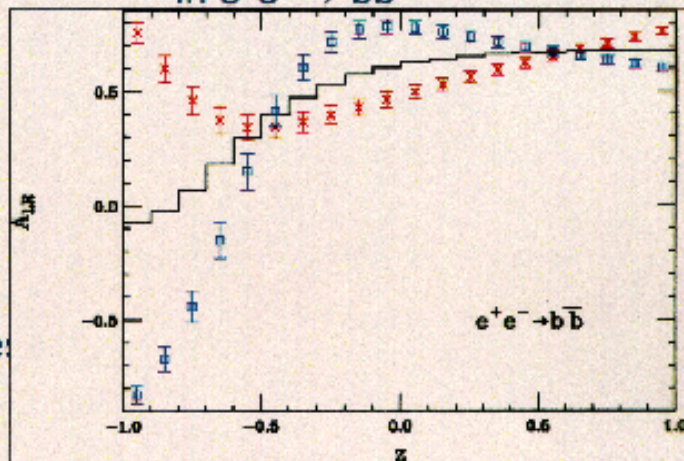


# Graviton Exchange



Deviations in SM processes  
Search for new processes

Polarized  $\angle$  Dist'btn  
In  $e^+e^- \rightarrow b\bar{b}$



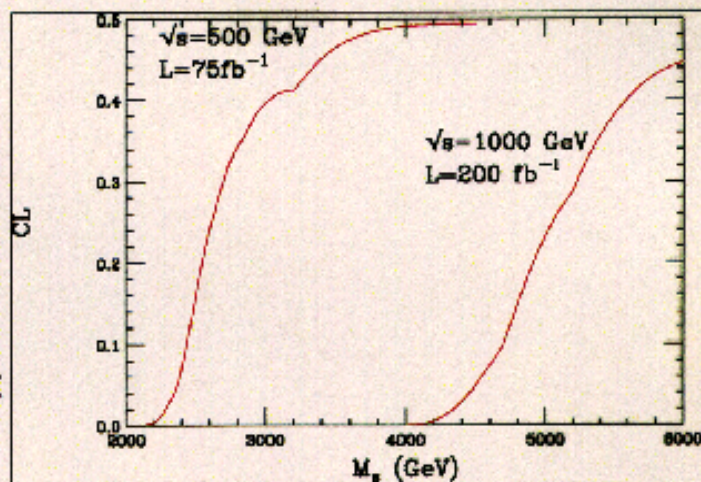
LC :	$e^+e^- \rightarrow f\bar{f}$	$\sqrt{s} = 500 \text{ GeV}$	4.1
		$\sqrt{s} = 1000 \text{ GeV}$	7.2
	$\gamma\gamma \rightarrow WW$	$\sqrt{s} = 1000 \text{ GeV}$	13.0
LHC:	$pp \rightarrow l^+l^-$		6.0

JLH  
Search Reach  
at 95% CL  
(in TeV)

Angular Dist'btns reveal  
spin-2 exchange!

Confidence Level of fit of  
spin-2 data to spin-1  
hypothesis

Spin-2 determined almost  
Up to kinematic limit !



JLH



# Large Extra Dimensions

$$M_{Pl}^2 = V_\delta M_*^{2+\delta}; \quad \text{Solves hierarchy if } M_* \sim \text{TeV}$$

Graviton Emission:



SM stuck to brane  
Gravity propagates in bulk  
Graviton appears as  $E_T$

Giudice, Rattazzi, Wells  
Han, Lykken, Zhang  
Mirabelli, Perelstein, Pes

Discovery Reach in TeV for  $M_*$

TESLA TDR	$e^+e^- \rightarrow \gamma + G_n$	$\sqrt{s} = 800 \text{ GeV}$	2	4	6
	LC	$P_{-,+} = 0$	5.9	3.5	2.5
	LC	$P_- = 0.8$	8.3	4.4	2.9
	LC	$P_- = 0.8, P_+ = 0.6$	10.4	5.1	3.3
ATLAS Study	$pp \rightarrow g + G_n$		2	3	4
	LHC		4 - 8.9	4.5 - 6.8	5.0 - 5.8

Discovery  
range wh  
effective  
theory is  
valid

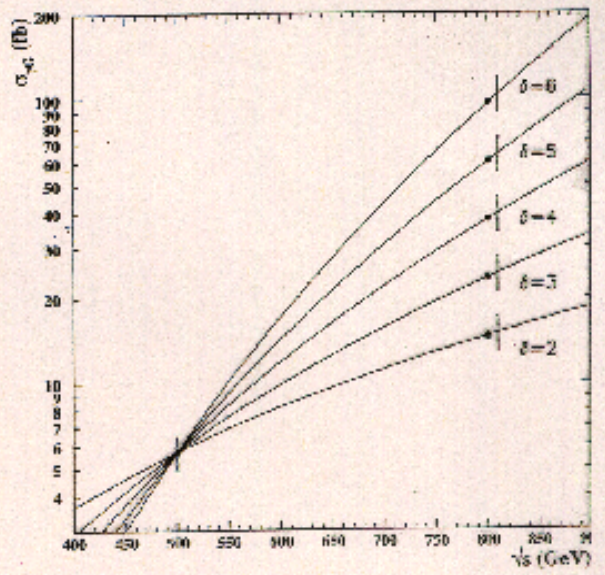
Hinchliffe + Vacavent

- Discovery likely at LHC/LC
- LC probes theory

Spectrum varies with  $\delta$   
as  $\sqrt{s}$  increases  
 $\Rightarrow$  Determine  $M_*$  and  $\delta$ !

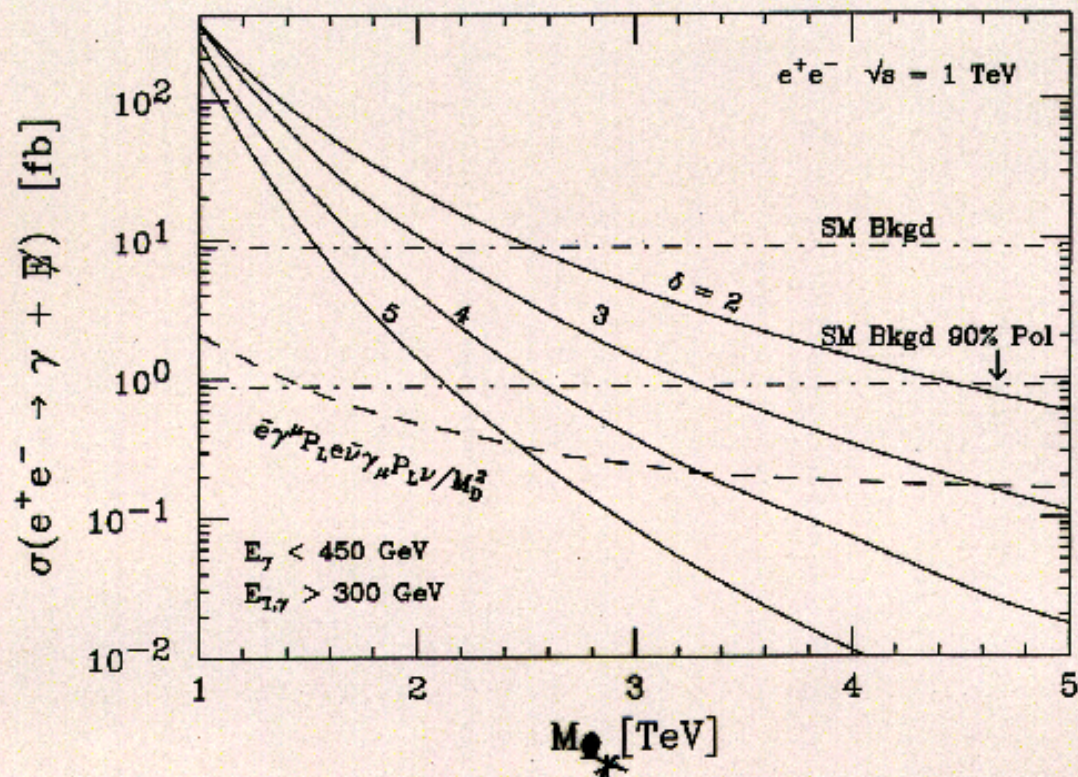
Wilson







# Graviton Emission : $e^+e^- \rightarrow \gamma + G^{(n)}$



Giudice  
Ratazzi  
We

Figure 2: Total  $e^+e^- \rightarrow \gamma + \text{nothing}$  cross-section at a 1 TeV centre-of-mass energy  $e^+e^-$  collider. The signal from graviton production is presented as solid lines for various numbers of extra dimension ( $\delta = 2, 3, 4, 5$ ). The Standard Model background for unpolarized beams is given by the upper dash-dotted line, and the background with 90% polarization is given by the lower dash-dotted line. The signal and background are computed with the requirement  $E_\gamma < 450$  GeV in order to eliminate the  $\gamma Z \rightarrow \gamma \nu \nu$  contribution to the background. The dashed line is the Standard Model background subtracted signal from a representative dimension-6 operator.

	2	4	6
$\rho_{-1+} = 0$	5.9	3.5	2.5
$\rho_{-} = 80\%$	8.3	4.4	2.9
$\rho_{+} = 80\%$	10.4	5.1	3.3
$\rho_{+} = 60\%$			

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Note: signal  $\uparrow$  w/  $\sqrt{s} \uparrow$   
background  $\downarrow$  w/  $\sqrt{s} \uparrow$





# LEP2 Lower 95% CL $M_S$ (Hewett) Limits (TeV)

Experiment	$e^+e^- \rightarrow \gamma G$						$e^+e^- \rightarrow ZG$						Color coding
	$n=2$	$n=3$	$n=4$	$n=5$	$n=6$		$n=2$	$n=3$	$n=4$	$n=5$	$n=6$		
ALEPH	1.28	0.97	0.78	0.66	0.57		0.35	0.22	0.17	0.14	0.12		$\leq 184$ GeV
DELPHI	1.38	1.02	0.84	0.68	0.58								$\leq 189$ GeV
L3	1.02	0.81	0.67	0.58	0.51		0.60	0.38	0.29	0.24	0.21		$> 200$ GeV
OPAL	1.09	0.86	0.71	0.61	0.53								$\lambda=-1$ GL $\lambda=+1$

## Virtual Graviton Exchange

Experiment	$e^+e^-$	$\mu^+\mu^-$	$\tau^+\tau^-$	$q\bar{q}$	$f\bar{f}$	$\gamma\gamma$	$W^+W^-$	$ZZ$	Combined
ALEPH	1.04 0.81	0.65 0.67	0.60 0.62	0.53/0.57 0.49/0.49 (bb)	1.05 0.84	0.81 0.82			0.75/1.00 (<189)
DELPHI		0.59 0.73	0.56 0.65		0.60 0.76	0.70 0.77			0.60/0.76 (ff) (<202)
L3	0.98 1.06	0.56 0.69	0.58 0.54	0.49	0.84 1.00	0.99 0.84	0.68 0.79	1.2 ? 1.2 ?	1.3/1.2 (<202) ?
OPAL	1.15 1.00		0.62 0.66		0.62 0.66	0.89 0.83		0.63 0.74	1.17/1.03 (<209)



# Supersymmetric Bulk

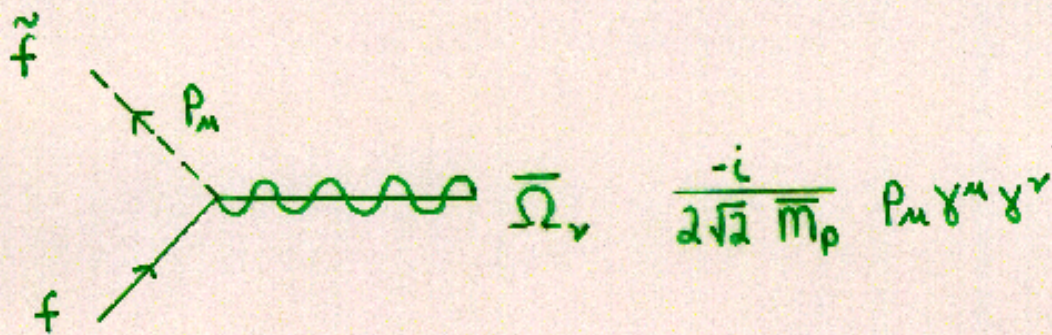
JLH, Sadri

Motivation : Embed ADD in string theory  
Stabilize size of dimensions

$N=2$  SUSY in bulk breaks to  $N=1$  SUSY on brane

Full supermultiplet in bulk

Compactify  $\Rightarrow$  KK towers of gravitons AND gravitinos



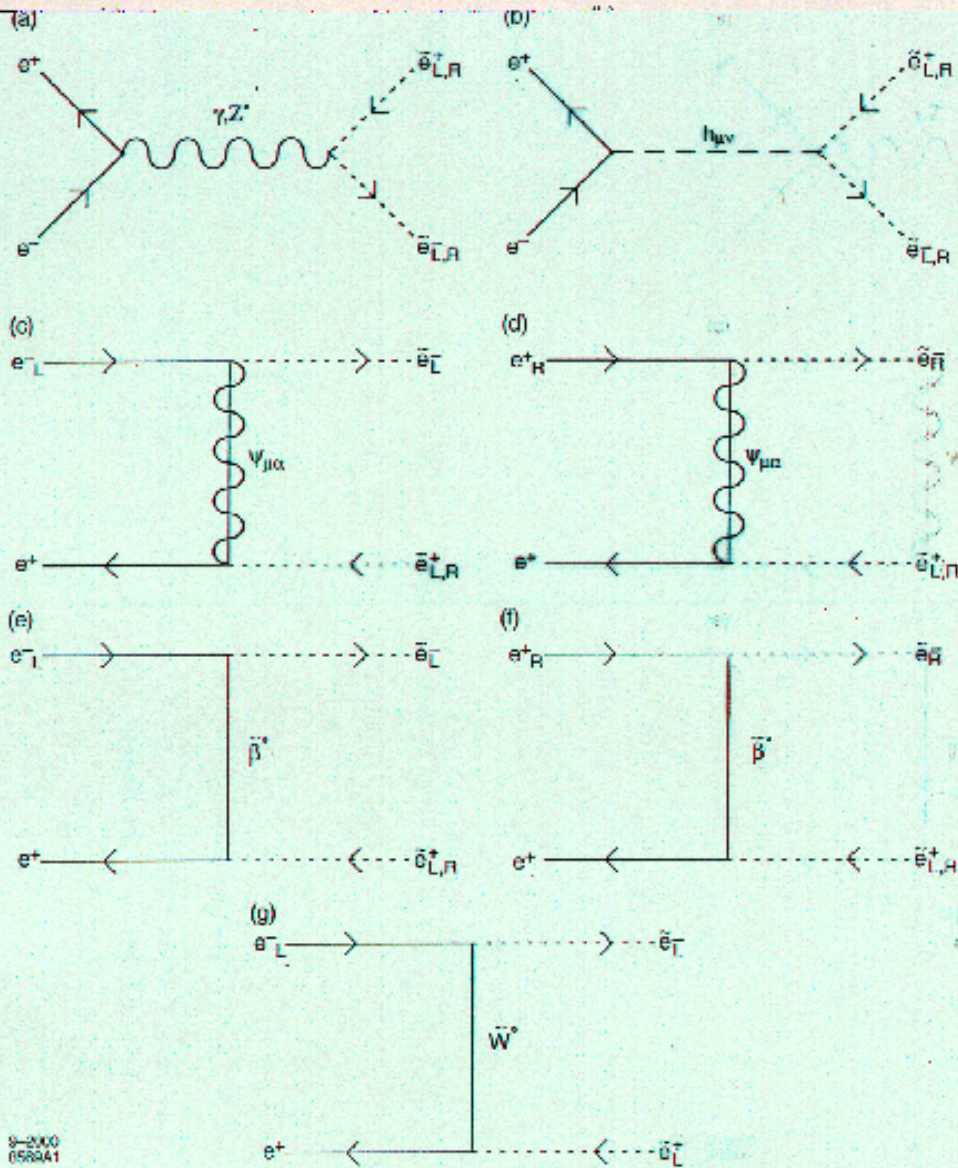
KK Gravitino Exchange : t-channel

$$\sum_{\tilde{n}} p^{\tilde{n}, \mu\nu} \sim -i \int_{m_0^2}^{\Lambda_c^2} dm_{\tilde{n}}^2 \rho(m_{\tilde{n}}^2) \frac{|m_{\tilde{n}}|^{\alpha-2}}{t - m_{\tilde{n}}^2}$$

$$\alpha = 0, 1, 2, 3$$



# Contributions to $e^+e^- \rightarrow \tilde{e}_{L,R}^+ \tilde{e}_{L,R}^-$



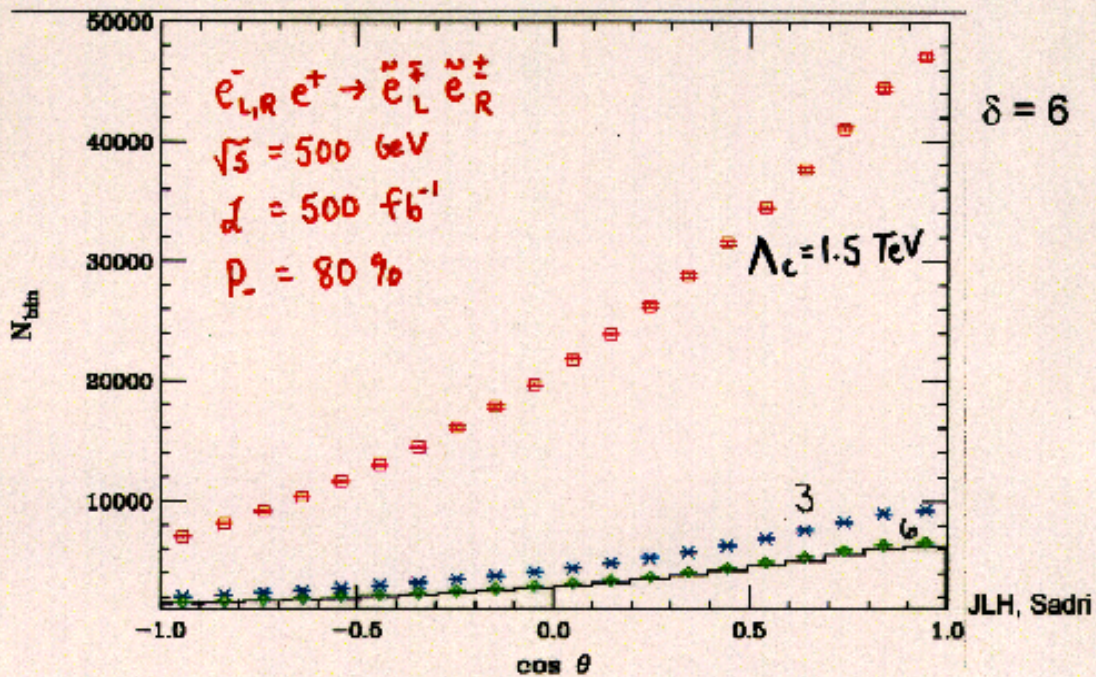
9-2000  
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Figure 2: Various processes contributing to scalar pair production.



# Supersymmetric Large Extra Dimensions

- Exchange KK towers of Gravitons and Gravitinos!
- Large t-channel contribution to  $e^+e^- \rightarrow \tilde{e}^+\tilde{e}^-$  from Gravitino exchange



Sensitivity in excess of  $M_* \sim 10 \text{ TeV}$  @  $\sqrt{s} = 500 \text{ GeV}$

Unique sensitive probe to a supersymmetric bulk!

LHC: Small signal in squark pair production  
 gluino production needs to be studied



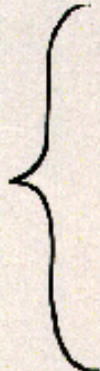
## TeV<sup>-1</sup> - size Extra Dimensions

Antoniadis

The SM goes into the bulk!

Many model building choices

- Are  $SU(3)_c$ ,  $SU(2)_L$ , +  $U(1)_Y$  all in the bulk? Same # of dims?
- Higgs in bulk or on brane?

- Fermions 
  - fixed points
  - in bulk
  - localized at various points on a thick brane

For simplicity, examine 1 extra dimension

compactified on  $S^1/\mathbb{Z}_2$  (circle of radius  $R$ , folded in half)



# TeV<sup>-1</sup> Extra Dimensions (Fat Branes)

- Arise naturally from string theory
- Mechanism to suppress proton decay

Antoniadis

Arkani-Hamed,  
Schmalt

Gauge bosons free to propagate in  $R_c \sim \text{TeV}$

⇒ Degenerate KK towers for  $\gamma/Z/W/g$

Discovery Reach for KK  $\gamma/Z$  (TeV): Ala Z' search

Direct	Run II	2 fb <sup>-1</sup>	1.1
	LHC	100 fb <sup>-1</sup>	6.3
Indirect	LC	$\sqrt{s} = 0.5 \text{ TeV}$	500 fb <sup>-1</sup> 13.0
	LC	$\sqrt{s} = 1.0 \text{ TeV}$	500 fb <sup>-1</sup> 23.0
	LC	$\sqrt{s} = 1.5 \text{ TeV}$	500 fb <sup>-1</sup> 31.0

Snowmass/LesHouches '01:

LHC Indirect Reach

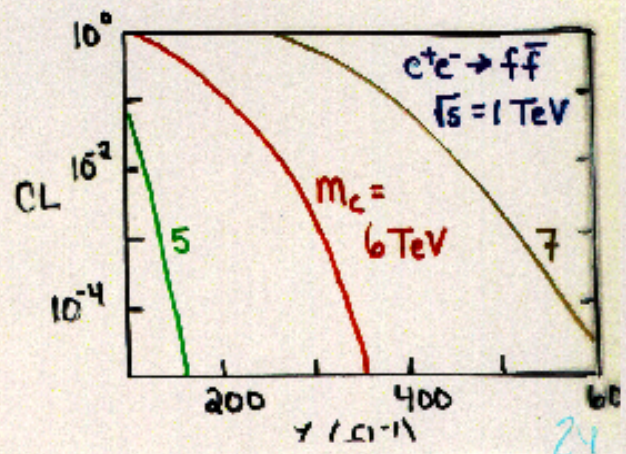
ATLAS Study:  $m_c \sim 8 \text{ TeV}$

2 Pheno studies JLH, Rizzo  
Cheung Landi

Can KK  $\gamma/Z$  be distinguished from GUT Z'?

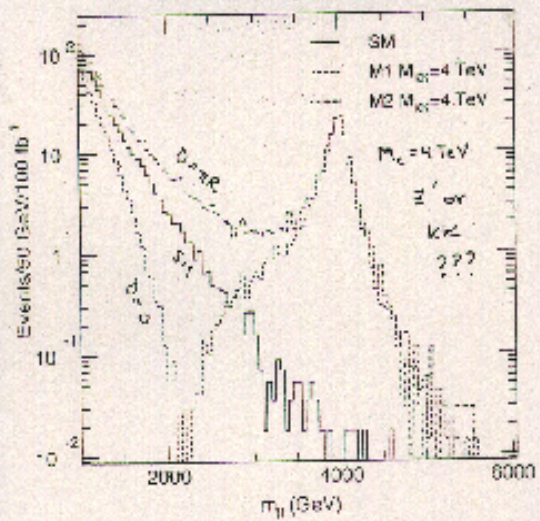
Rizzo

LHC:	$m_c = 4 \text{ TeV}$	5 TeV
D=0	CL < 10 <sup>-10</sup>	CL ~ 0.3%
D=πR	CL ~ 70%	CL ~ 100%





ATLAS  $e^+e^-$  preliminary

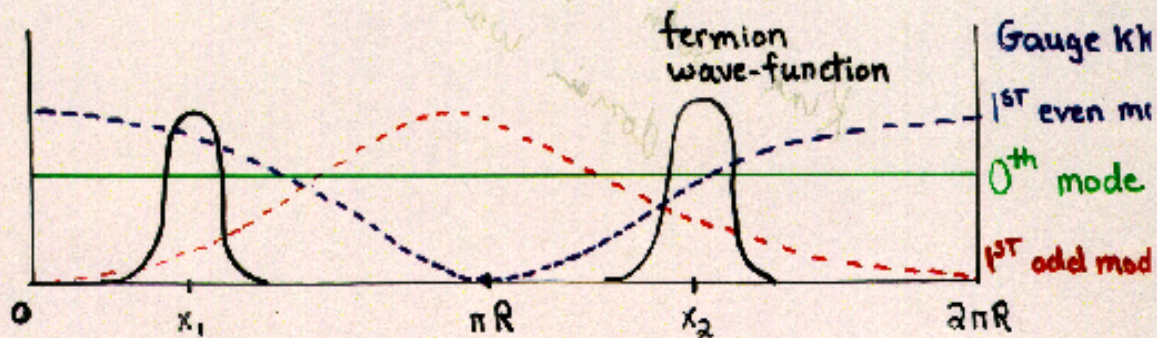




# Separated Fermions

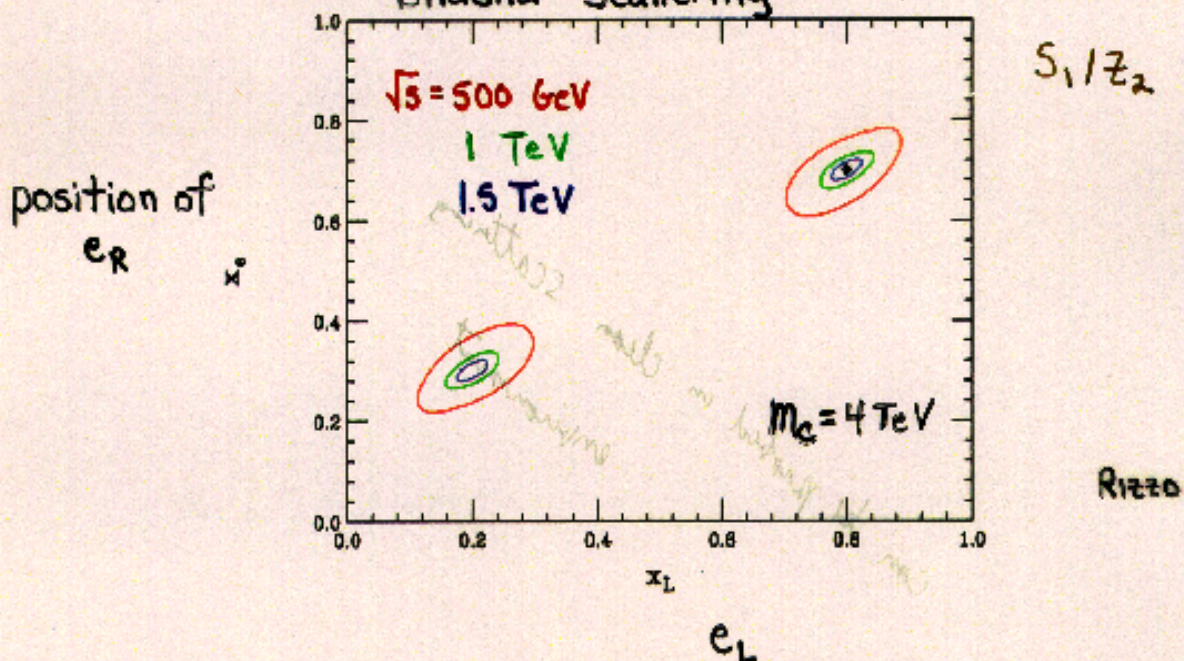
Arkani-Hamed, Schmaltz

Fermions can be localized at different points in a thick brane



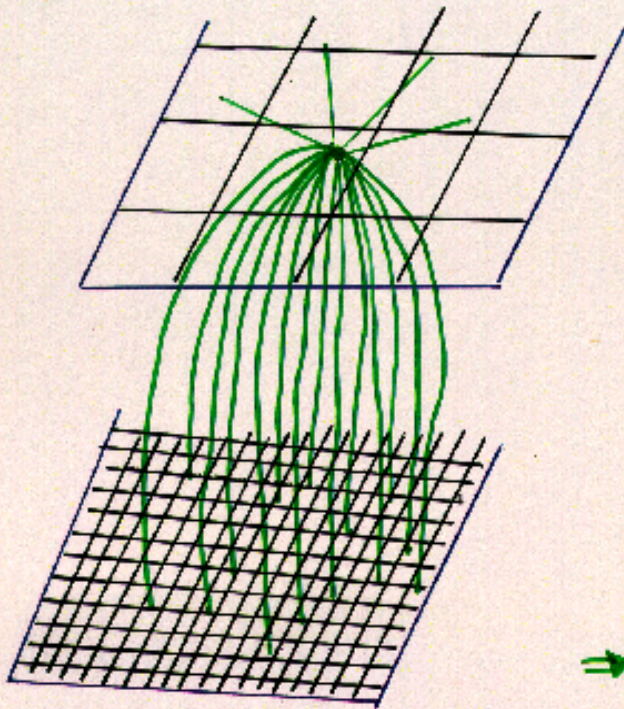
Gauge KK couplings probe relative fermion locations!

## Bhabha Scattering





## Non-Factorizable Curved Geometry - 'Warped' Space



Area of each grid  
is equal

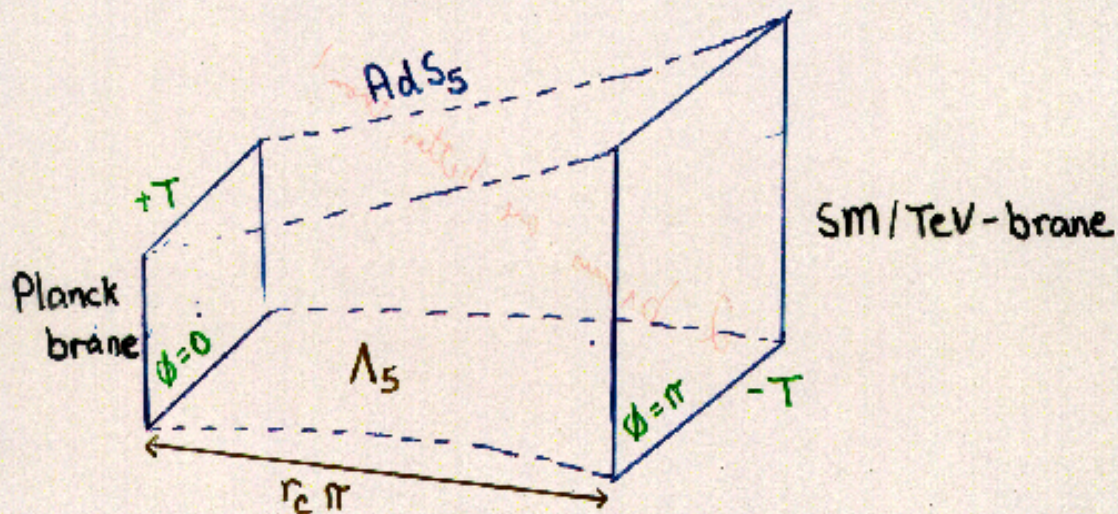
Field lines spread out  
faster with more volume

⇒ Drop to bottom brane

Gravity appears weak on top brane!



## Localized Gravity ala Randall-Sundrum



Bulk = Slice of AdS<sub>5</sub>

Two 3-branes at S<sub>1</sub>/Z<sub>2</sub> orbifold fixed points

5-D, non-factorizable geometry

Solutions to Einstein's Eqn [w/ 4-D Poincaré']

$$ds^2 = e^{-2kr_c|\phi|} \eta_{\mu\nu} dx^\mu dx^\nu + r_c^2 d\phi^2$$

Warp factor

$$0 \leq |\phi| \leq \pi$$

r<sub>c</sub> = Compactification radius

where  $\Lambda_5 = -24 m_5^3 k^2$



## 4-D Effective Action:

$$\bar{m}_{\text{pl}}^2 = \frac{m_s^3}{k} (1 - e^{-2kr_c\pi}) \Rightarrow k \sim m_s \sim \bar{m}_{\text{pl}}$$

no additional hierarchie

## Physical scales:

$$\Lambda_\emptyset = e^{-kr_c|\emptyset|} \bar{m}_{\text{pl}}$$

For TeV-brane at  $\emptyset = \pi$ 

$$\Lambda_\pi = e^{-kr_c\pi} \bar{m}_{\text{pl}} \sim \text{TeV} \quad \text{if } kr_c \sim 11-12$$

 $\Rightarrow$  hierarchy generated by an exponential!

stabilized via Goldberger + Wise

Phenomenology governed by

$$k/\bar{m}_{\text{pl}} \text{ and } \Lambda_\pi$$

 $\Rightarrow$  only 2 free parameters!

## 5-D curvature:

$$|R_5| = 20k^2 < m_s^2$$

(neglecting higher-order curvature terms)

suggests  $k/\bar{m}_{\text{pl}} < 0.1$ 

$$\Rightarrow r_c \sim \frac{100}{\bar{m}_{\text{pl}}}$$



## 4-D Effective Theory:

Linear expansion of flat metric

$$G_{\alpha\beta} = e^{-2kr_c|\vartheta|} (\eta_{\alpha\beta} + \kappa h_{\alpha\beta}) \quad \kappa = 2m_5^{-3/2}$$

Expand into KK tower

$$h_{\alpha\beta}(x, \vartheta) = \sum_{n=0}^{\infty} h_{\alpha\beta}^{(n)}(x) \frac{\chi_n^{(n)}(\vartheta)}{\sqrt{r_c}}$$

Employ gauge  $\eta^{\alpha\beta} \partial_\alpha h_{\beta\gamma}^{(n)} = 0$  +  $\eta^{\alpha\beta} h_{\alpha\beta}^{(n)} = 0$

Demand  $\int_{-\pi}^{\pi} d\vartheta e^{-2kr_c|\vartheta|} \chi_h^{(m)} \chi_h^{(n)} = \delta_{mn}$

$$\chi_h^{(n)}(\vartheta) = \frac{e^{2kr_c|\vartheta|}}{N_n} \left[ J_2\left(\frac{m_n}{k} e^{kr_c|\vartheta|}\right) + \alpha_n Y_2\left(\frac{m_n}{k} e^{kr_c|\vartheta|}\right) \right]$$

For TeV-brane:

$$m_n = x_n k e^{-kr_c\pi} \quad \text{with } J_1(x_n) = 0$$

$$= x_n \frac{k}{\bar{m}_{pl}} \Lambda_{\pi}$$

$\Rightarrow$  KK excitations are not evenly spaced!



## Interactions

$$\mathcal{L} \sim \frac{1}{m_5^{3/2}} T^{\alpha\beta}(x) h_{\alpha\beta}(x, \vartheta = \pi)$$

$$\equiv \frac{1}{\bar{m}_{pl}} T^{\alpha\beta}(x) h_{\alpha\beta}^{(0)}(x) - \frac{1}{\Lambda_\pi} T^{\alpha\beta}(x) \sum_{n=1}^{\infty} h_{\alpha\beta}^{(n)}(x)$$

Zero-mode decouples

TeV-suppressed  
 $\rightarrow$  can be directly produced!

## Phenomenology

Davoudiasl, JLH, Rizzo

PRL 100  
 PLB 100  
 PRD 101  
 PLB 100  
 hep-ph/0010066

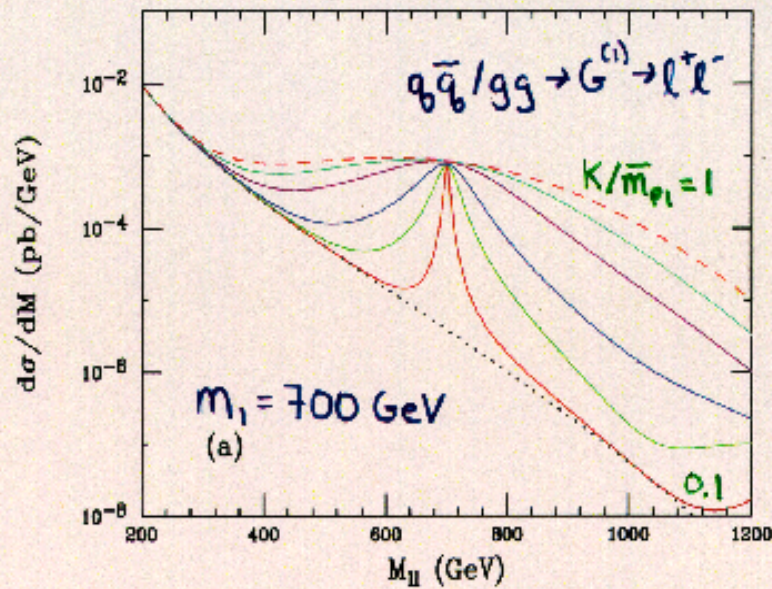
- Graviton resonance production
- Graviton contributions to EW oblique parameter
- 'Light, skinny' Gravitons [ $k/\bar{m}_{pl} \lesssim 0.01$ ]  
 Graviton emission
- Below resonance exchange  
 "Contact interaction" limits



# KK Graviton Drell-Yan Spectrum

Davoudiasl, J  
Rizzo PRL'

Tevatron



LHC

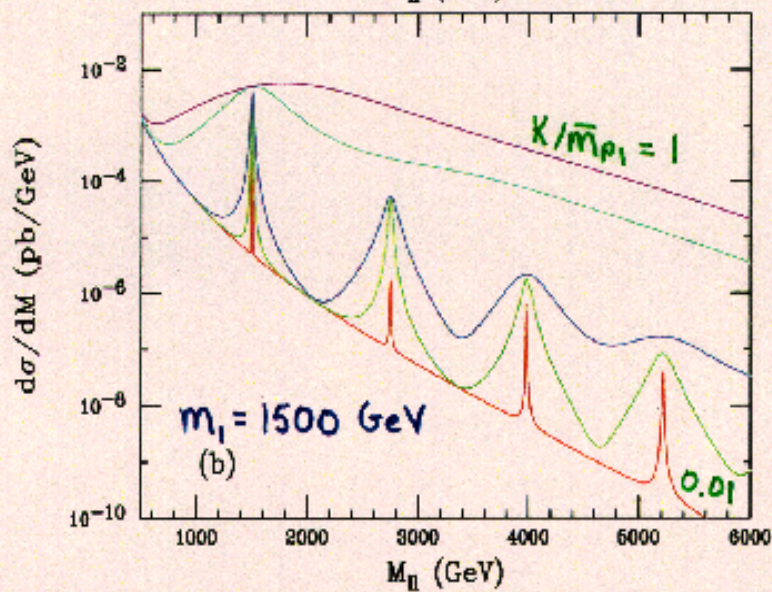


Figure 17: Drell-Yan production of a (a) 700 GeV KK graviton at the Tevatron with  $k/\bar{M}_{Pl} = 1, 0.7, 0.5, 0.3, 0.2,$  and  $0.1,$  respectively; (b) 1500 GeV KK graviton and its subsequent tower states at the LHC. From top to bottom, the curves are for  $k/\bar{M}_{Pl} = 1, 0.5, 0.1, 0.05,$  and  $0.01,$  respectively.



# Constraints from Drell-Yan + di-jet production

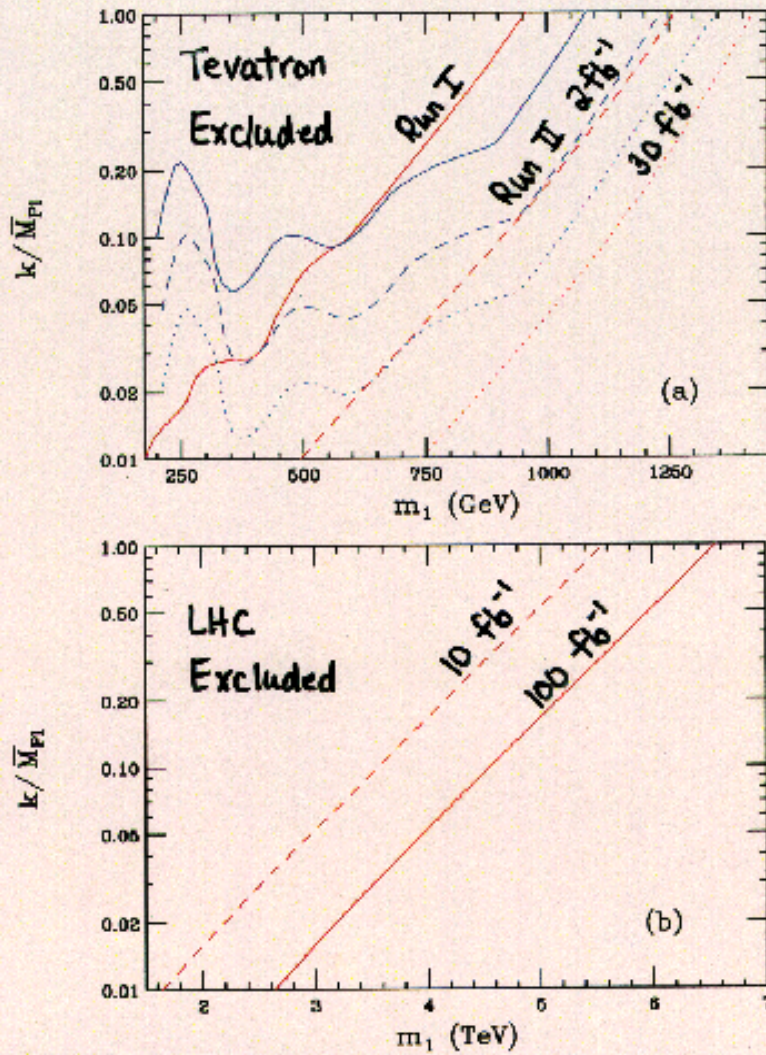
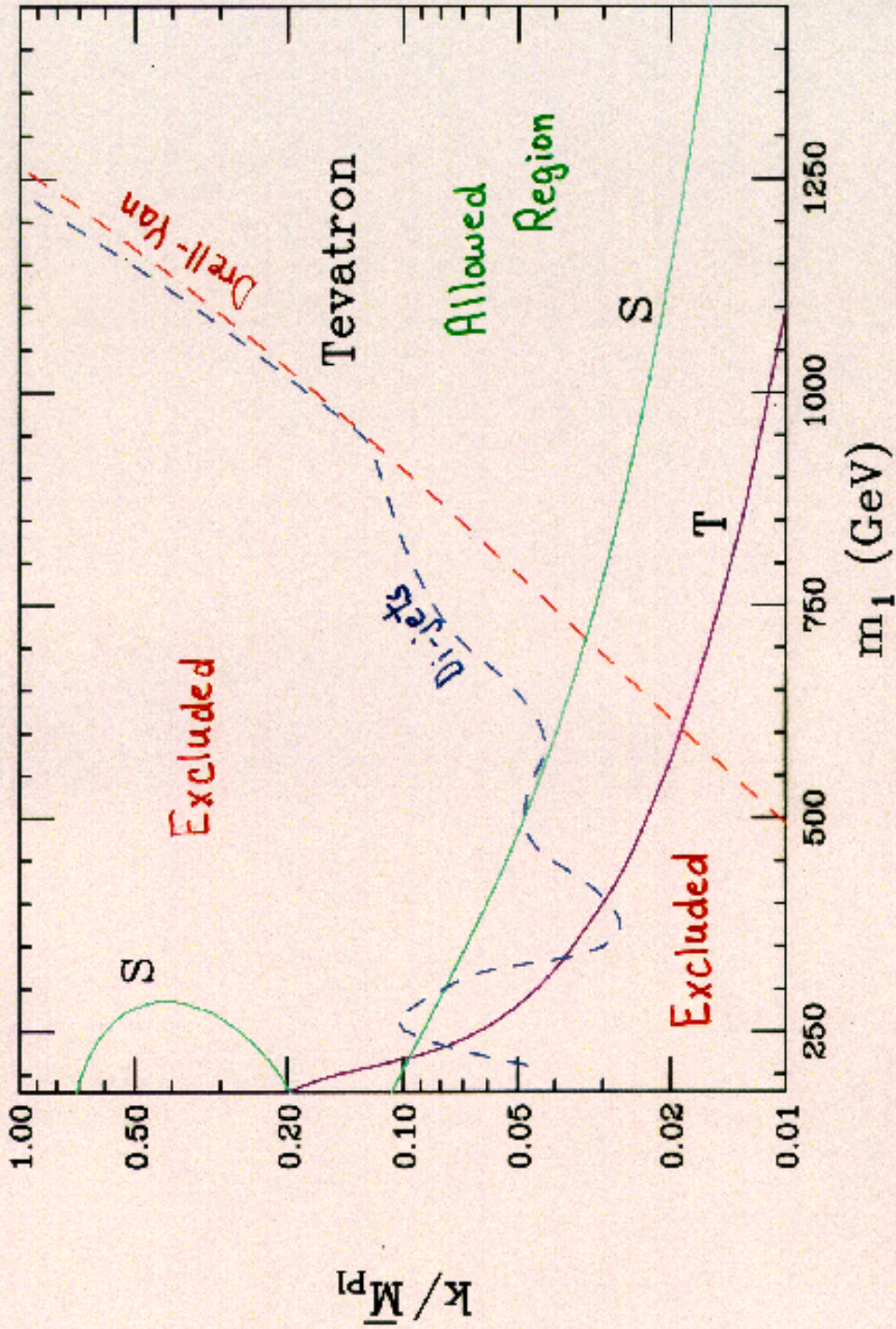


Figure 1: Exclusion regions for resonance production of the first KK graviton excitation in (a) the Drell-Yan (corresponding to the diagonal lines) and dijet (represented by the bumpy curves) channels at the Tevatron and (b) Drell-Yan production at the LHC. (a) The solid curves represent the results for Run I, while the dashed, dotted curves correspond to Run II with 2, 30  $\text{fb}^{-1}$  of integrated luminosity, respectively. (b) The dashed, solid curves correspond to 10, 100  $\text{fb}^{-1}$ . The excluded region lies above and to the left of the curves.

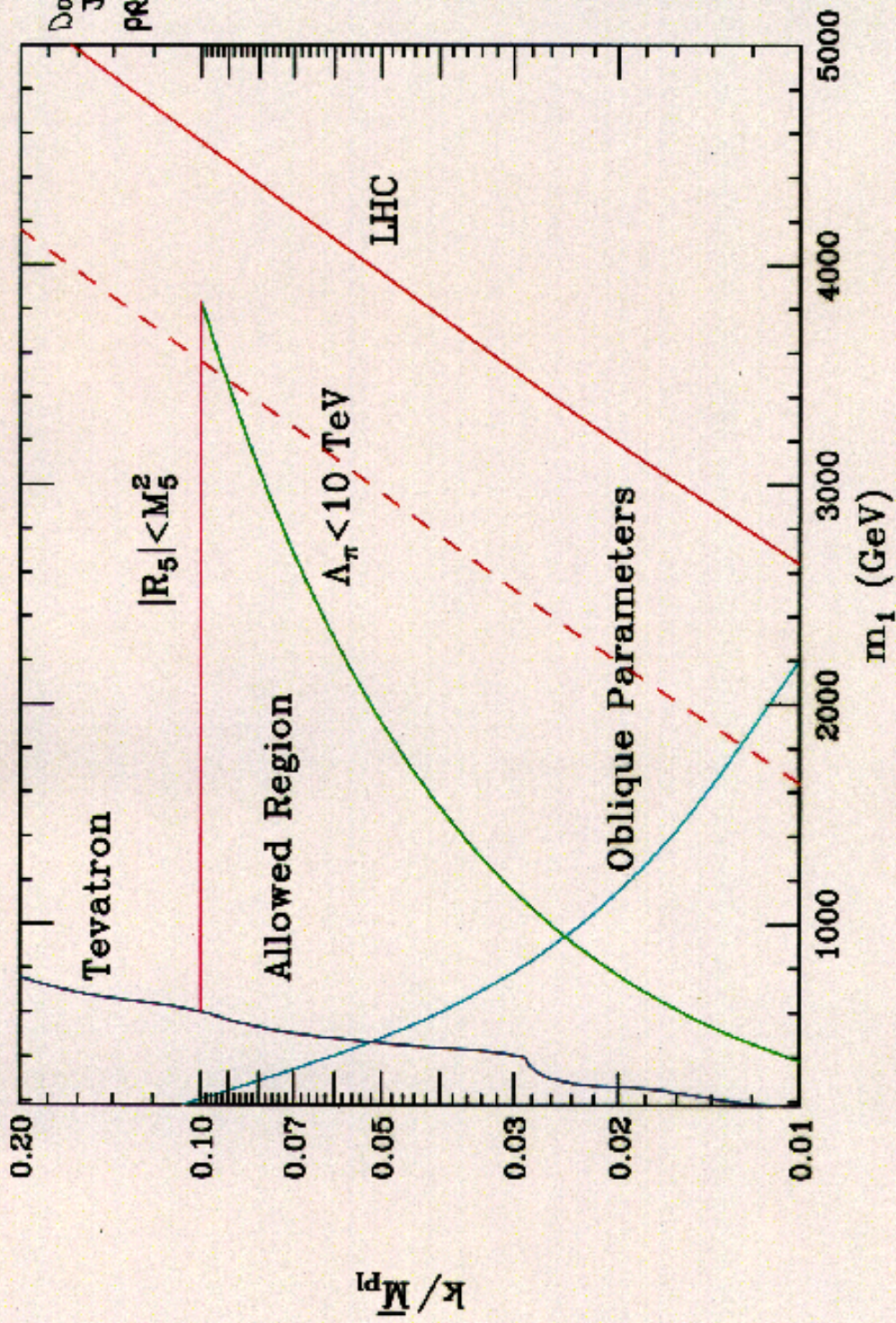


# Present Constraints on RS Model

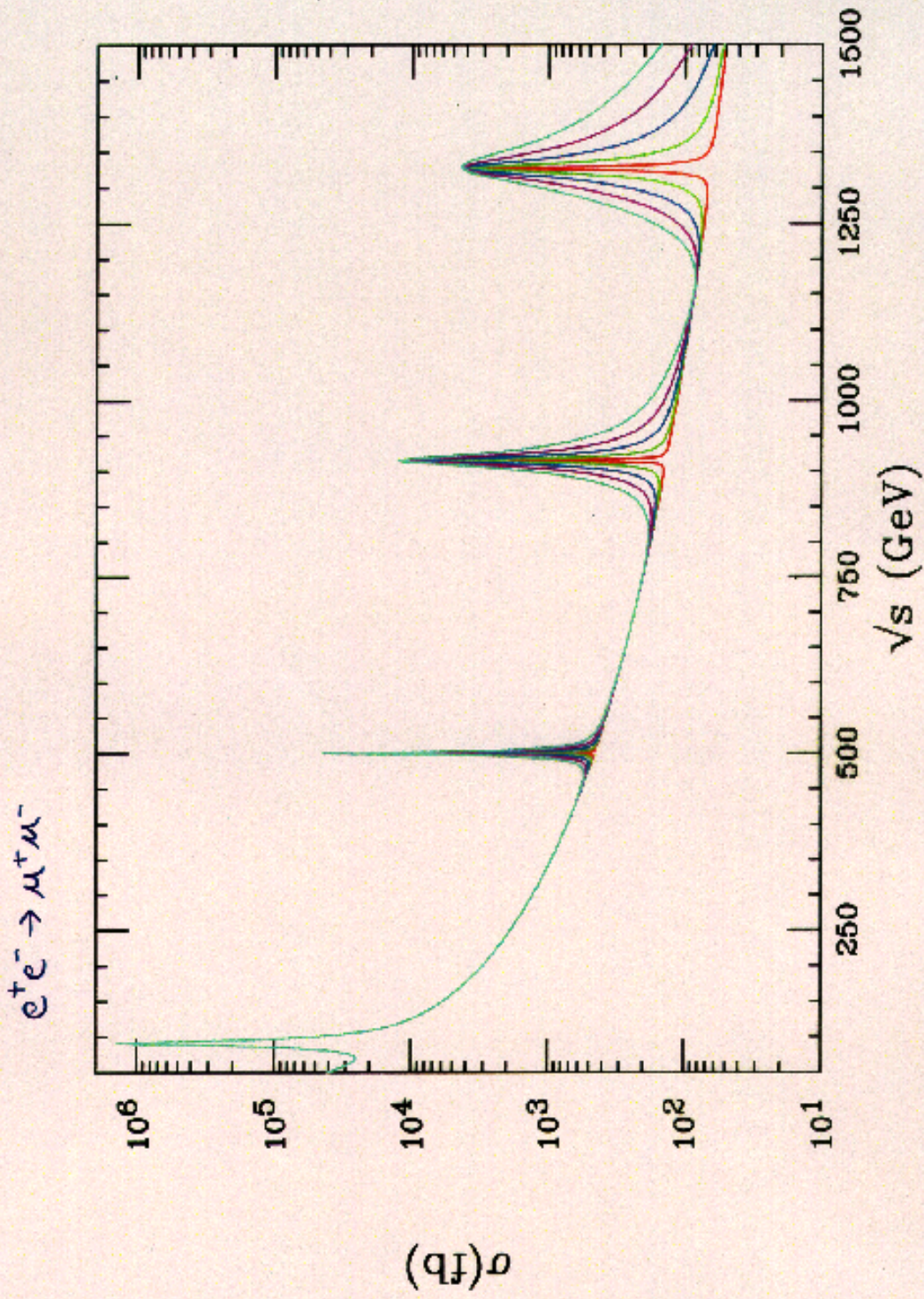




# Summary of Theory + Exp't Constraints

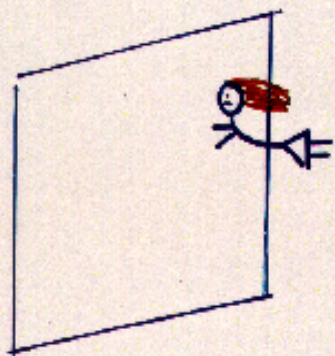








## Peeling the SM off the Wall



Davoudiasl, JLH, Rizzo  
PLB 100  
PRD 01

- Gauge Fields in Bulk Alone

EW Data  $m_A > 25 \text{ TeV!}$  {violates curvature constraint}

- Fix: Add Fermions in the bulk

⇒ Introduces 3<sup>rd</sup> parameter

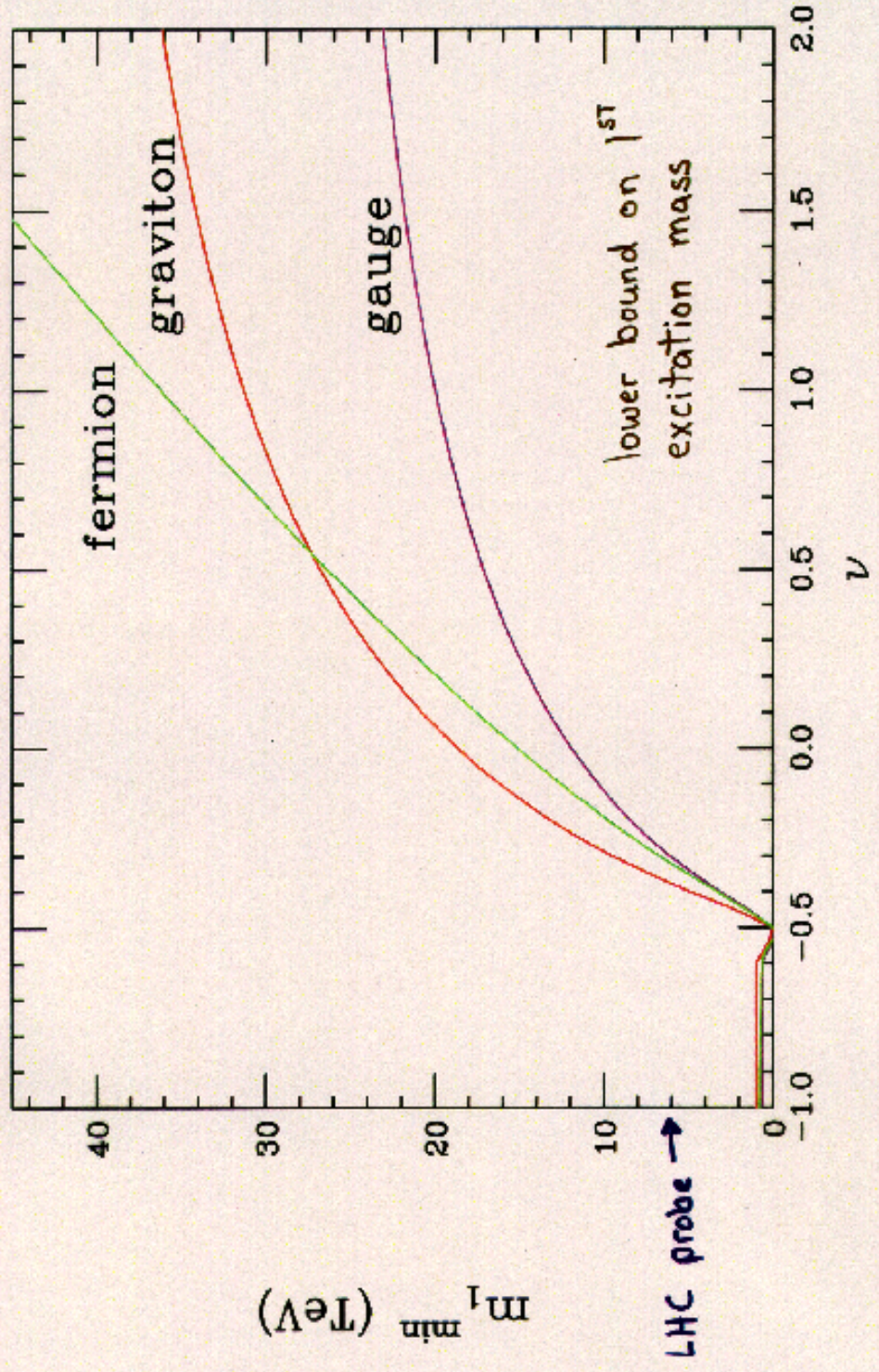
$$m_f^{\text{bulk}} = \gamma K, \quad \gamma \sim \mathcal{O}(1)$$

Zero-mode KK fermions couple more weakly than Wall fields

⇒ Serious reduction in collider sensitivity!

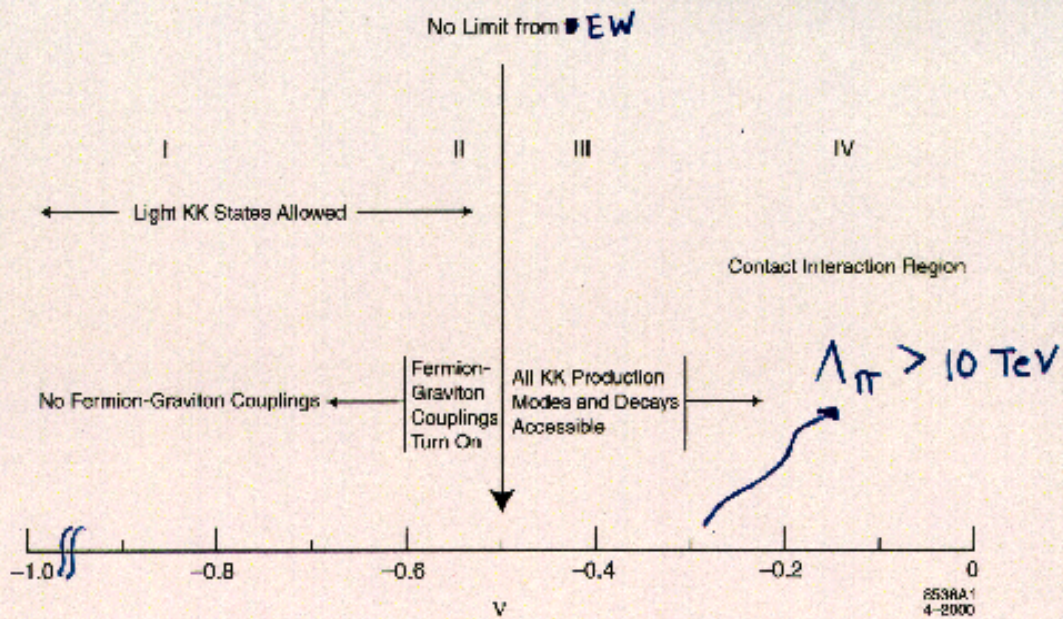


# Precision EW Data Constraints





# Suggests Distinct Phenomenological Regions



Constraint from  
No fine-tuning of  
Yukawa's



## Summary:

### Large Extra Dimensions:

Gravity only in bulk.

Finely spaced KK tower. [ $\Sigma$  states yields observability]

Consequences for Cavendish-type expt.

Large Astro/Cosmo constraints for  $n=2$ .

Collider Probes: Graviton Emission } Present:  $m_* \gtrsim \text{TeV}$   
Exchange } Future:  $m_* \gtrsim 6-7$

### TeV<sup>-1</sup> Extra Dimensions:

Large Model Building Choices!

SM in bulk

Fermions: Fixed points -  $m_c \gtrsim 4 \text{ TeV}$

Localized - Natural suppression of p dk

Bulk - Evades Limits!  $m_c \gtrsim 400 \text{ GeV}$

### Localized Gravity:

No new hierarchies

Direct Graviton Resonance production

LHC probes parameter space: SM on wall

Drastic change in phenomenology: SM off wall



# The Physics of Extra Dimensions

Spring '98: Extra Dimensional Revolution!

Plethora of Models now exist

- Began classification of models and resulting phenomenology  
Snowmass 2001 Davoudiasl, JLT, Riee

Use Extra Dimensional Geometry to:

- Solve hierarchy problem
- Generate EWSB
- Generate SUSY Breaking
- Generate flavor breaking
- Generate  $\nu$  masses (see-saw from the bulk)
- Suppress proton decay
- ✓ • Grand Unification possible with power law running Dienes Randal
- Progress on cosmological constant problem

Rash of New Ideas!

Implications for Tabletop/Flavor/Collider/Astro-Cosmological/Precision EW Physics!

